



Preliminary results of MARE in Milan

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Outline



- Physical motivation and the calorimetric approach
- MARE-1 in Milano
- Re option
- Ho option

Physical motivation



direct neutrino mass measurement



General experimental requirements:

- High statistics at the beta spectrum end-point
- high energy resolution ΔE

Calorimeters measure the **entire spectrum** at once:

- low $\mathsf{E}_{_0}\,\beta$ decaying isotopes for more statistics near the end-point
- best choice ¹⁸⁷Re:

$$- E_0 = 2.5 \text{ keV}, \tau^{1/2} = 4 \times 10^{10} \text{ y}$$

• other option ¹⁶³Ho EC:

•
$$- E_0 \approx 2.6 \text{ keV}, \tau^{1/2} \approx 4600 \text{ y}$$

15 eV \rightarrow ¹⁸⁷Re (E₀=2.47keV) & calorimeters

Calorimeters: source ⊆ detector



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MARE



MARE-1: a few eV direct neutrino mass measurement

Two isotopes under investigations:

- > 187Re high statistics measurement
 - > asses systematics
 - test large arrays
 - Iower limit to few eV
- 163Ho high statistics measurement R&D for 163Ho production
 - measure Q_{EC}
 - study spectrum shape
 - asses systematics

MARE-1 in Milan: Milano/FBK/Wisconsin/NASA

- $m_{v^e} < 2 \text{ eV/c}^2$
- 10¹⁰ events 300 sensors
- 8 arrays of Si:P thermistors with AgReO₄ absorbers
- energy resolution 25 eV @ 2.6 keV

This experiment is needed:

- because it's the only possible one with present technology
- To investigate systematics in thermal calorimeters

very important to cross-check spectrometer results

MARE-1 detectors/Re

• 187 Re β -decay

- ${}^{187}\text{Re} \rightarrow {}^{187}\text{Os} + \text{e-} + \nu_{\text{e}}$ $E_0 = 2.47 \text{ keV}$
- i. a. 63% and τ =42.3 Gy
- Single crystal of silver perrhenate (AgReO₄)
 - mass ~ 500 μ g per pixel (A_{β}~ 0.3 decay/sec)
 - regular shape ($600 \times 600 \times 250 \ \mu m^3$)
 - low heat capacity due to Debye law
- 6x6 array of Si:P semiconductors (NASA-GSFC)
 - pixel: 300x300x1.5 μm³
 - high energy resolution
 - developed for X-ray spectroscopy with HgTe absorber (ASTRO-E2)







Cryogenic set-up/Re





MARE 1 /Re



All the problems concerning the cryogenic set-up have been solved.



Thanks to the improvements added to the cryogenic set-up the detector target performances have been achieved.



- First spectra acquired
- Completed assembly of the first array





Measured 7 pixels: $\Delta E_{ave} \sim 30eV @ 1,5keV$

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First array /Re





Thermal coupling

- Araldit or ST1266: thermistor/spacer
- ST2850: spacer/AgReO₄



- > 6 silicon spacers are attached with Araldite Normal
- > 10 with Araldite Rapid
- > 15 with **ST1266**

First array/Re



A run aimed to test the performance of this setup is ongoing, after which the absorbers will be glued also on the second array. With two arrays, a sensitivity of 4.5 eV at 90% C.L. is expected in three years running time.





Energy [eV]

MARE 1 with Rhenium: sensitivity



Estimation of the sensitivity on neutrino mass over the years by increasing the detectors number from year to year.



MARE-Ho: ¹⁶³Ho EC measurement



electron capture from shell > M1 163 Ho + e⁻ \Rightarrow 163 Dy* + v_o A. De Rujula and M. Lusignoli, Phys. Lett. B 118 (1982) 429 10¹² 5×104 - Q_{EC} = 2.20 keV N1 $m_v = 0 \, \text{eV}$ 1011 Q_{EC} = 2.55 keV M1 4×10⁴ N2 $m_v = 5 \, \text{eV}$ 70.12 € / 0.12 ¢ counts / 0.3 eV M2 $m_{\rm V} = 10 \, {\rm eV}$ / stupe $Q_{\rm ec} = 2.555 \, \rm keV$ pile-up 1×104 10' $N_{ev} = 10^{14}; f_{pp} = 10^{-6}; \Delta E_{FWHM} = 2eV$ 10⁶ 0.5 2.5 2.55 2 2.53 2.54 2.5 energy [keV] energy [keV]

Calorimetric measurement of non-radiative Dy atomic de-excitations
Breit Wigner M,N,O lines have an end-point at the Q value
rate at end-point may be as high as for ¹⁸⁷Re but depends on Q_{FC}

✓ Q_{EC} ? Measured: Q_{EC} = 2.3÷2.8 keV. Recommended: Q_{FC} = 2.555 keV

 $\checkmark \tau$ $_{_{1/2}} \approx$ 4570 years: few active nuclei are needed

✓ can be implanted in any suitable microcalorimeter absorber
✓ ¹⁶³Ho production by neutron irradiation of ¹⁶²Er enriched Er

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MARE-Ho

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Goals:

Production of metal absorber with Ho-163 metal homogeneously embedded in.

Issues:

Production of Ho-163: neutron activation of enriched Er-162 oxide

Er-162(n, γ)Er-163 →Ho-163 (3 processes in the last 2 years)

- Purification: under study with different methods (PSI, LANL)
- ✓ Reduction to metal form: recently demonstrated (Genova) [see G. Pizzigoni Poster D208].
- Embedding in the absorber: first test done



MARE /Ho option



A run aimed to test the new absorbers made of Sn/Ho/Sn is ongoing @ Milano-Bicocca. The absorbers are glued on Si thermistors.

The cryogenic set-up and electronics are the one used in the Mibeta experiment as well as the dilution unit.



Detector holders mounted under the MC

Cold electronic JFET 10 detectors: 9 with Ho and 1 with Sn



Zoom of one single detector



Al bonding wires

Sn/Ho/Sn absorber Glued on Si thermistor

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Average pulses and first spectra



Average pulse of detector 3



Conclusion



First array of MARE-1 has been assembled

 \rightarrow 31 thermistors are equipped with AgReO4 absorbers

The goal performances of the detectors have been achieved \rightarrow first spectra were acquired obtaining a resolution of ~28eV @ 1,5keV

First samples of Ho(m) are produced and embedded into absorbers \rightarrow A first run to test Ho absorber produced by Genova is ongoing

In the meanwhile new detector technology under investigation \rightarrow Ho EC measurement with MKIDs (M. Faverzani poster B306)