(1000x) Stability boosting, and characterization, of high resolution spectrographs using an externally dispersed interferometer #47

David J. Erskine^a, J. Edelstein^b, E. Wishnow^c, M. Sirk^c, E. Linder^b, D. Fratanduono^a, (a) Lawrence Livermore Nat. Lab., (b) Lawrence Berkeley Nat. Lab., (c) UC Berkeley Space Sciences

Both On-sky and Off-sky uses benefit Doppler precision

Extreme Precision Radial Velocity 5 Mtg. March 27-30, 2023, Santa Barbara, CA

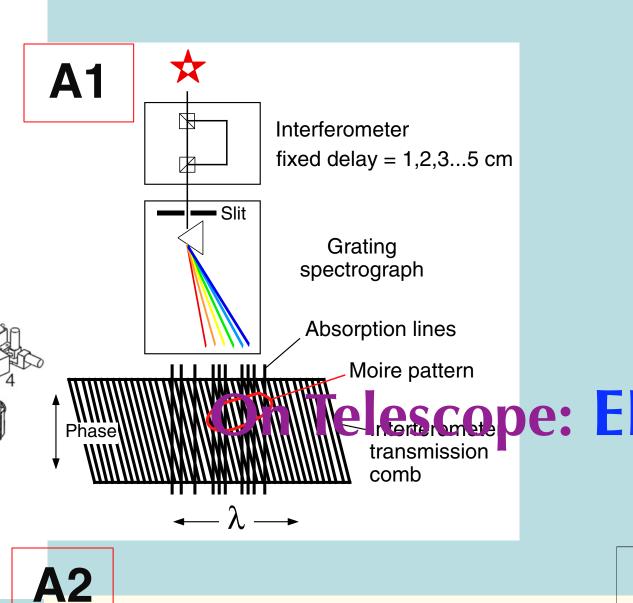


Plate holding optics: (steering mirrors, interferom., TV camera aser probe)

On-sky stability topic

EDI

Delav

KPF

Off-sky characterization topic

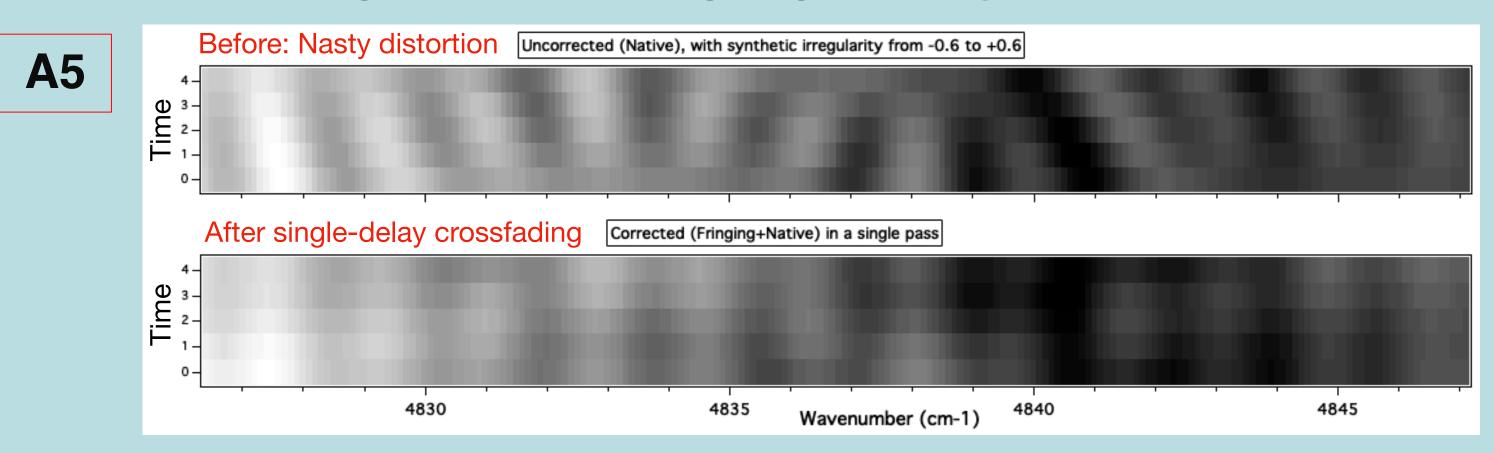
B1

Lamp

On-sky stability topic

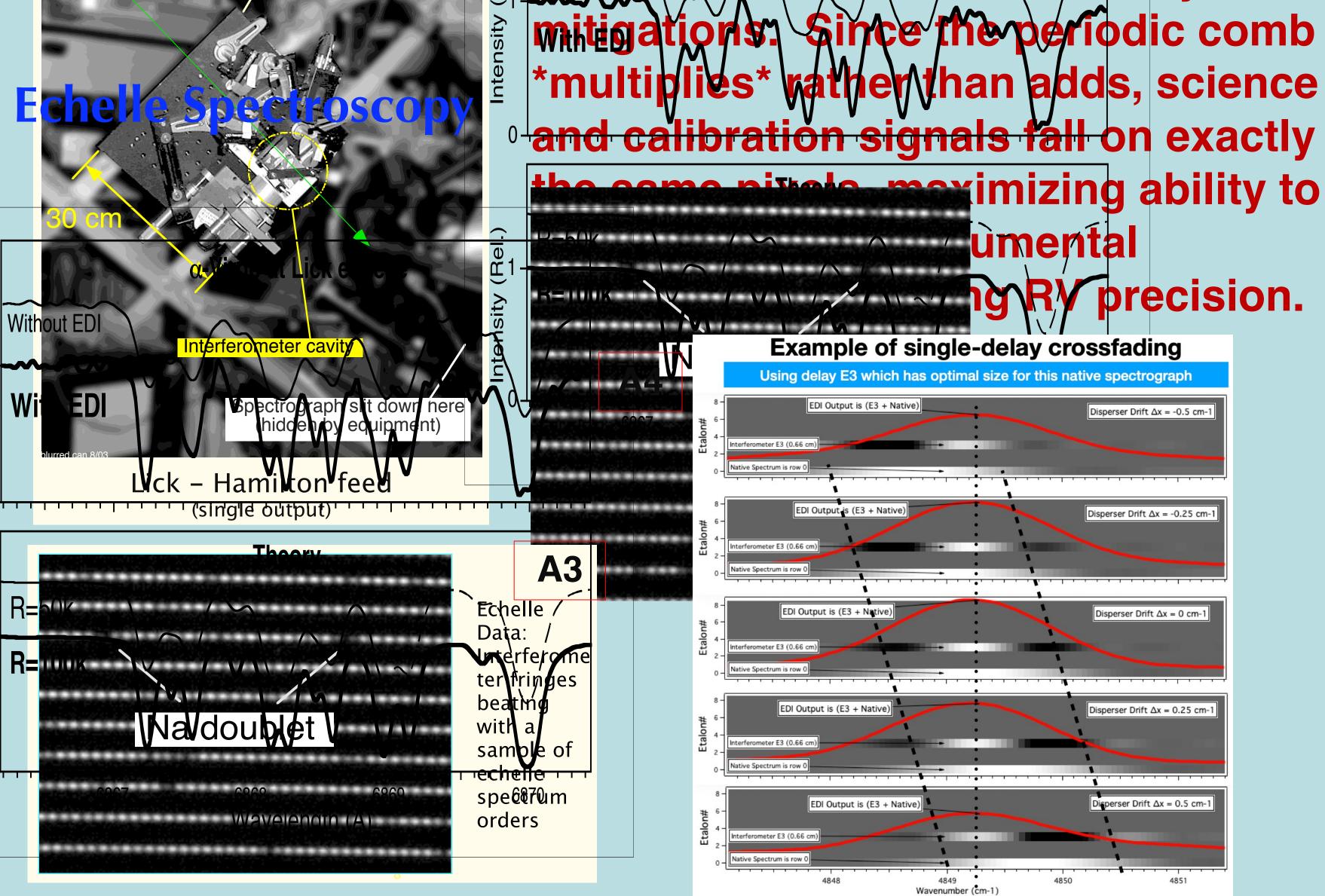
"Crossfading EDI" uses an in-series interferometer to stabilize against unknown and irregular on-sky spectrograph drifts by combining fringing and nonfringing components. These react oppositely in phase to a disperser or detector wavelength drift, and thus can be made to cancel in analysis.^{*a-Virgo at Lick echelle* **analysis**. This stability gain (~1000x)} multiplies conventional stability

Demo stabilizing the most egregious type of drift: bipolar

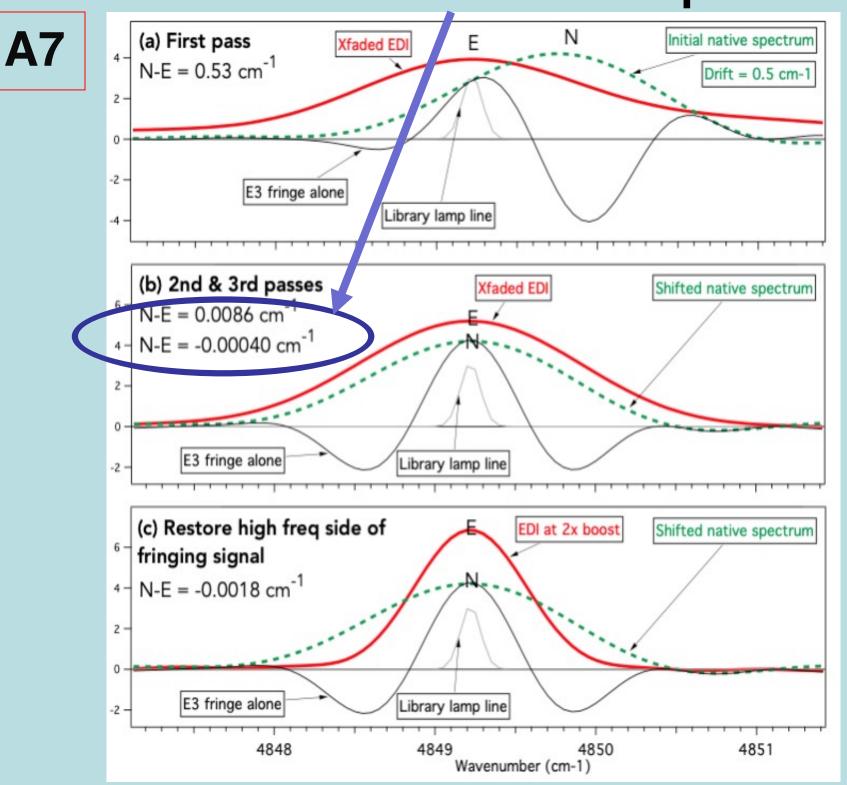


You can't repair this type of drift by a simple translation, but crossfading fixes it!

Externally Dispersed Interferometer (EDI) scheme. EDI inserted into beam prior to Hamilton echelle spectrograph at Lick Obs. (A3) Snippet of echelle spectra around sodium doublet (589 nm) showing extremely periodic interferometer comb multiplying stellar Demo of deliberately shifting raw data (dashed lines) spectrum. (A4 showing native spectral shift)--yet the crossfaded output (red peak) does not shift. (A5) Demo on the worst kind of shift: irregular and Demo on data from Hale telescope. (A7) Demo showing bipolar. 1300x stability after 3 iterations, and optional 2x resolution boost.



1300x stability: 0.53 cm⁻¹ initial insult reduced to 0.0004 cm⁻¹ in three passes



Native spectrum (row 0) shifts from -0.5 to 0.5 cm-1, but when combined with E3 fringing signal produces stable output peak position (red)

isperser Drift $\Delta x = -0.5$ cm

perser Drift $\Delta x = -0.25$ c

sperser Drift $\Delta x = 0$ cm-

perser Drift $\Delta x = 0.25$ cm

perser Drift $\Delta x = 0.5 c$

A6

Fiber A, HD962

Fiber B, HD962

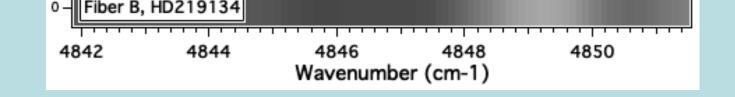
Fiber A, HD962

Fiber A, HD219134

Fiber B, HD962

Fiber A, HD219134

Fiber B. HD21913



Before

After

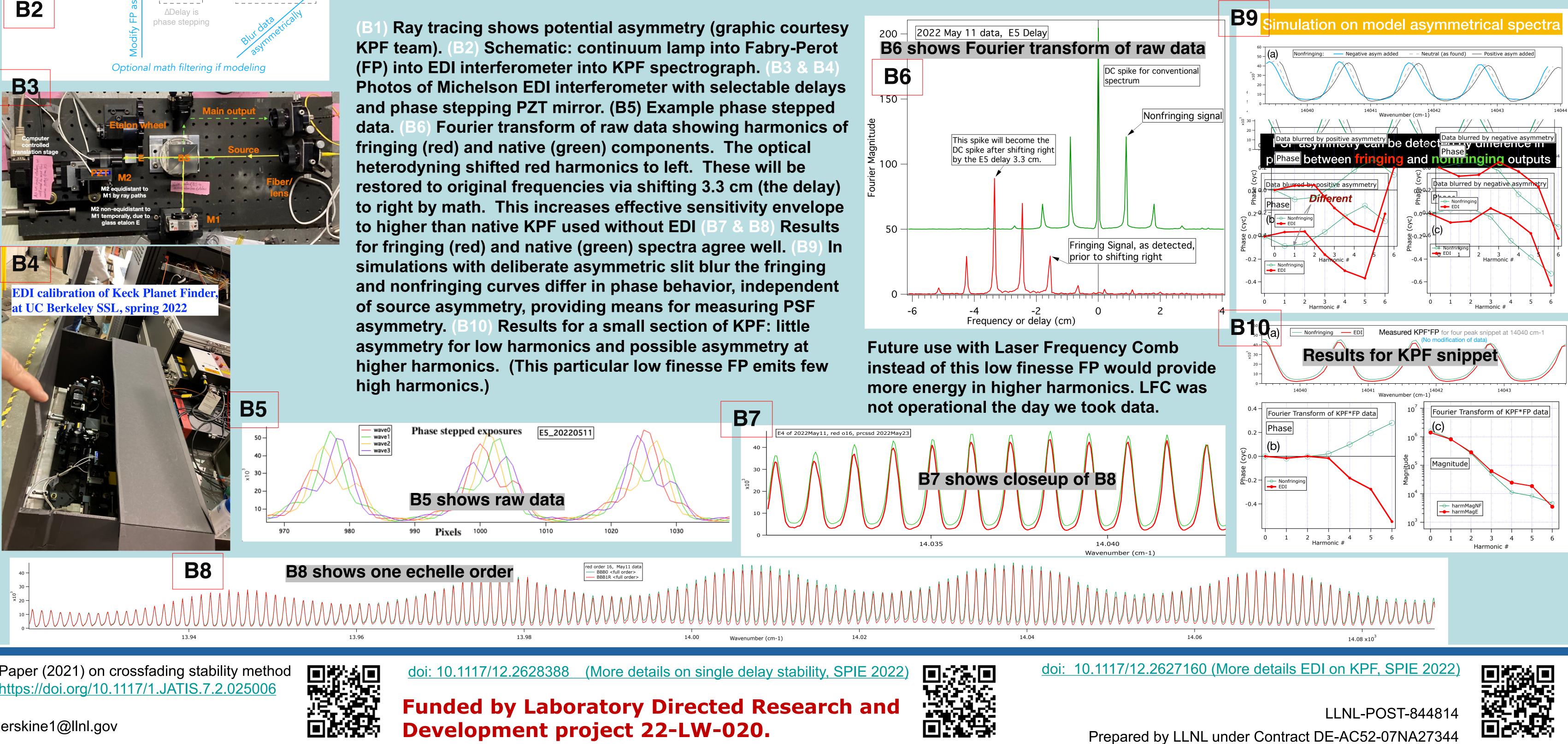
Uncorrected (Native)

Corrected (Native+Fringing)

Off-sky characterization topic

A symmetric PSF is optimal for Doppler stability. We used an EDI with the Keck Planet Finder spectrograph during off-sky engineering tests at UC Berkeley to characterize PSF asymmetry A white light illuminated Fabry-Perot generated high frequency spectral information at discrete frequencies (simplifying analysis). By comparing fringing and nonfringing signal components that observe the same source spectrum simultaneously but using different signal routes (due to the heterodyning of fringing), we are able to tease out the effect of spectrograph point spread function (PSF) asymmetry independent of source asymmetry.

Schematic: continuum lamp into Fabry-Perot KPF team). (B2 (FP) into EDI interferometer into KPF spectrograph. and phase stepping PZT mirror. (B5) Example phase stepped data. fringing (red) and native (green) components. The optical heterodyning shifted red harmonics to left. These will be restored to original frequencies via shifting 3.3 cm (the delay) to right by math. This increases effective sensitivity envelope to higher than native KPF used without EDI (B7 & B8 Results for fringing (red) and native (green) spectra agree well. 9) In simulations with deliberate asymmetric slit blur the fringing



Paper (2021) on crossfading stability method https://doi.org/10.1117/1.JATIS.7.2.025006

erskine1@llnl.gov