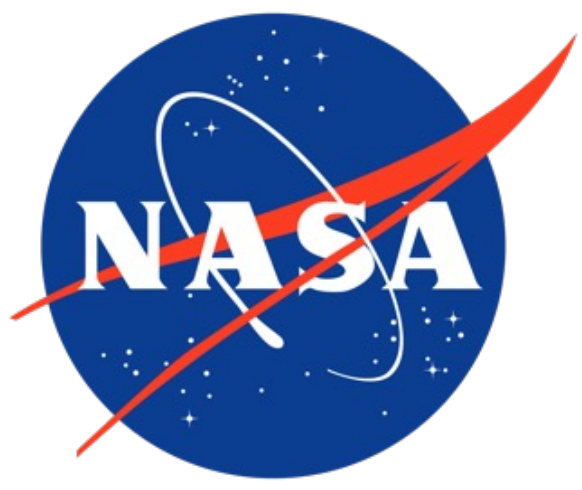


Orbits & Inclinations of Planet Host Binaries



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Introduction

- Stellar companions can greatly affect planet formation & evolution.
- Since 50% of planet hosts have companions (Horch+ 2014), we need to study how planet formation is different in single- versus multi-star systems to get the full picture!
- But the dynamical properties of planet host binaries (e.g., inclination and eccentricity) have not been well studied observationally.

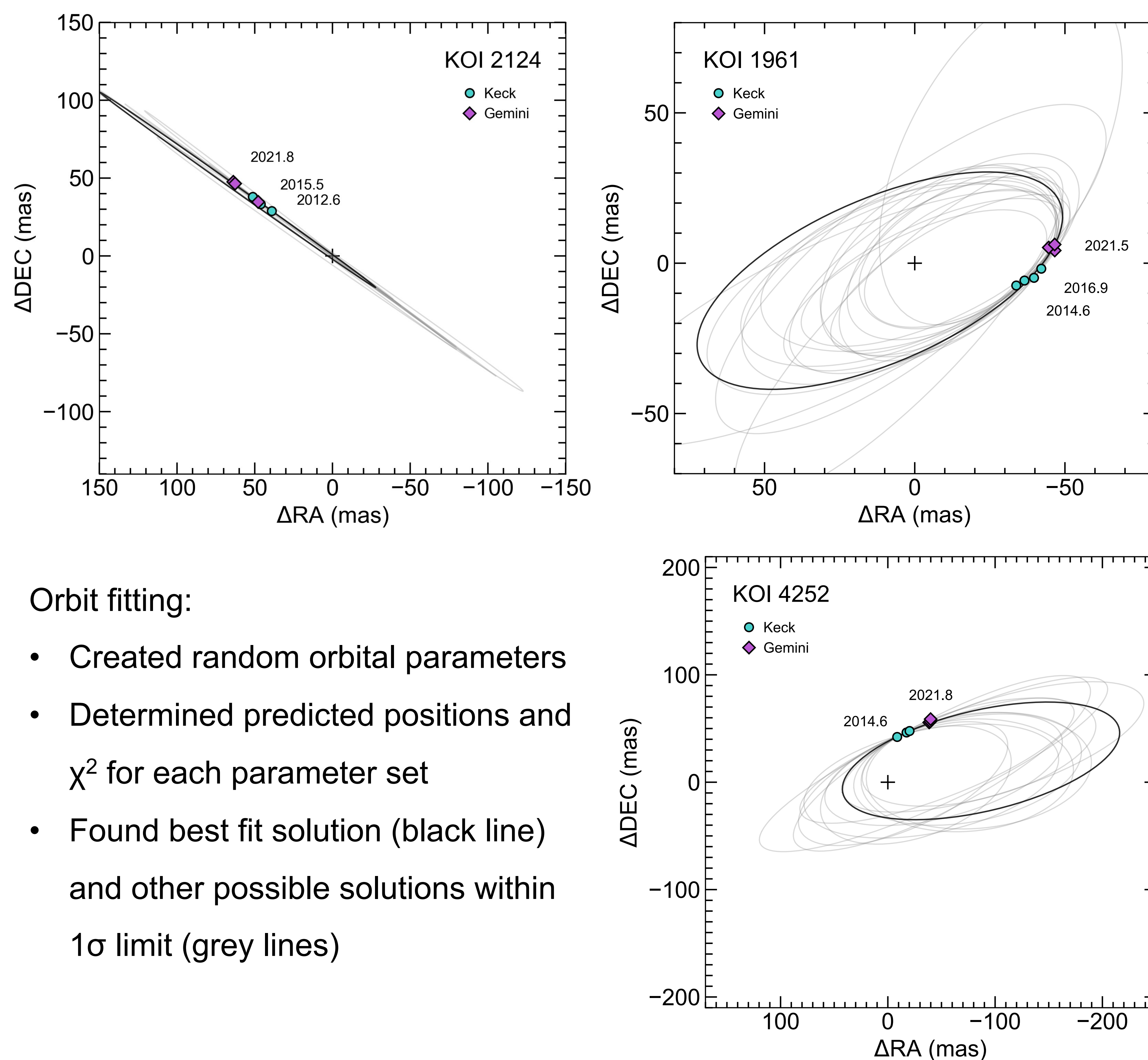
Observations

- We are monitoring 40 planet host binaries from Kepler & TESS (with $a < 100$ AU) to measure their visual & spectroscopic orbits.
- New astrometry from speckle interferometry with Gemini, + past astrometry from Keck & WIYN (Furlan+ 2017, Dupuy+ 2022, Howell+ 2021)

→ 11 systems with preliminary visual orbits



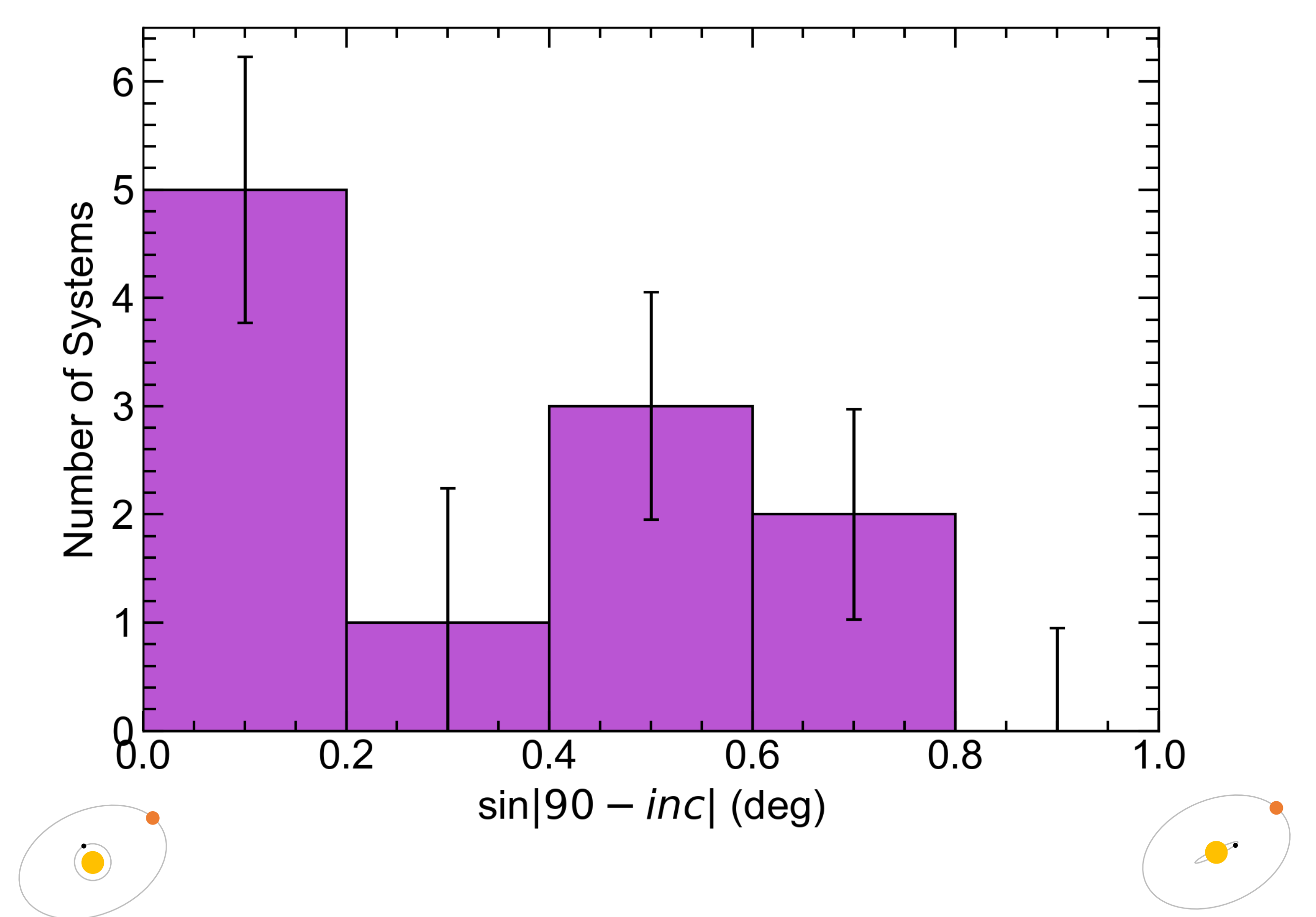
Example Visual Orbits



Orbit fitting:

- Created random orbital parameters
- Determined predicted positions and χ^2 for each parameter set
- Found best fit solution (black line) and other possible solutions within 1σ limit (grey lines)

Binary – Planet Orbital Alignment



- We found that our initial sample of close binary hosts ($a < 100$ AU) shows preferential alignment between the binary & planet orbital planes.
- This is consistent with the results for wide binary hosts ($a \sim 1000$ AU) from Christian+ 2022.

(Note – this is the *minimum* alignment, because we do not know the on-sky orientation of the planet.)

Future Work

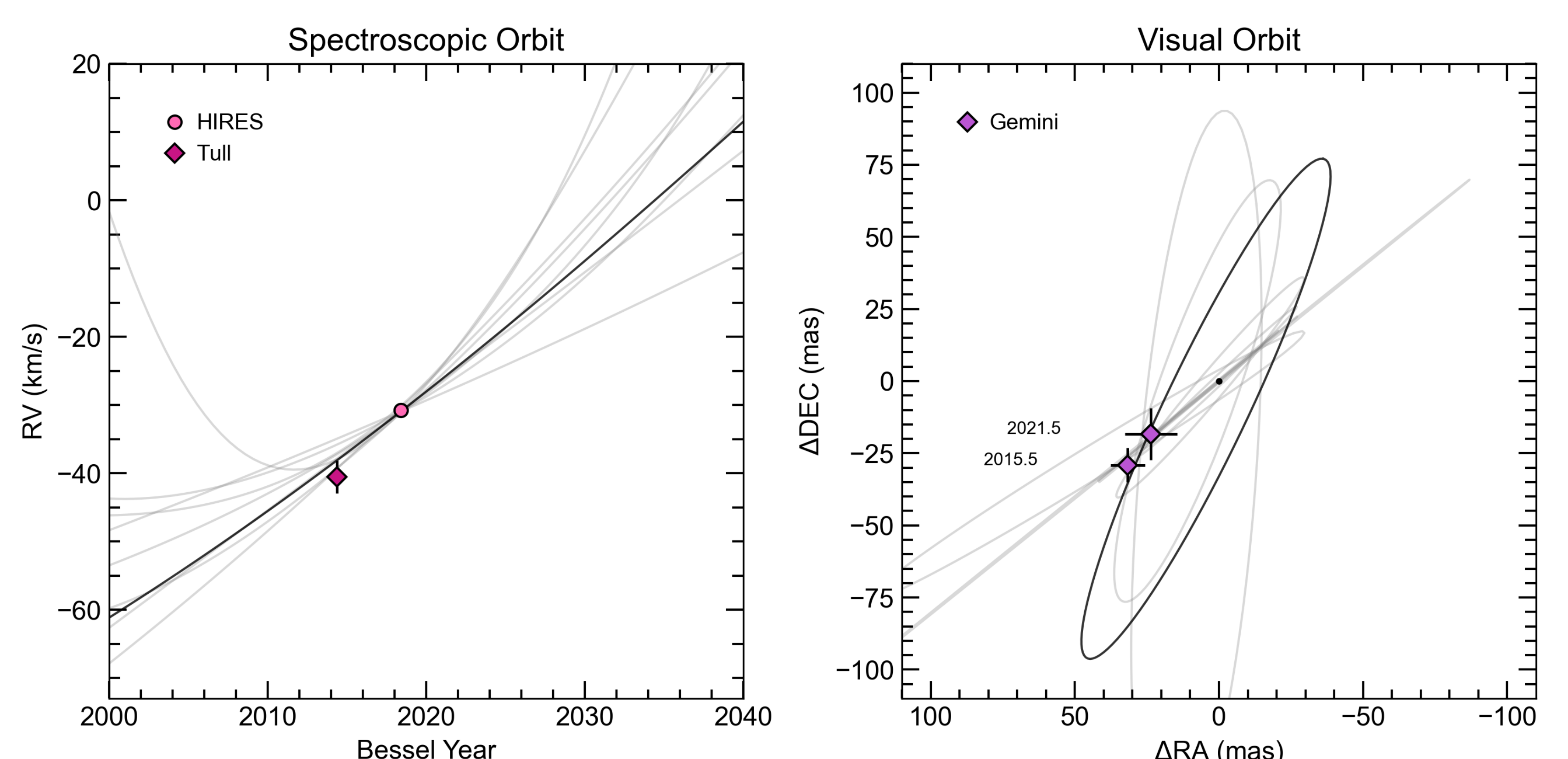
- Measuring binary-planet alignments for full sample
- Spectroscopic follow-up with NEID, HIRES, CHIRON in 2023A
- Combined spectroscopic & visual orbital solutions
- Determining binary environments amenable to planet formation!

References

Christian et al. 2022, AJ, 163, 207
 Dupuy et al. 2022, MNRAS, 512, 648
 Furlan et al. 2017, AJ, 153, 71
 Horch et al. 2014, ApJ, 795, 60
 Howell et al. 2021, AJ, 161, 164

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Preliminary VB+SB1 orbital solution for KOI 5971... but we need more observations to constrain the orbit!