

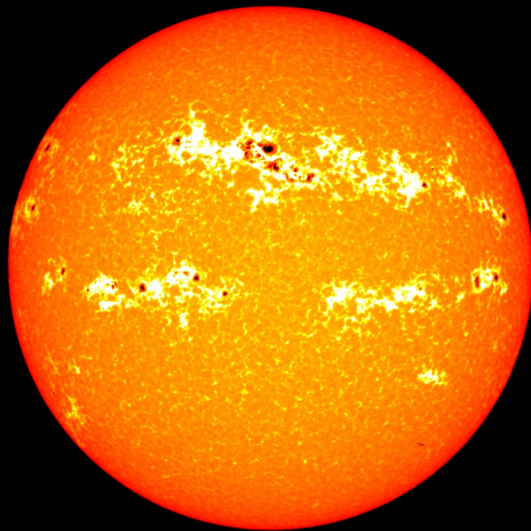
Correcting for Unresolved Companions in *Kepler* Occurrence Rates

KTSKTP2 – Wednesday, 2/5/2025

Galen Bergsten, David Ciardi, Catherine Clark, Jessie Christiansen,
Ilaria Pascucci, Kevin Hardegree-Ullman, Mike Lund



Facula



Dracula



NOT demographics in multi-star systems

see e.g., Sullivan et al. (2023); Clark et al. (2024)

Dominic Oddo (#1.15)

Catherine Clark (#1.28)

Daniel Fabrycky (#2.04)

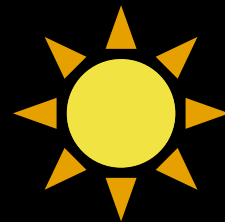
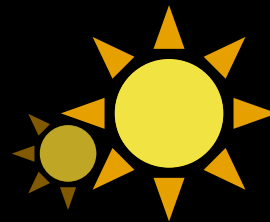
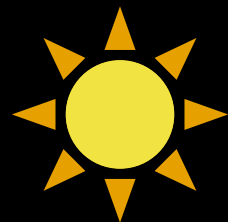
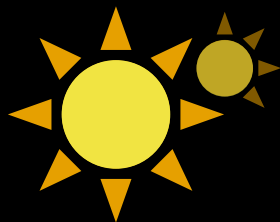
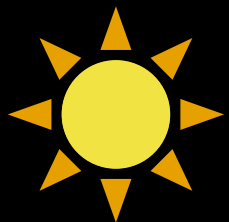
Kendall Sullivan (#4.10)

Colin Littlefield (#5.07)

Instead, demographics assuming single stars...

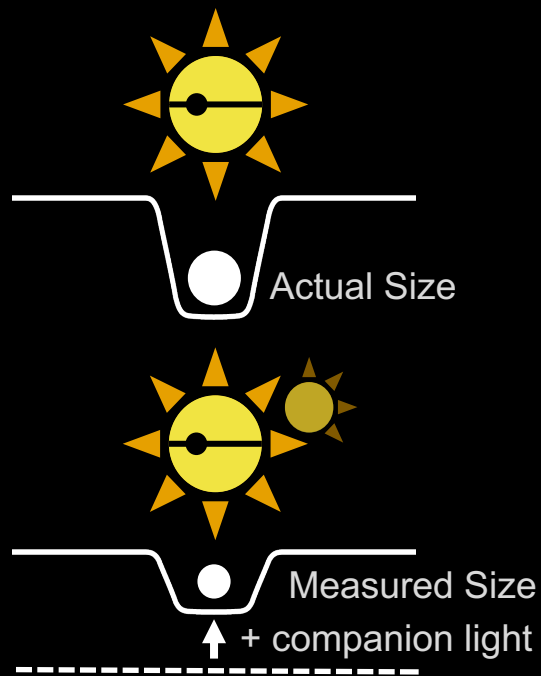
(... that aren't always singles)

see e.g., Savel et al. (2020)



Unresolved stellar companions mess with our measurements of:

Planet Radii



Detection Efficiency

10% Complete to
 $1 R_{\oplus}$ Planets

+

+50% underestimation of
planet radii

=

10% Complete to
 $1.5 R_{\oplus}$ Planets

Occurrence Rates

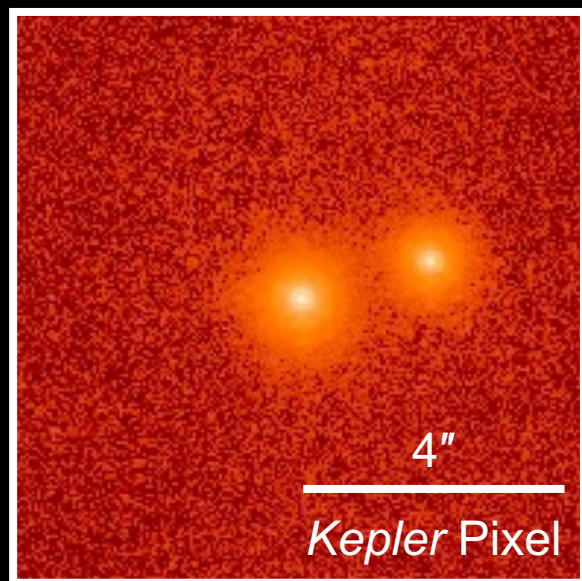
already a
hot mess



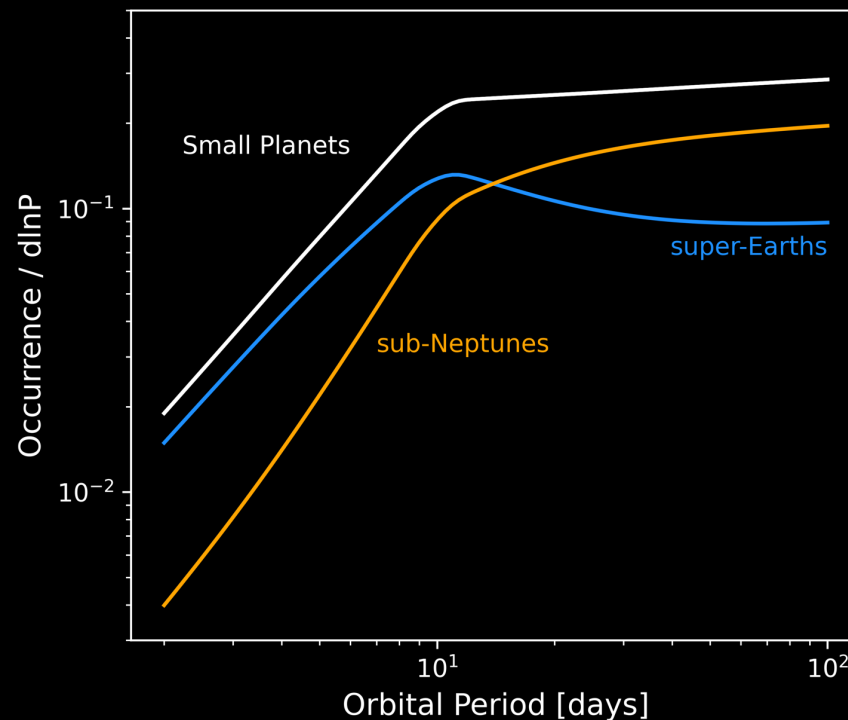
a considerably
hotter and
messier mess

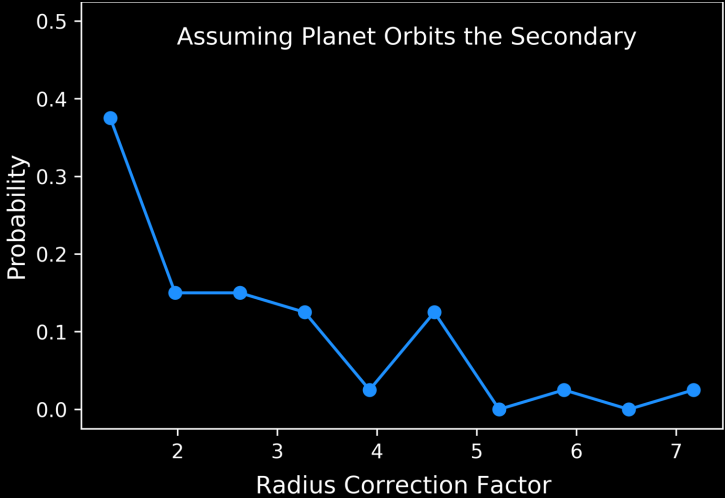
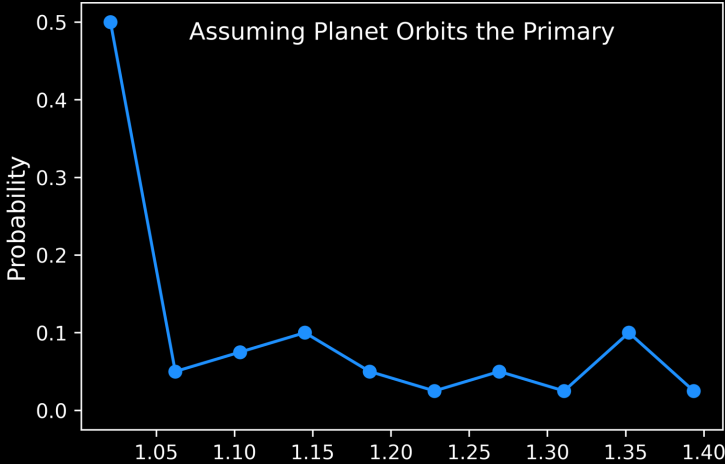
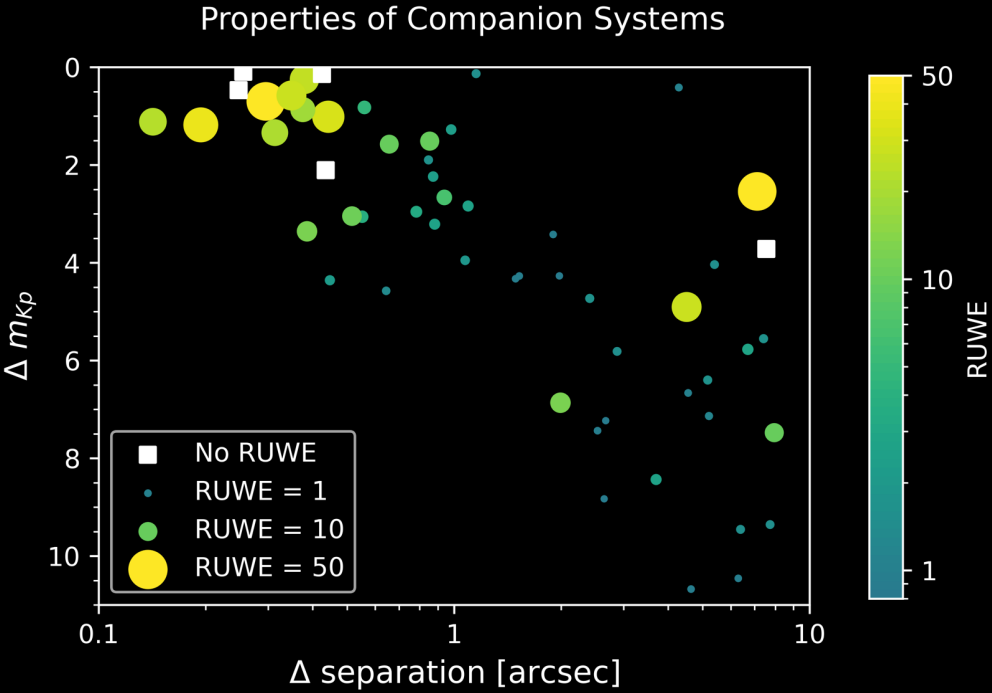
200 *Kepler* stars

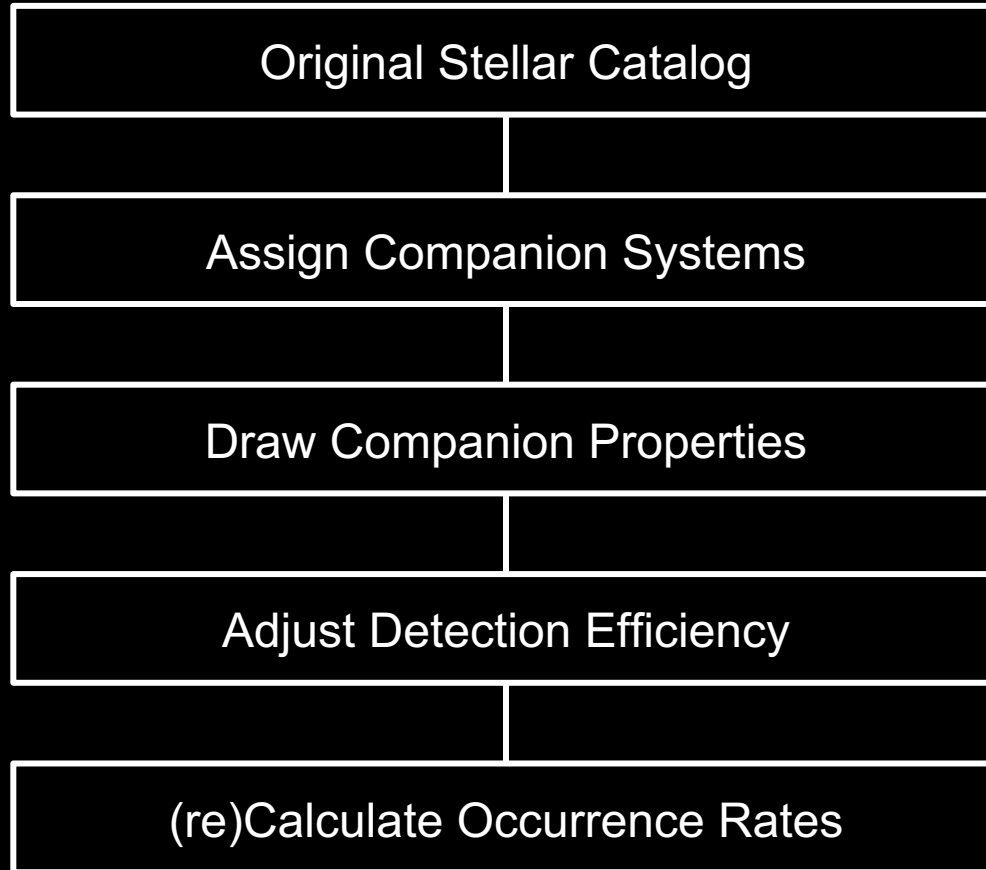
40 systems with
(previously) unresolved
companions



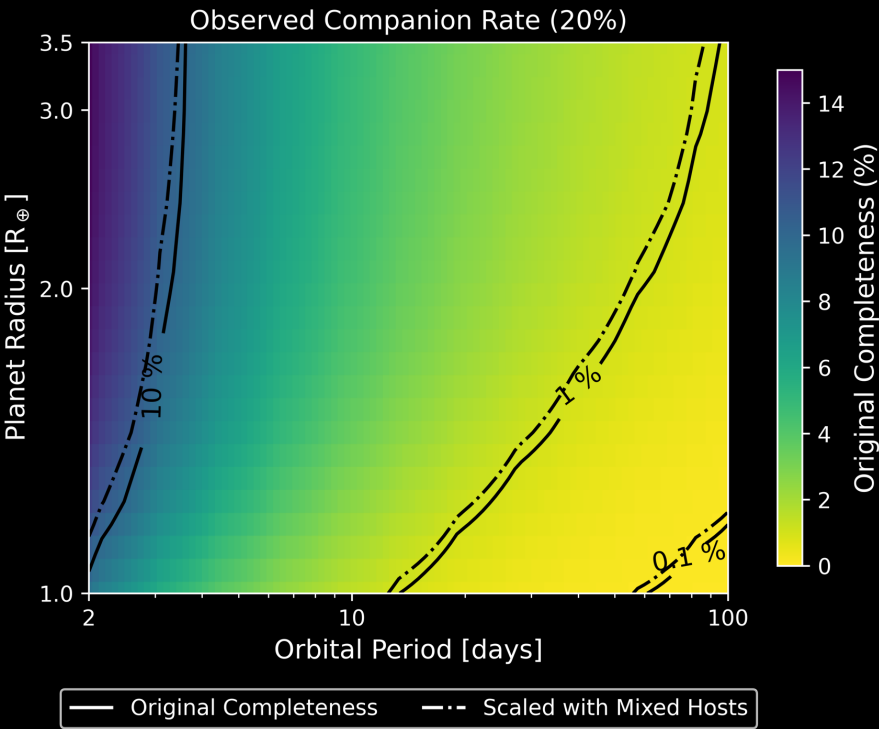
Occurrence Rates from
Bergsten et al. (2022)



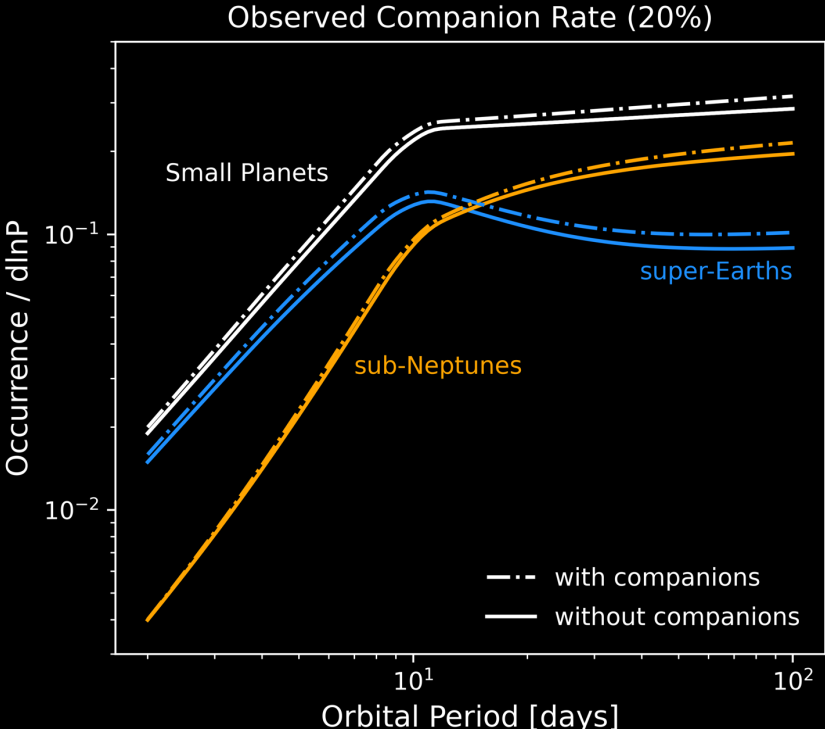


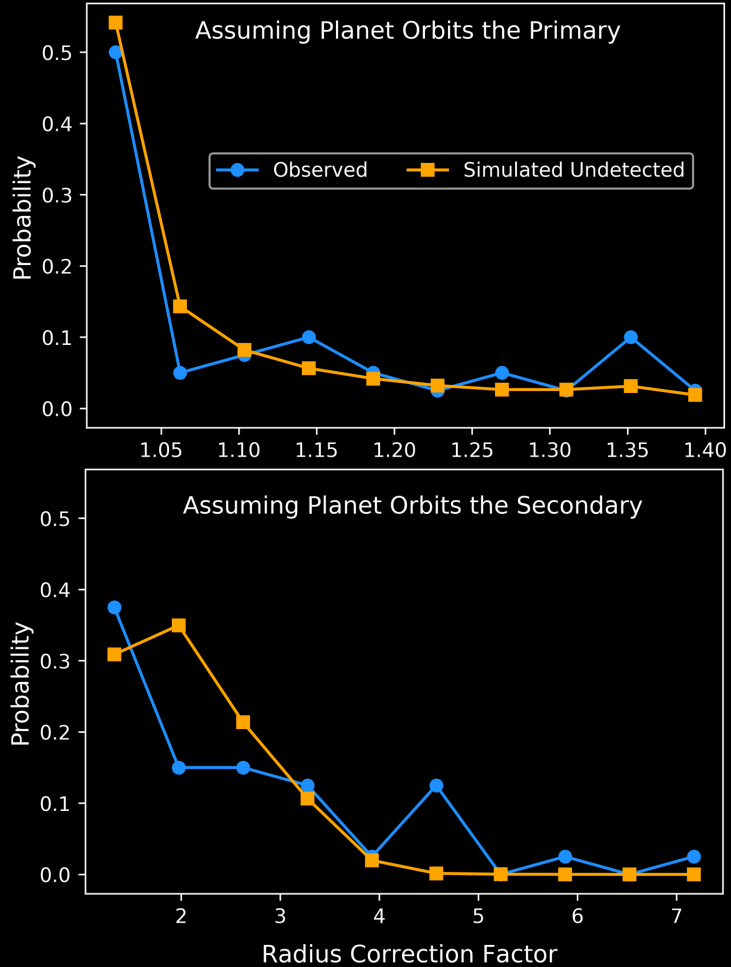
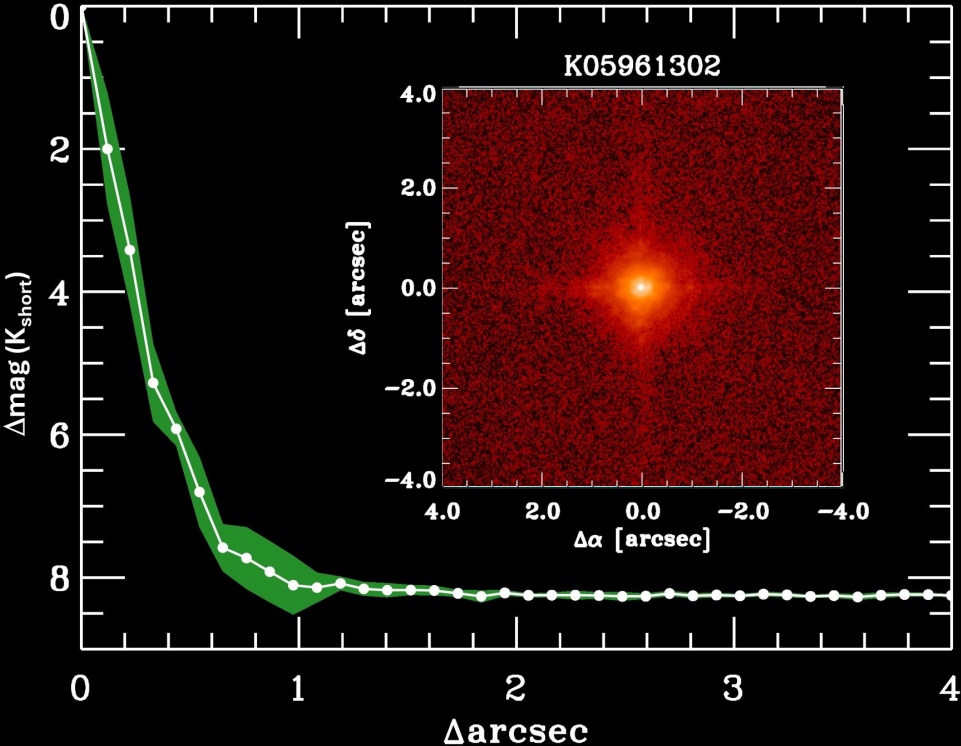


Adjusted Detection Efficiency

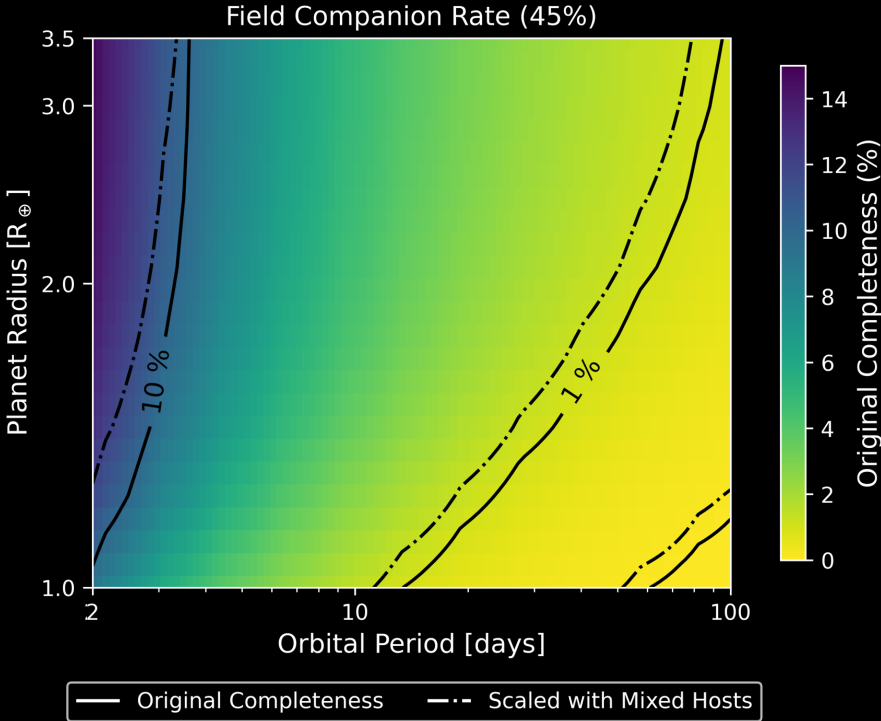


Recalculated Occurrence Rates

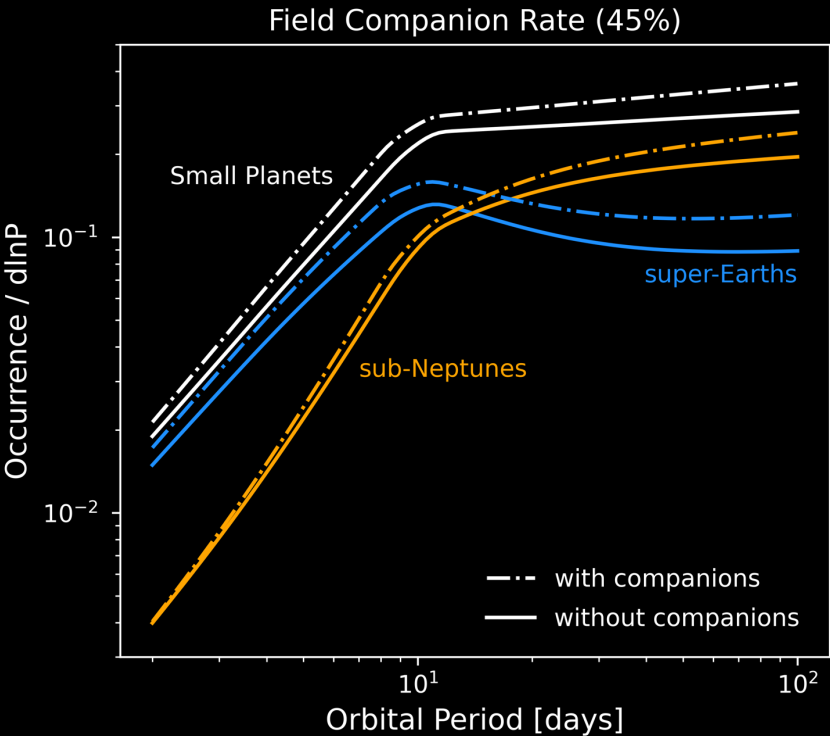


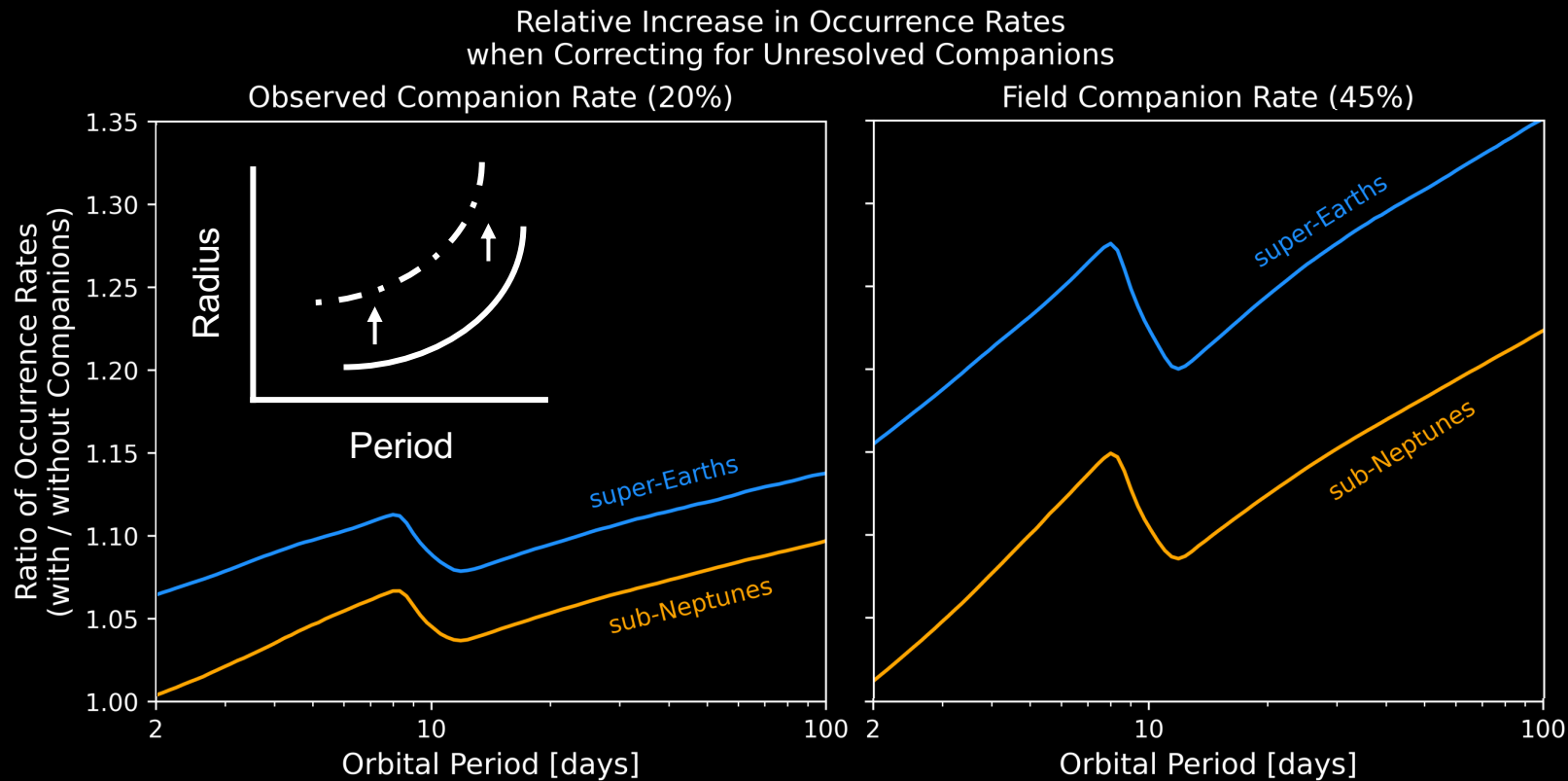


Adjusted Detection Efficiency



Recalculated Occurrence Rates





bump @ 10 days due to (insignificant) difference in free parameters

η_{\oplus} – the frequency of Earth-sized, Habitable Zone planets

	η_{\oplus} [%]	Relative Increase
Assuming Singles	$8.4^{+2.2}_{-3.2}$	--
Observed Rate (20%)	$9.9^{+2.8}_{-3.9}$	1.18
Field Rate (45%)	$12.3^{+3.5}_{-4.8}$	1.46

Doesn't matter which models you use!

Correcting for unresolved companions
can lead to an average 1.2-1.4x
relative increase in η_{\oplus}

