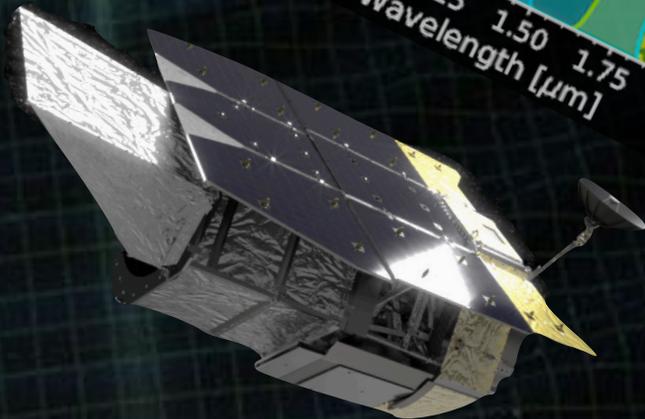
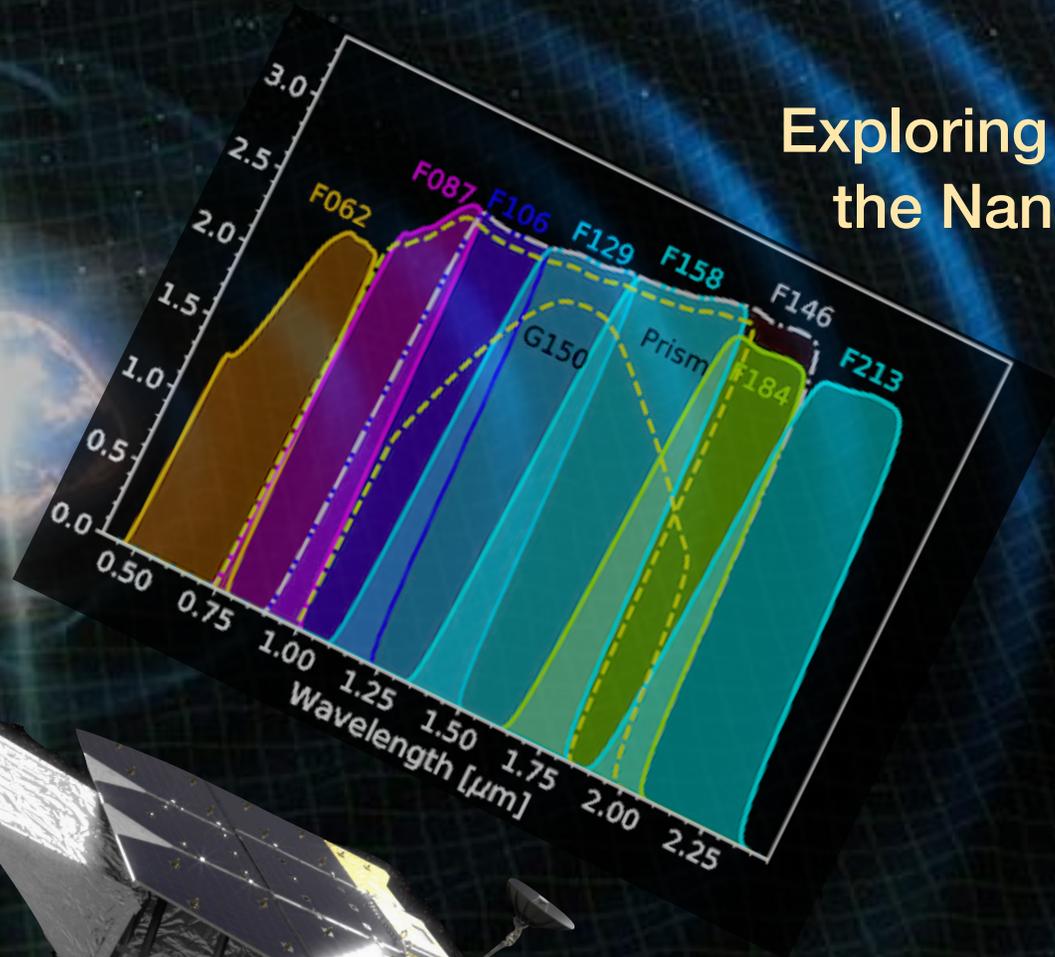


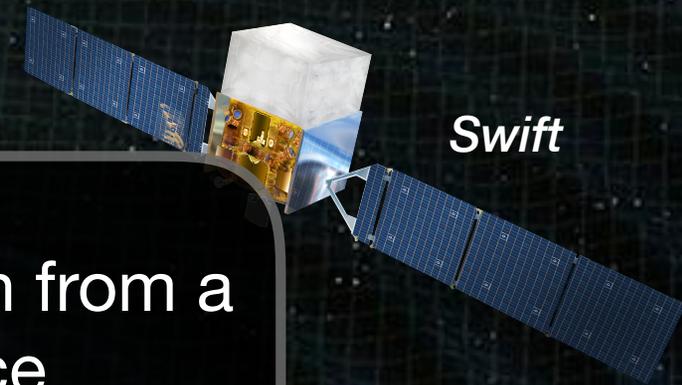
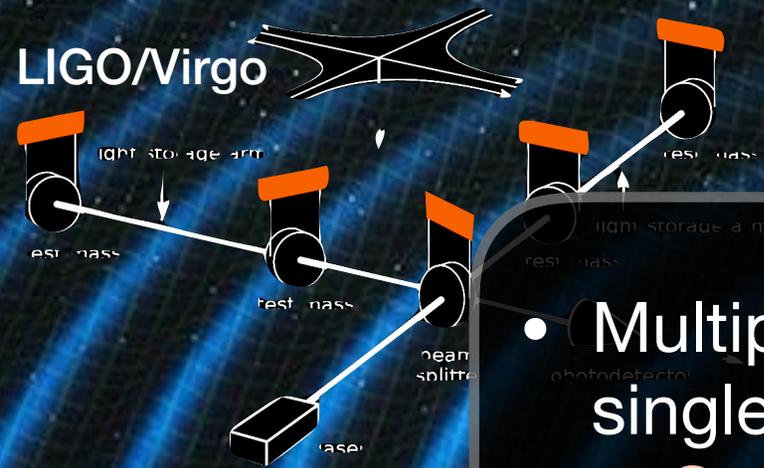
Multi-messenger Astronomy with the Roman Space Telescope

Exploring the Transient Universe with the Nancy Grace Roman Telescope
Feb 9, 2022

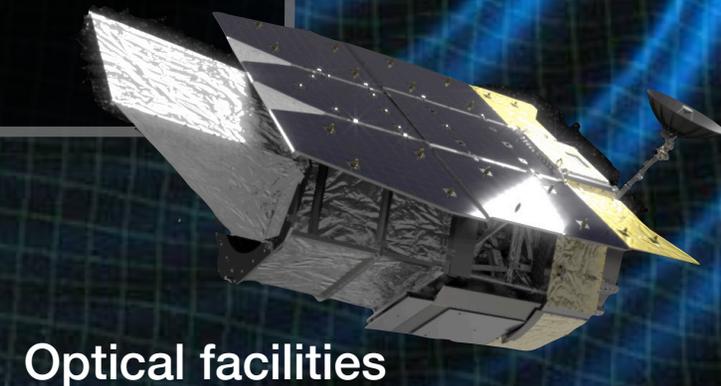


Jennifer Barnes
KITP, UC Santa Barbara

Multi-messenger astronomy defined

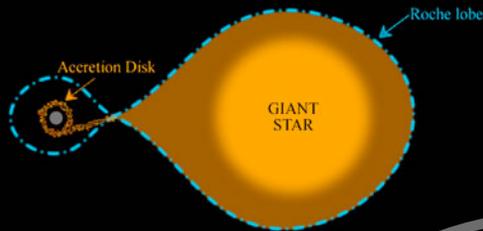


- Multiple types of emission from a single astrophysical source
 - Gravitational waves (GW)
 - Electromagnetic radiation (EM)
 - Neutrinos
- Focus today: binary neutron star and neutron star-black hole mergers

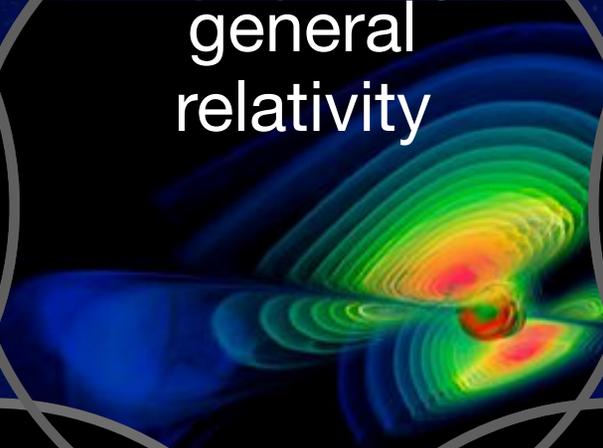


Compact object mergers overlap with many important questions

stellar binary evolution



general relativity



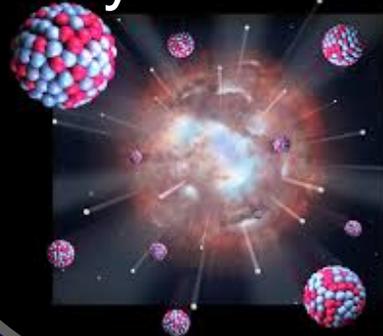
astrophysical jets



neutron star equation of state

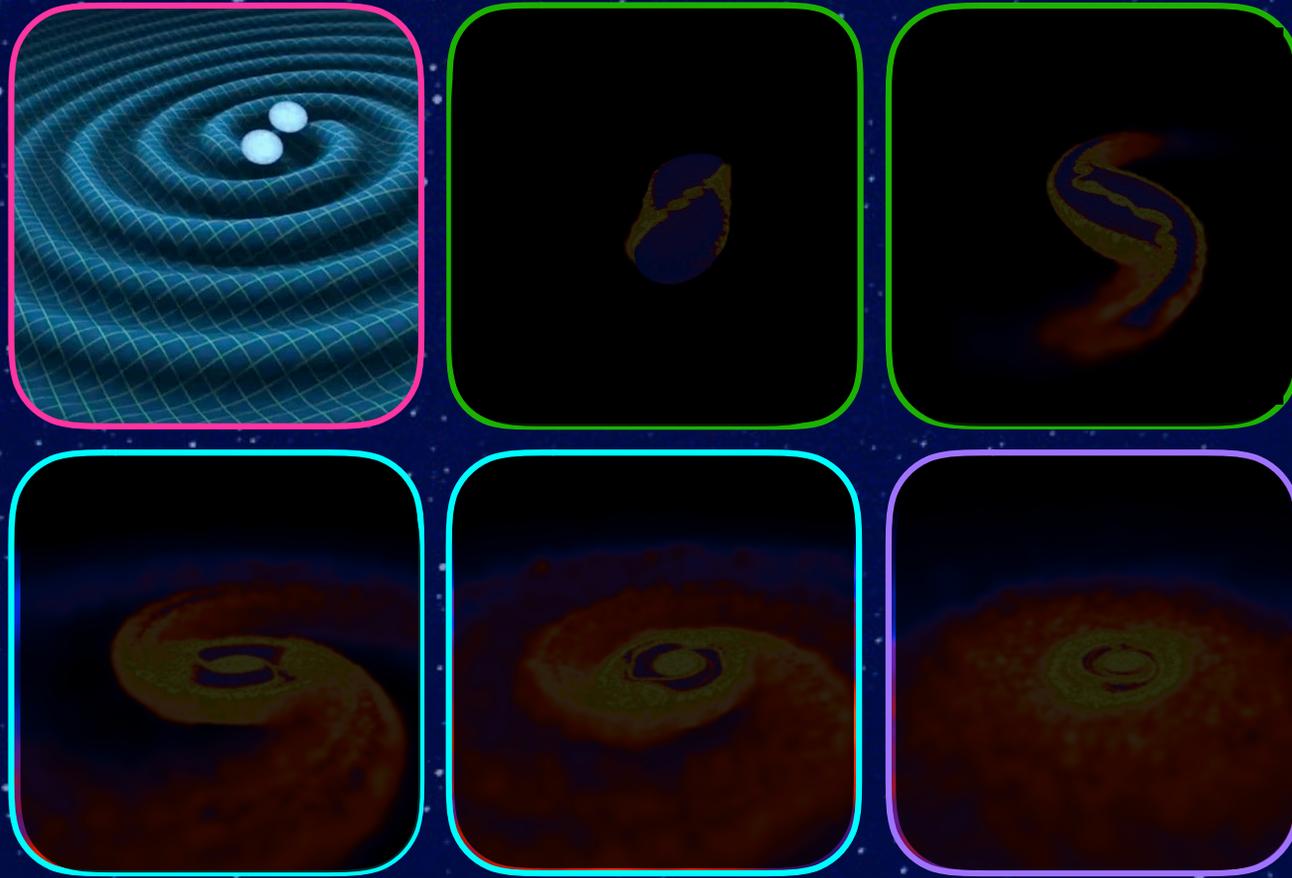


origin of the heavy elements



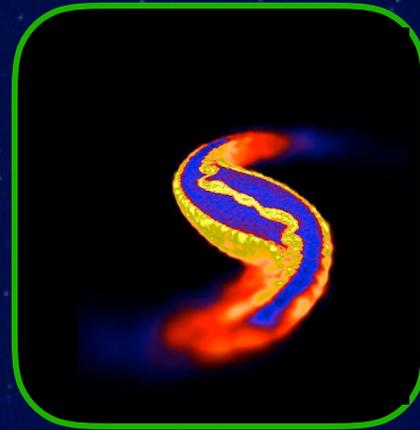
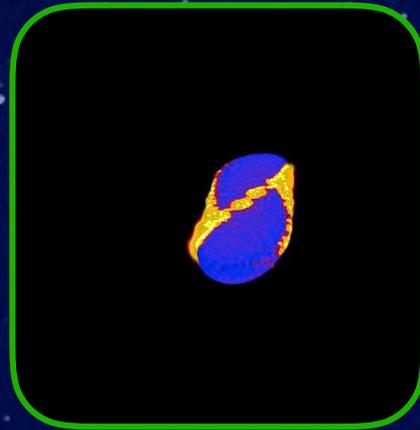
Compact object mergers are a stellar danse macabre

final orbits:
strong GW
source

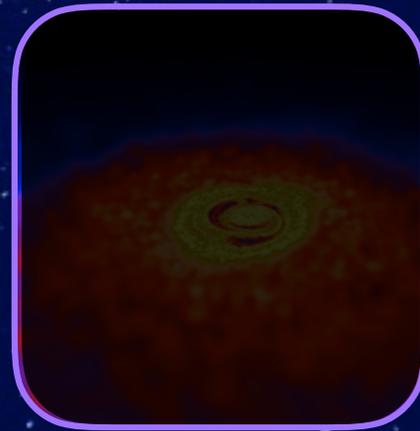
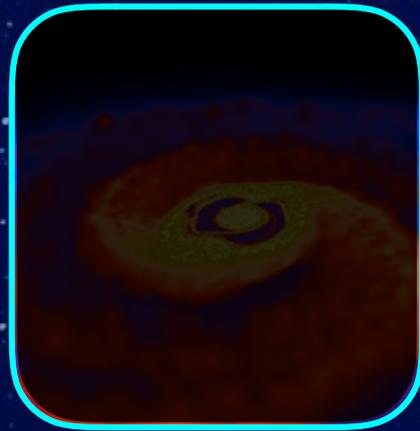
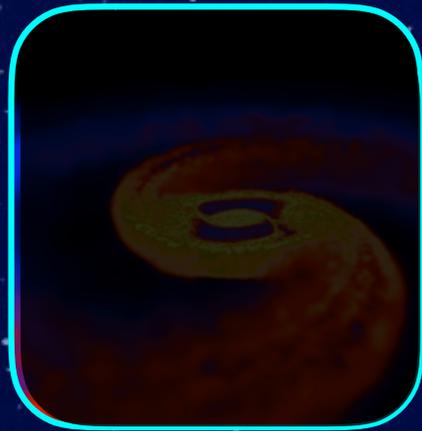


Compact object mergers are a stellar danse macabre

final orbits:
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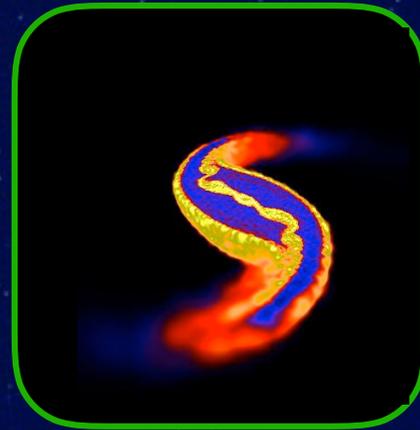
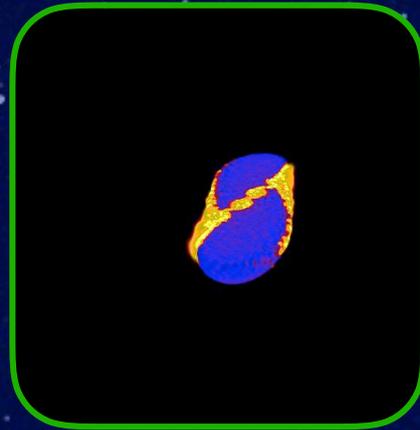
merger: neutron
star disrupts,
central remnant
forms



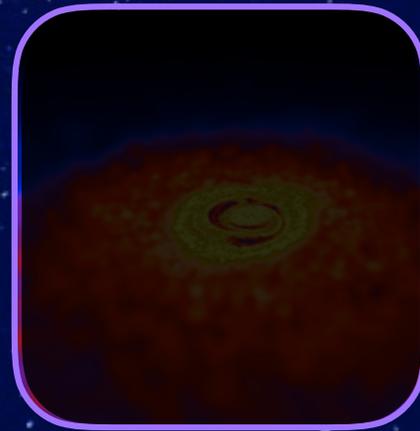
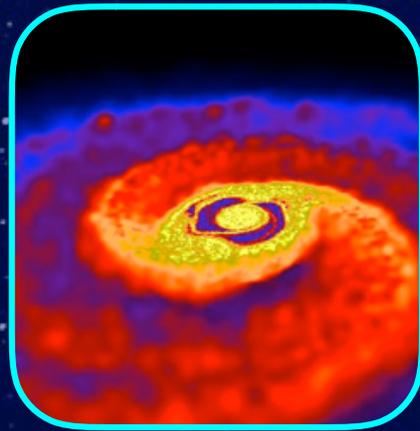
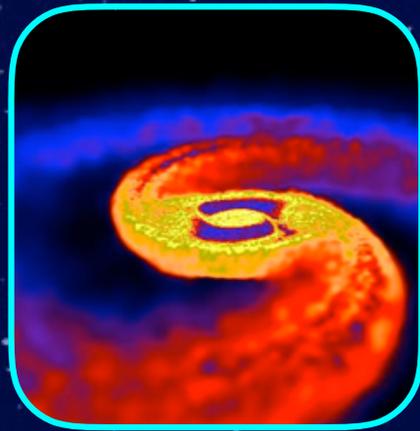
Compact object mergers are a stellar danse macabre

merger: neutron star disrupts, central remnant forms

final orbits: strong GW source

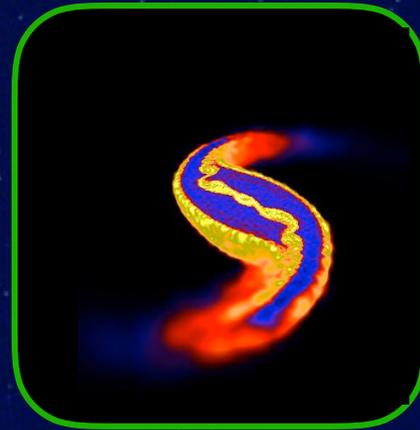
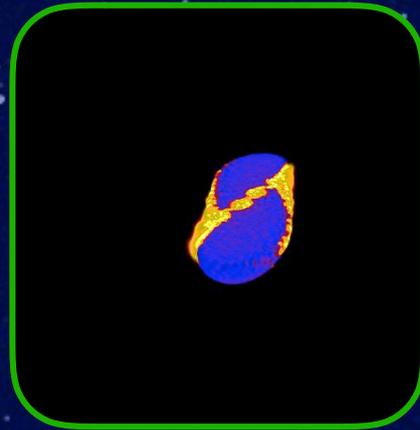


ejecta: some material escapes; some is bound



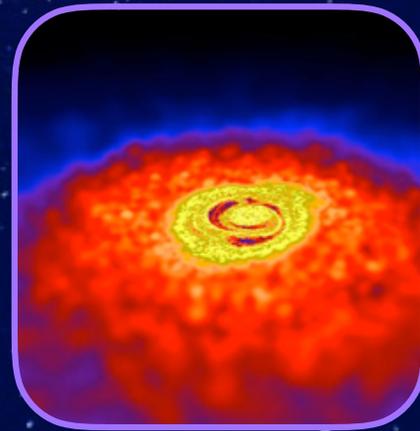
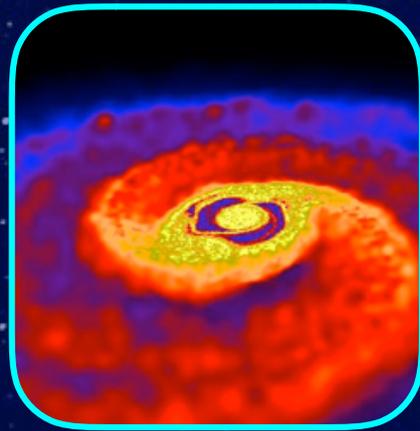
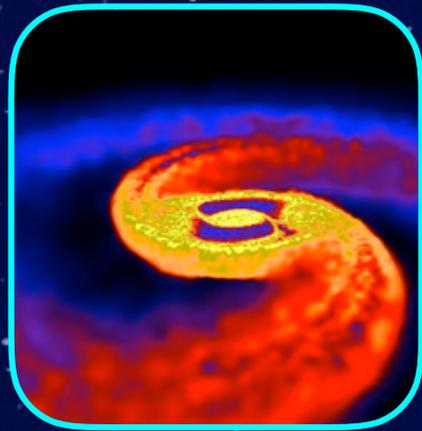
Compact object mergers are a stellar danse macabre

final orbits:
strong GW
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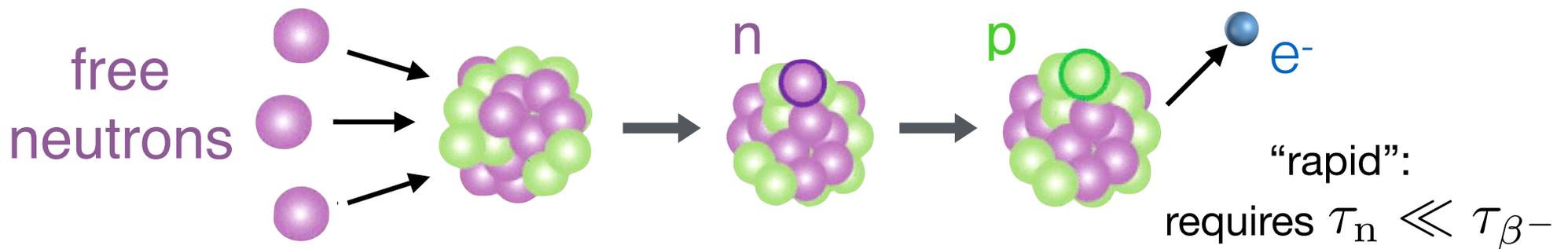
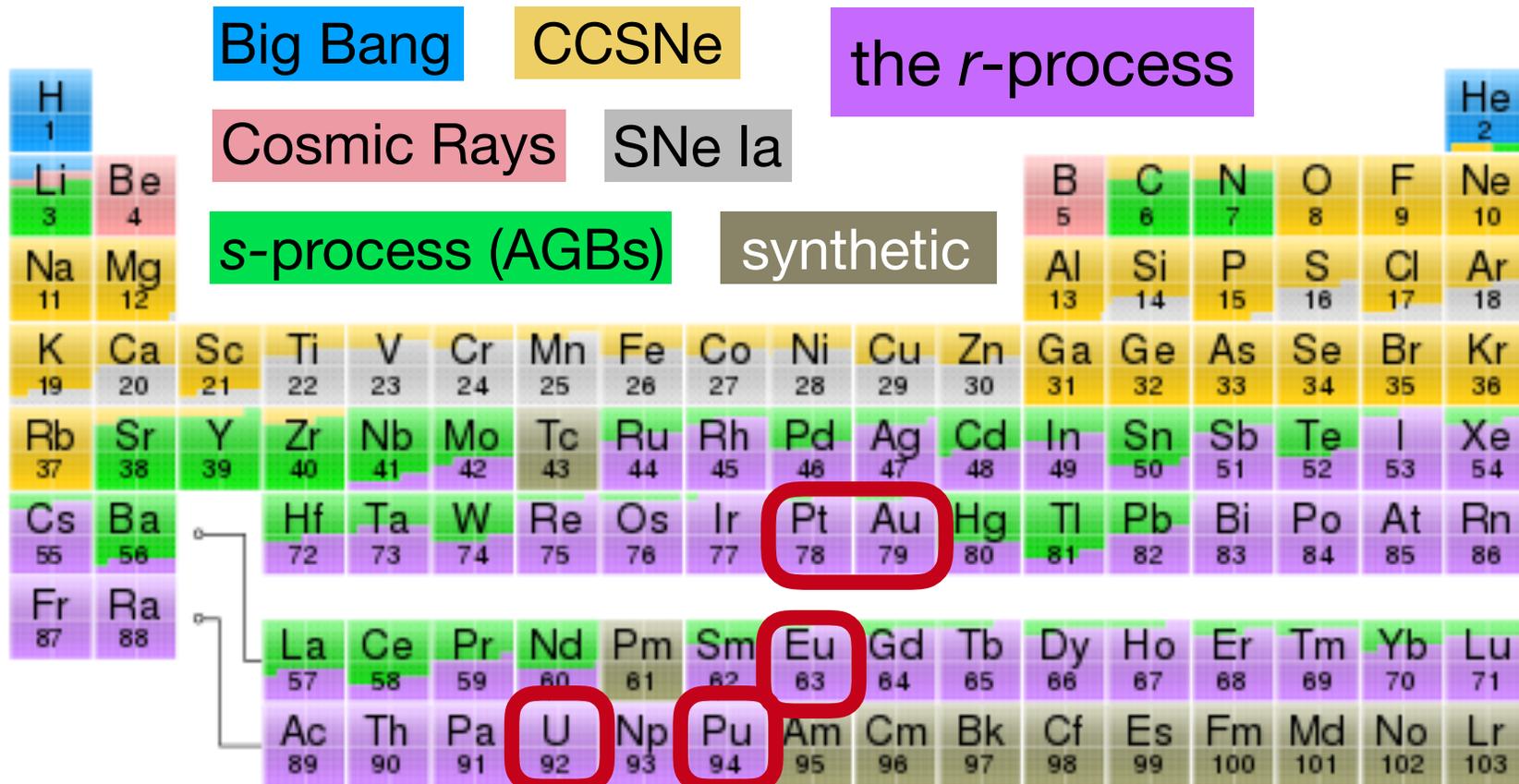
merger: neutron
star disrupts,
central remnant
forms

ejecta: some
material
escapes; some
is bound



final: a central
NS or BH, an
accretion disk,
unbound ejecta

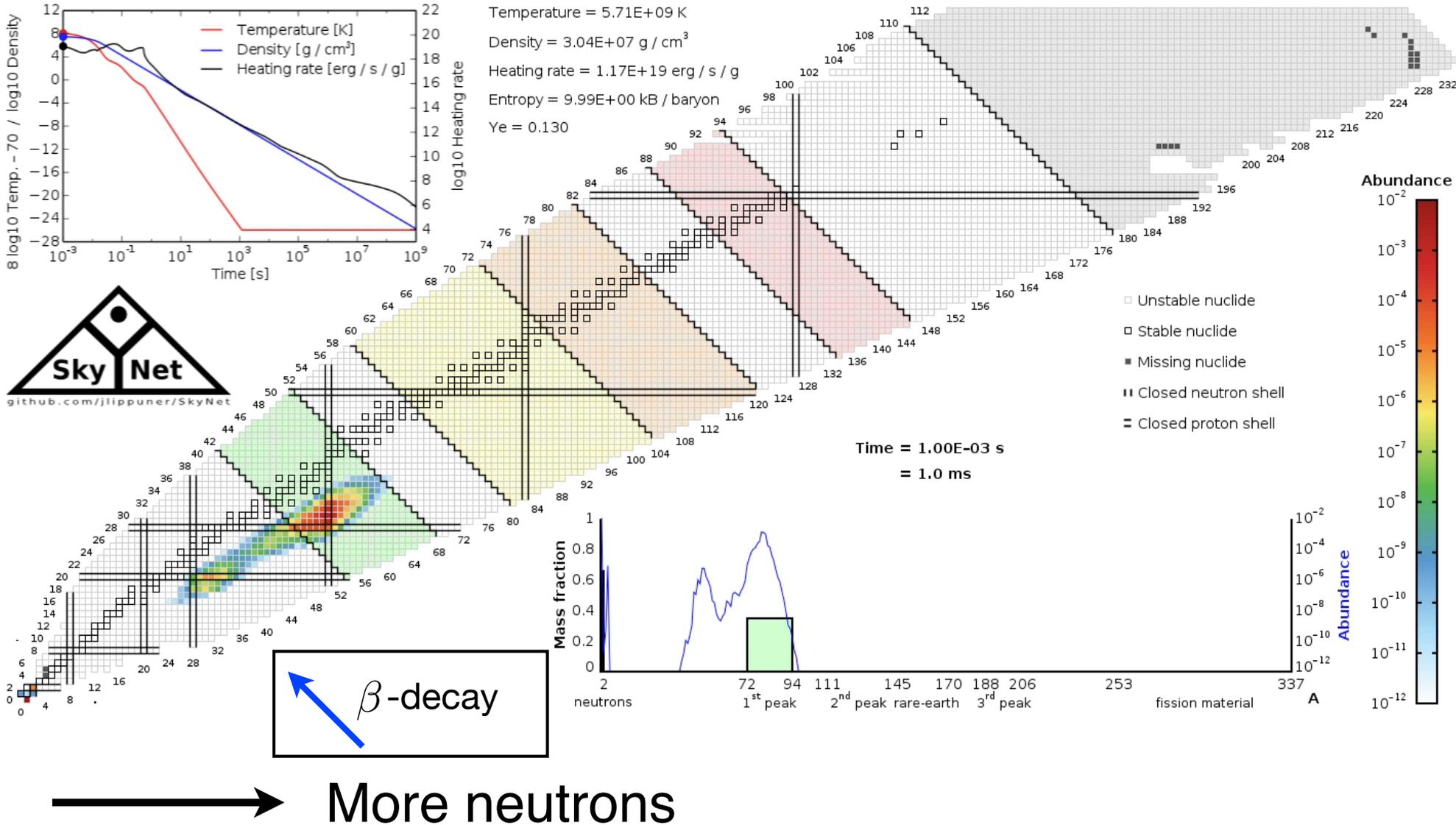
The *r*-process produces ~half of elements heavier than Fe



The r -process produces ~half of elements heavier than Fe

courtesy J. Lippuner

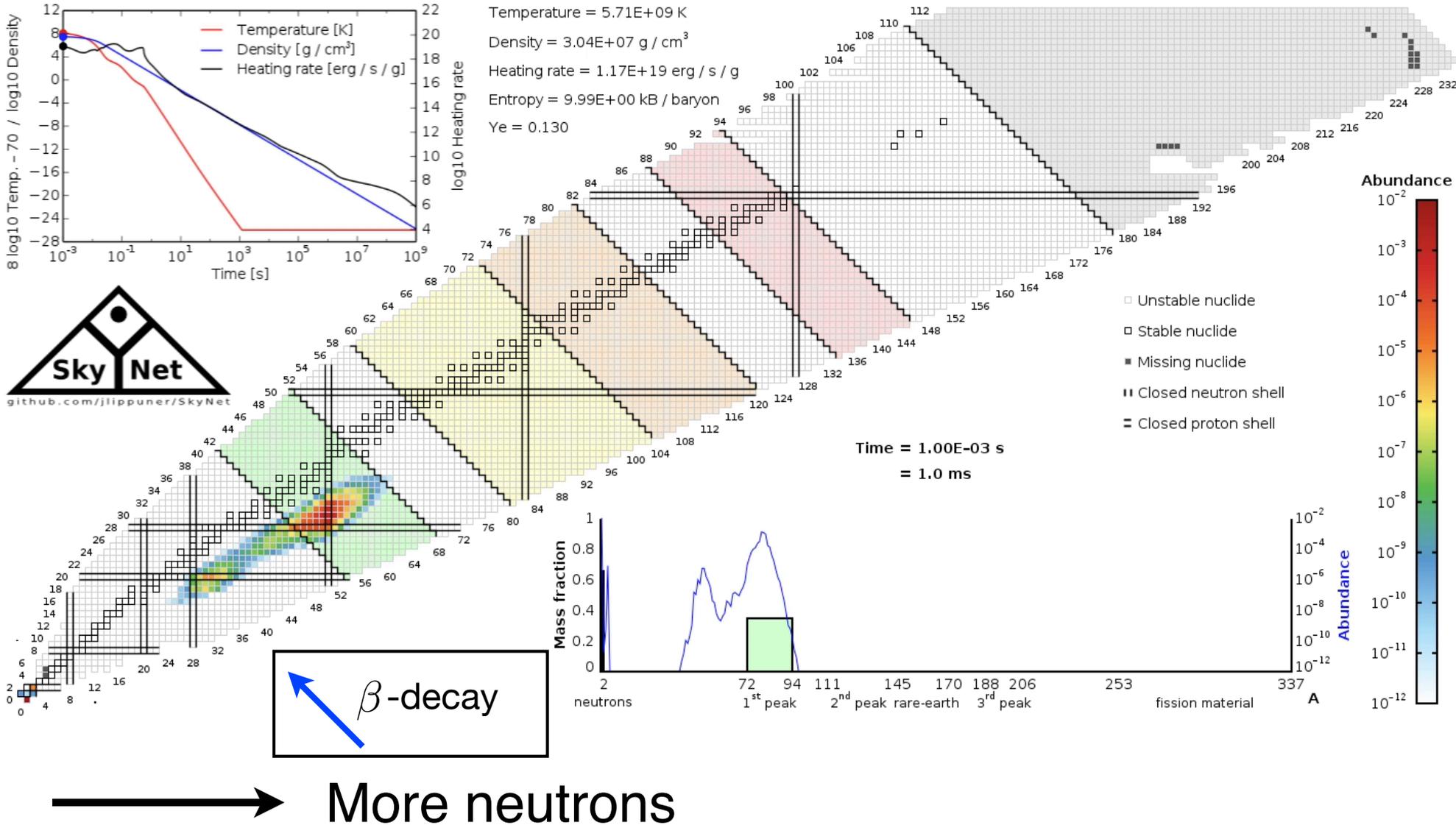
More protons
↑



The r -process produces ~half of elements heavier than Fe

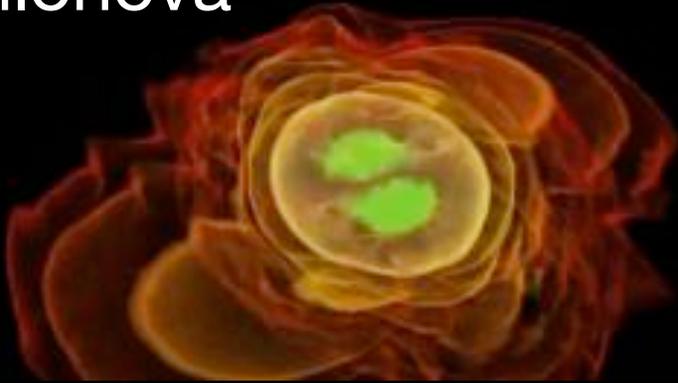
courtesy J. Lippuner

More protons



The decay of synthesized elements powers a “kilonova”

“kilonova”

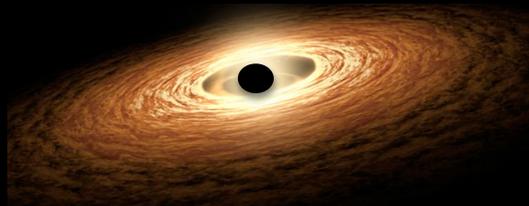


- Mildly relativistic **neutron-rich** unbound material
 - Synthesis of heavy elements
- An expanding cloud heated by radioactive decays

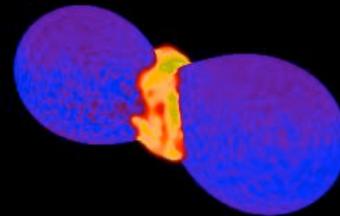
tidally
stripped



disk
outflows

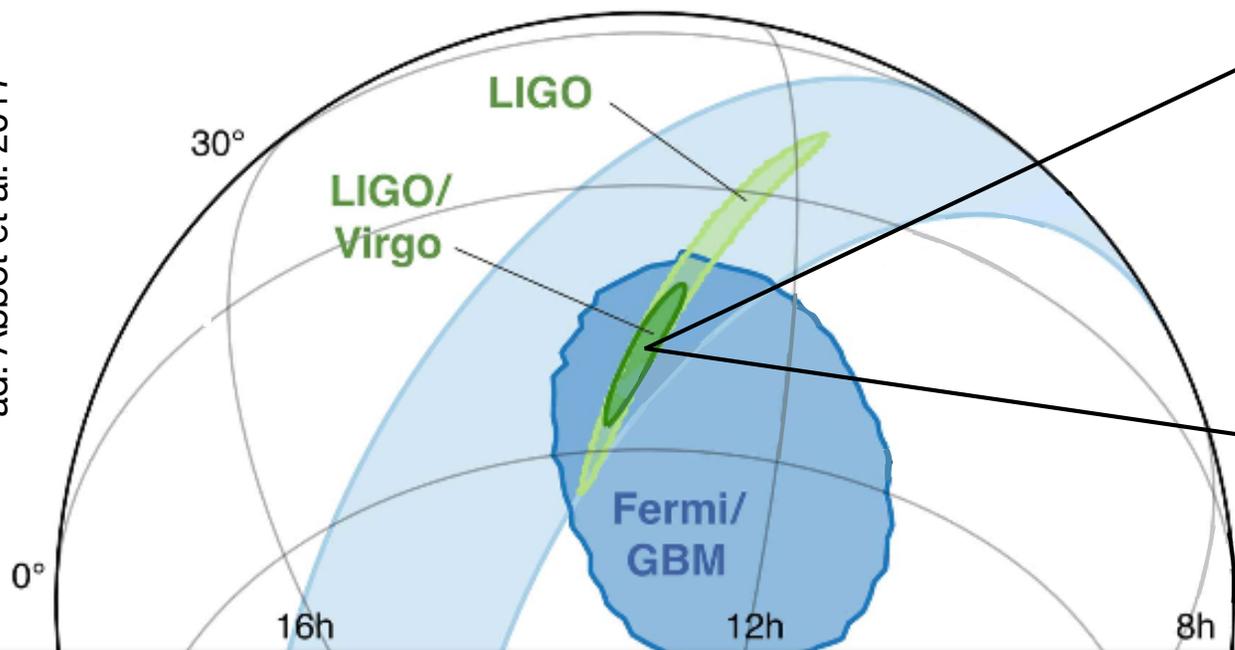


dynamically
squeezed

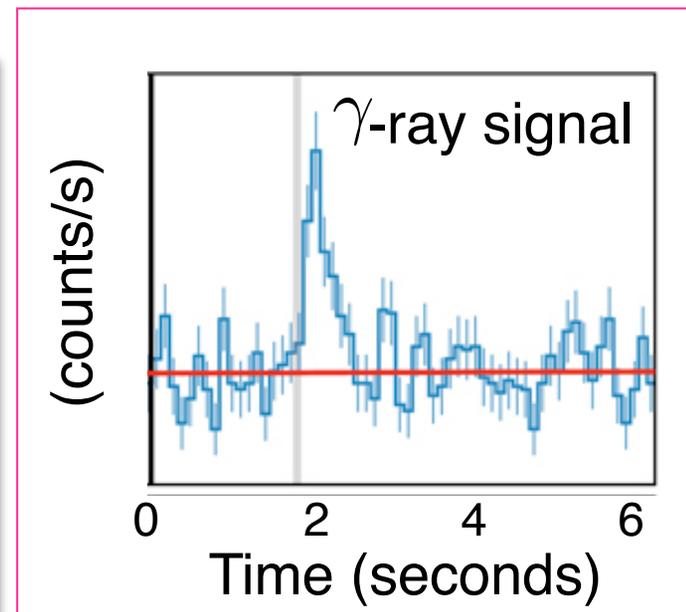
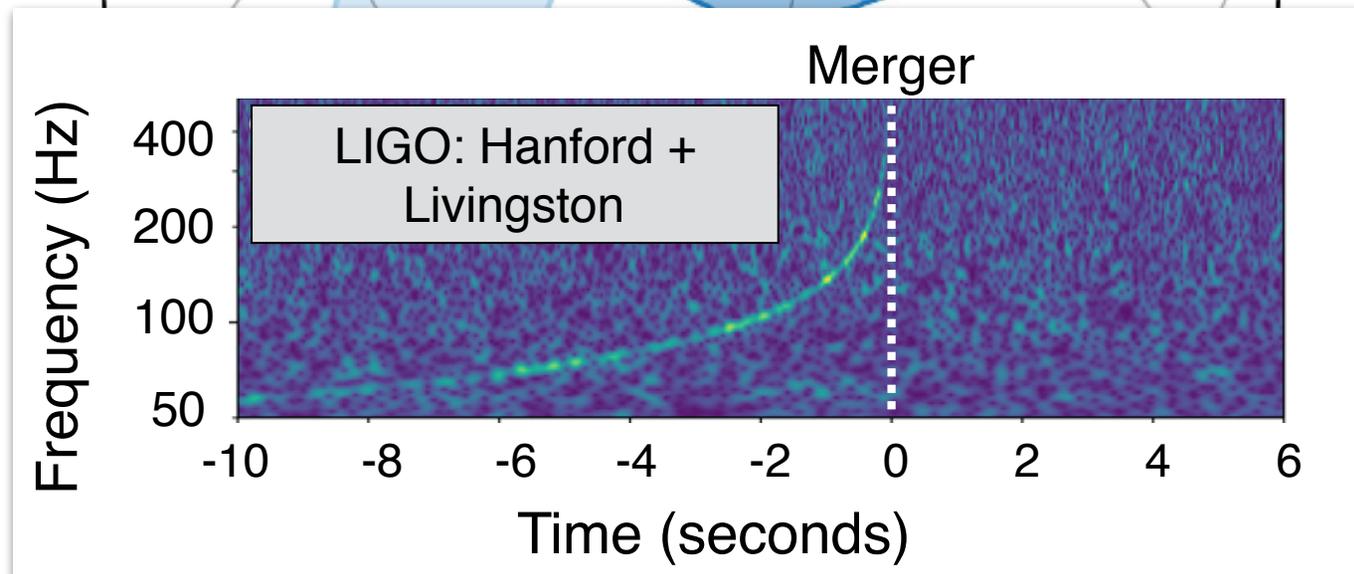
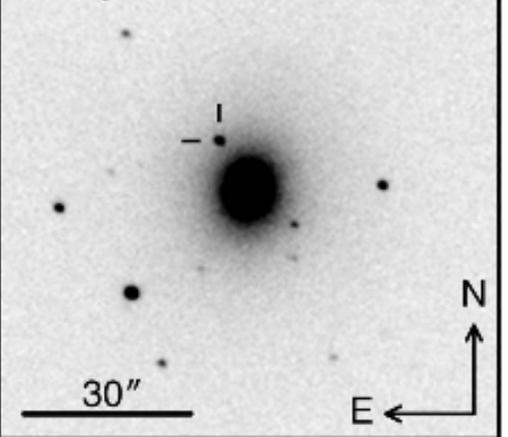


GW170817: the first neutron-star merger

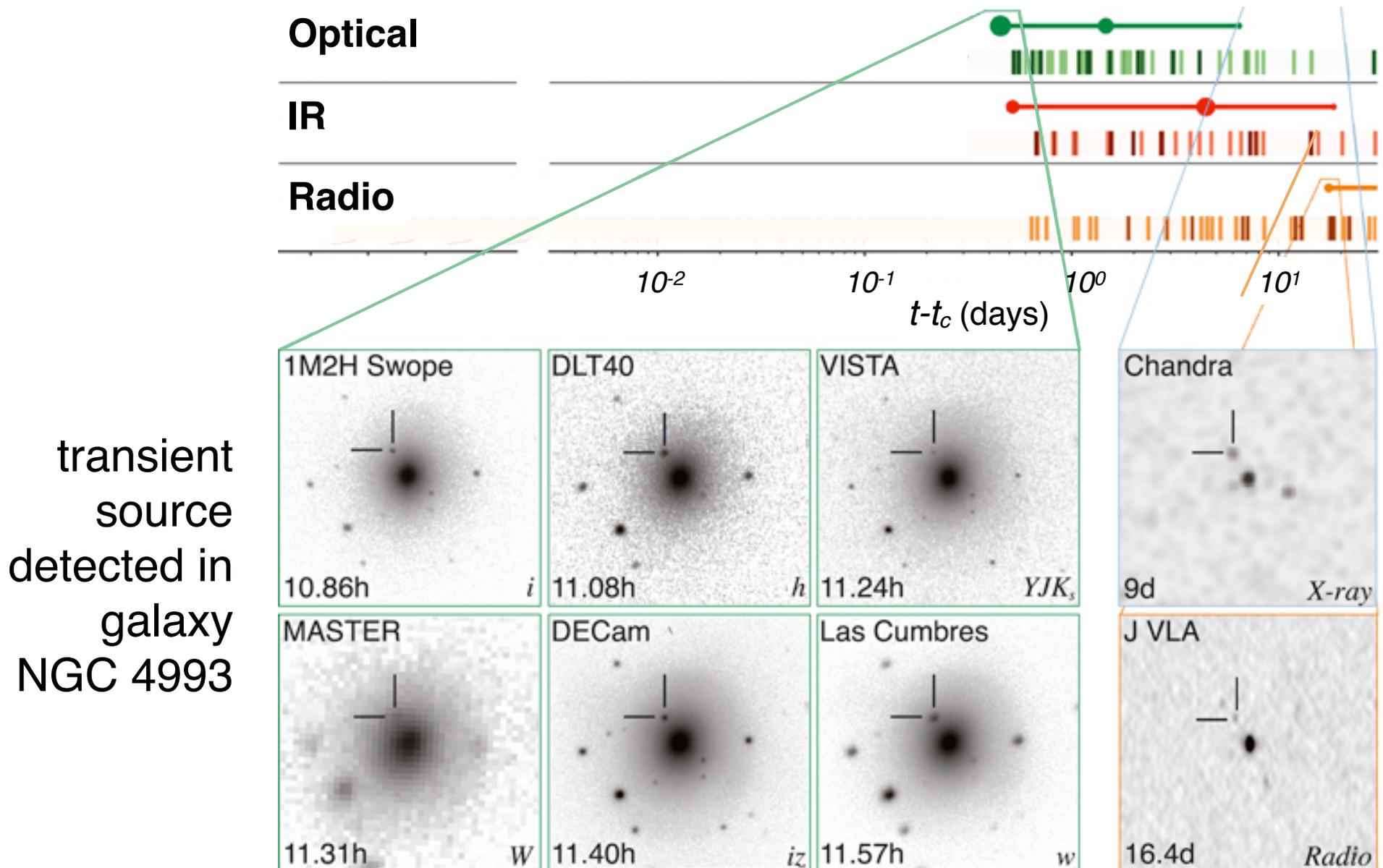
ad. Abbot et al. 2017



Swope +10.9 h

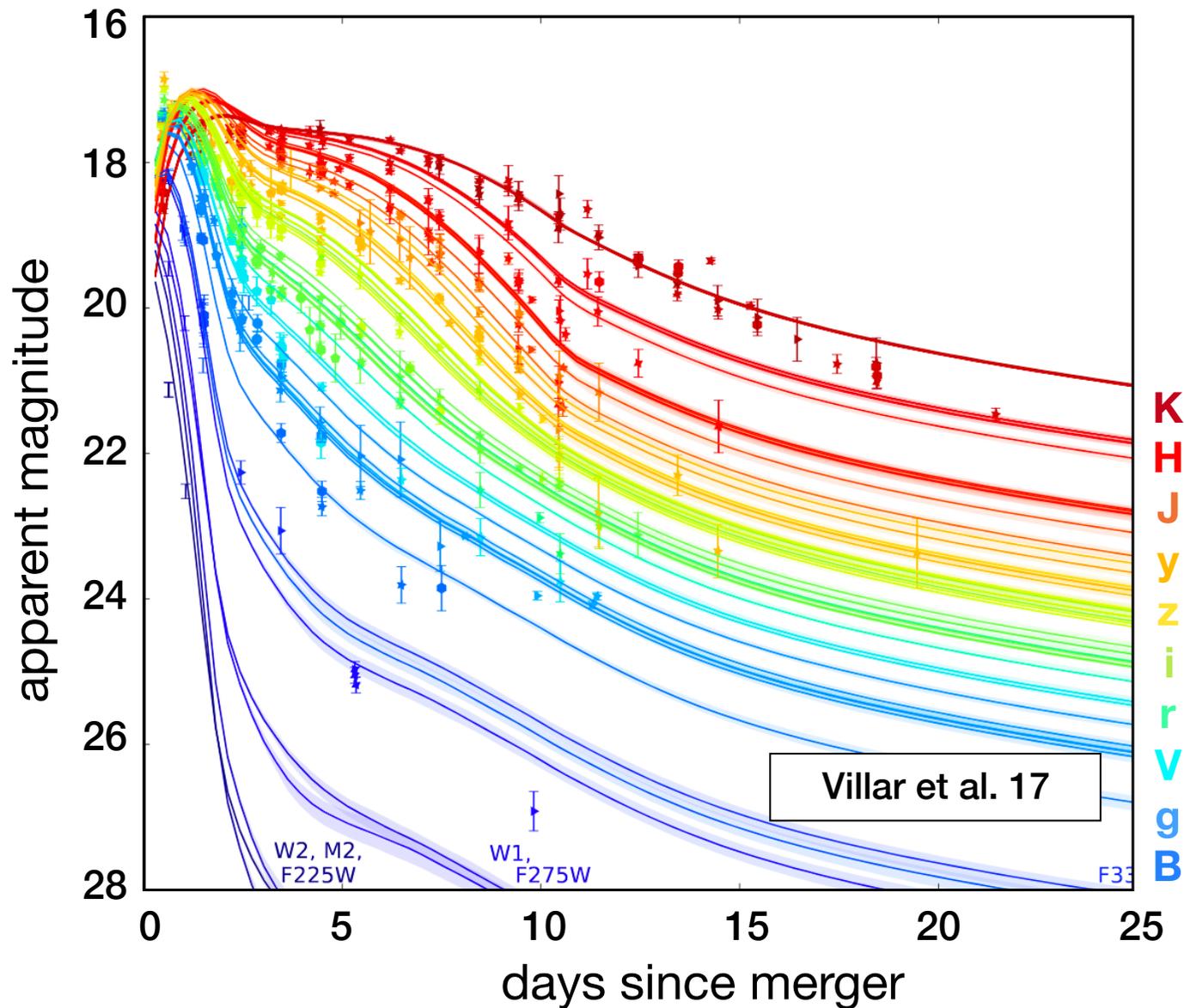


GW170817: the first neutron-star merger

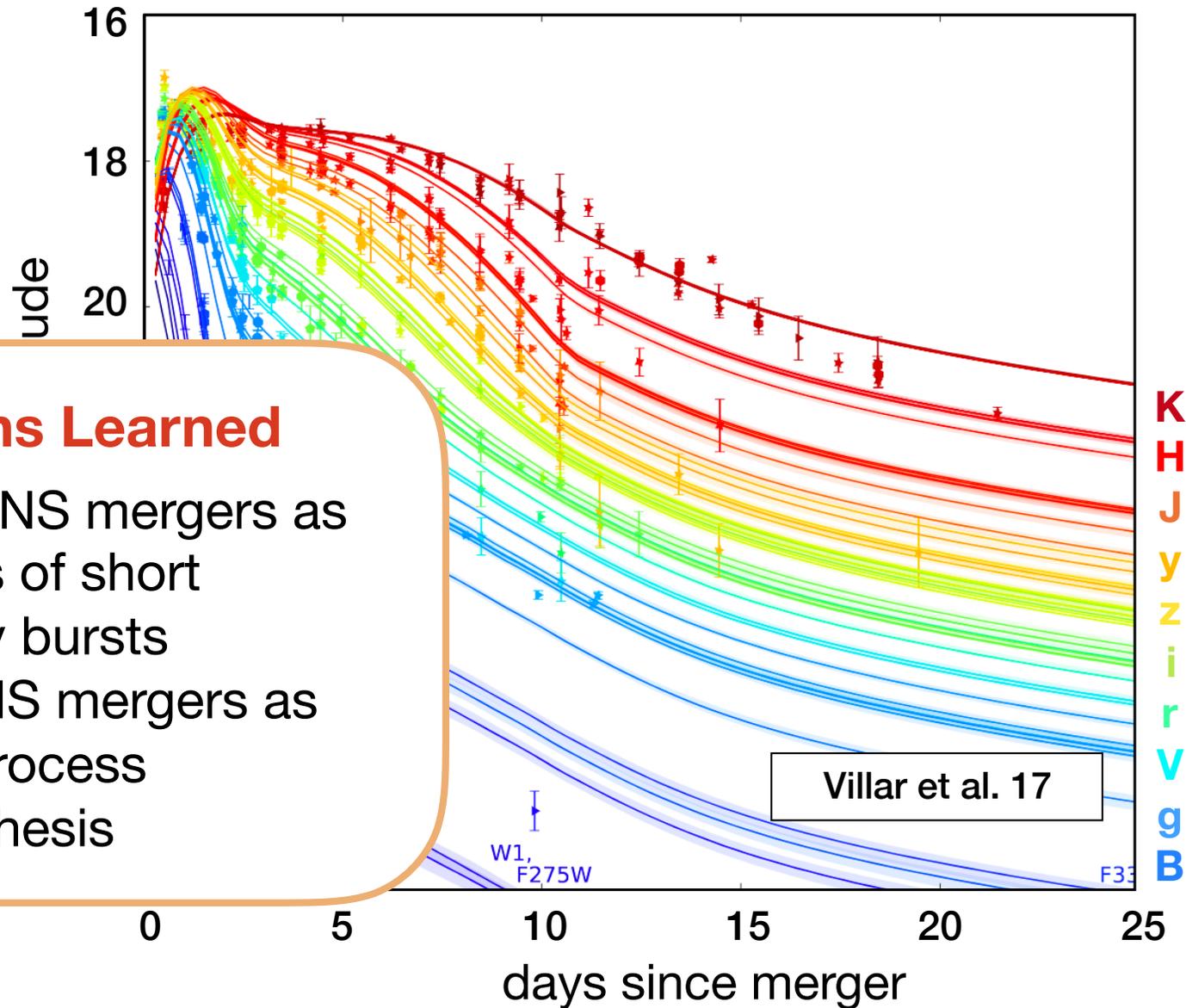


GW170817: the first neutron-star merger

transient
source
detected in
galaxy
NGC 4993



GW170817: the first neutron-star merger

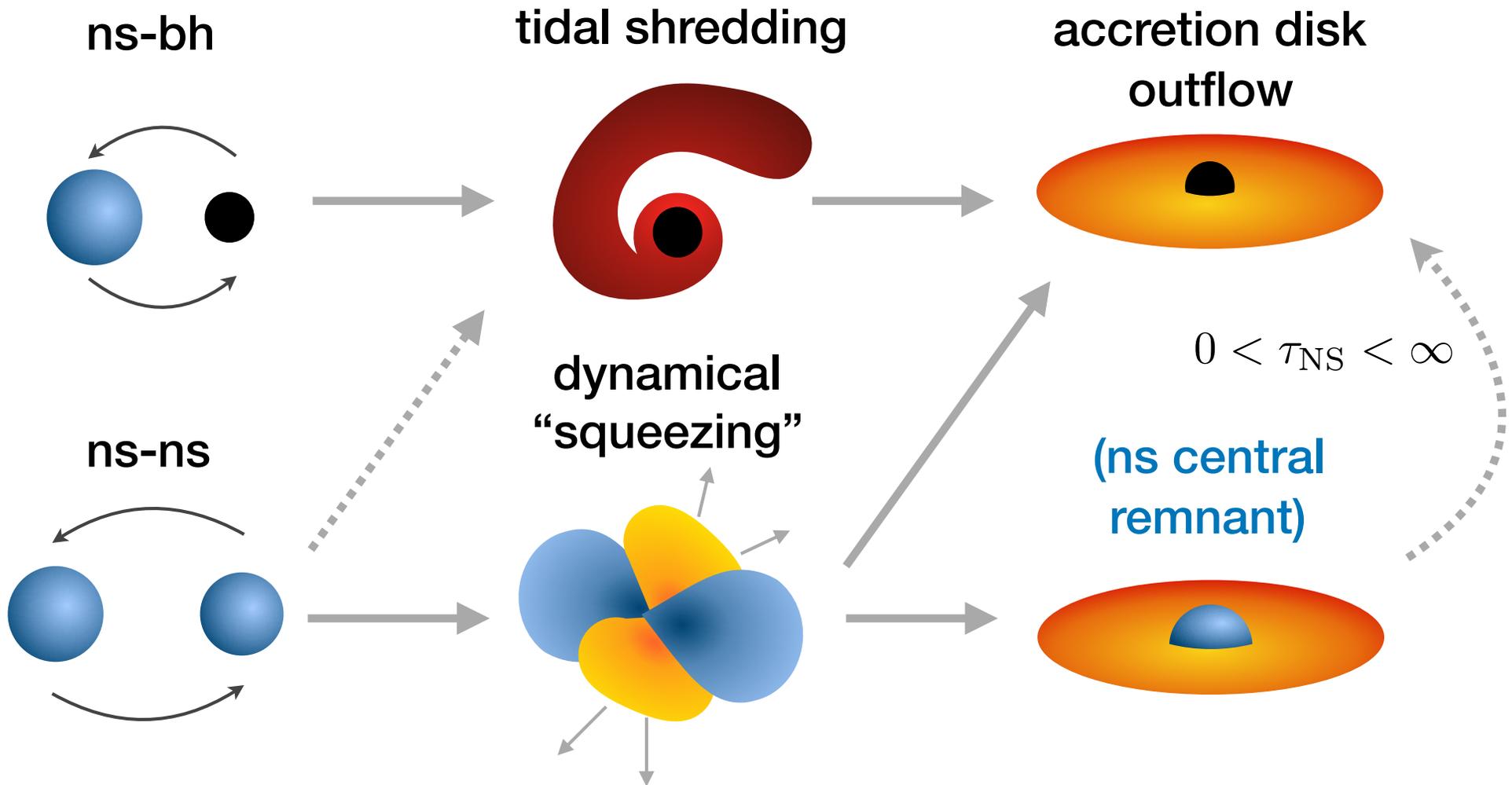


Lessons Learned

- Confirmed NS mergers as progenitors of short gamma-ray bursts
- Identified NS mergers as sites of *r*-process nucleosynthesis

A clearer picture of the ejecta is crucial for progress

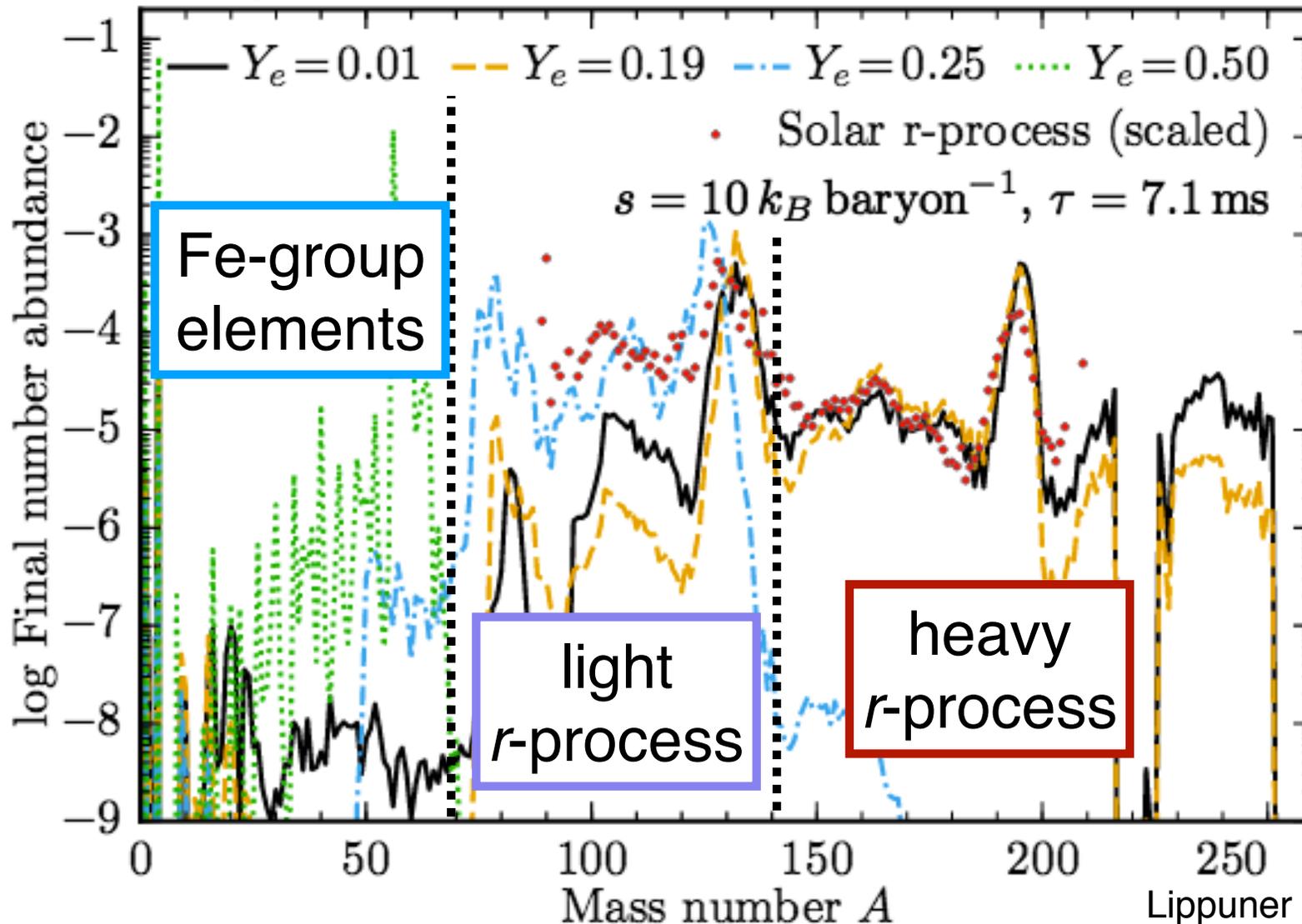
How much mass was ejected, at what velocity, and with what composition?



Interpreting the GW170817 kilonova

Outcomes of the r -process

fewer free n per seed \leftarrow \rightarrow more free n per seed

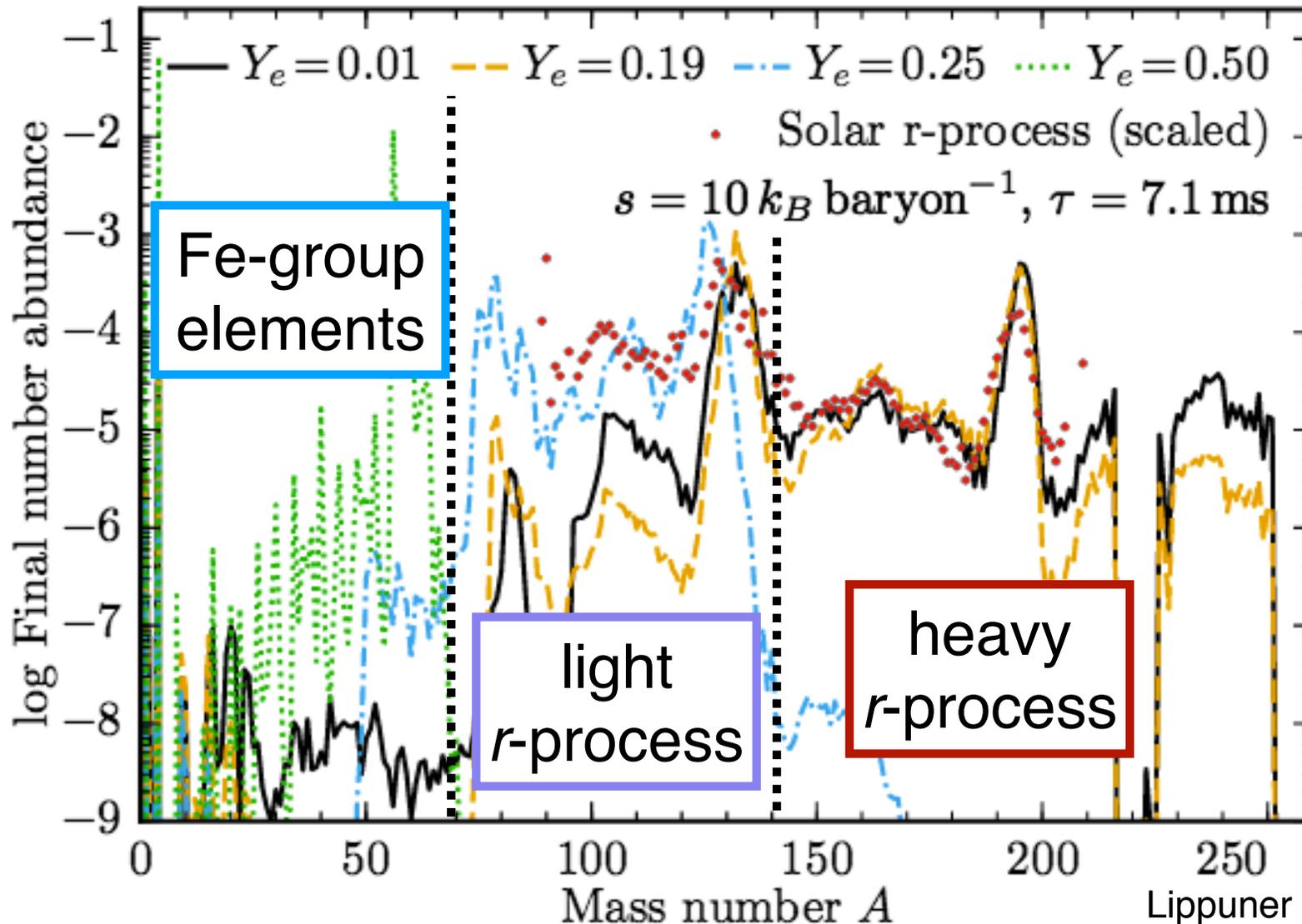


$$Y_e = \frac{p}{p + n}$$

Interpreting the GW170817 kilonova

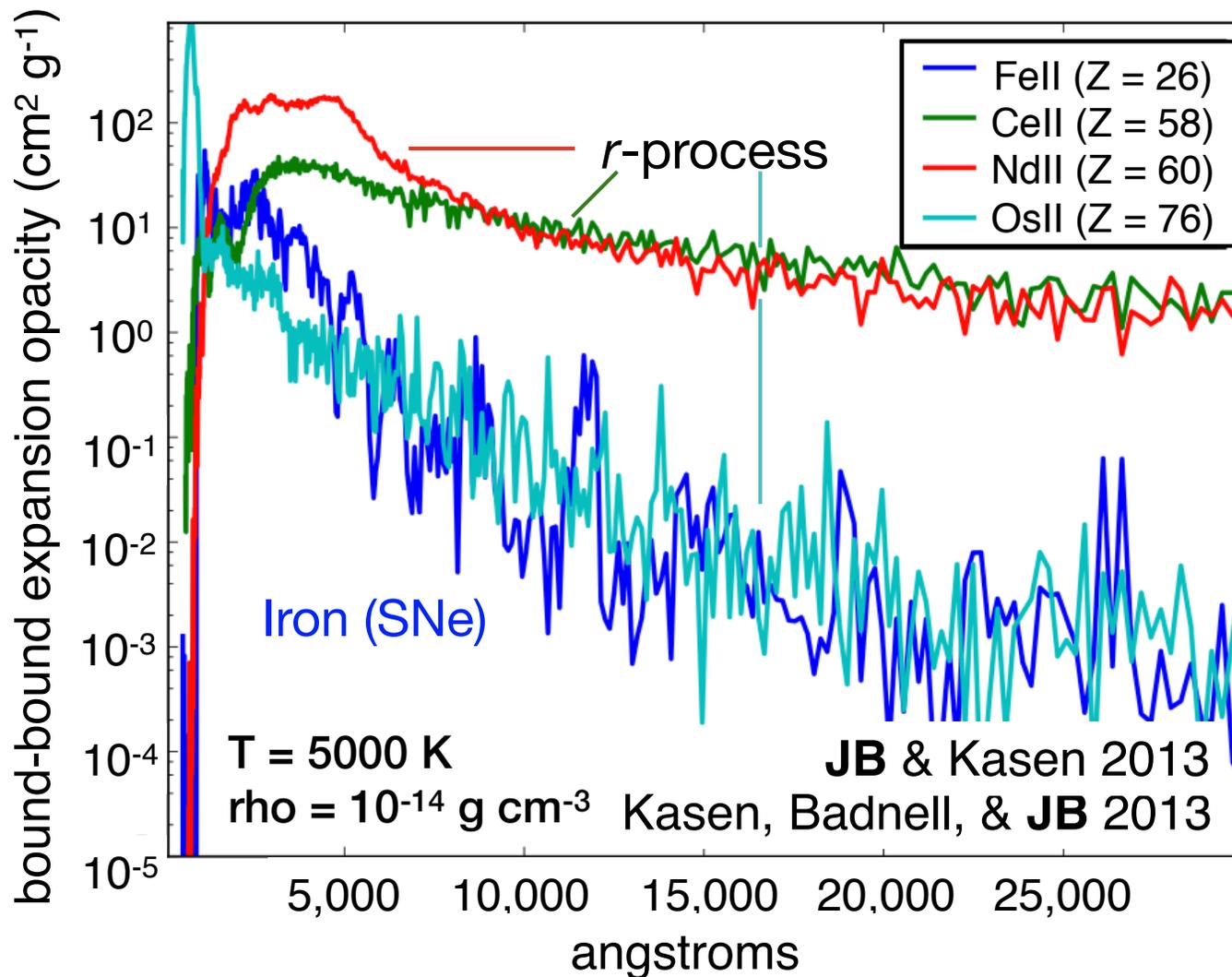
Outcomes of the r -process

fewer weak interactions \longleftrightarrow more weak interactions



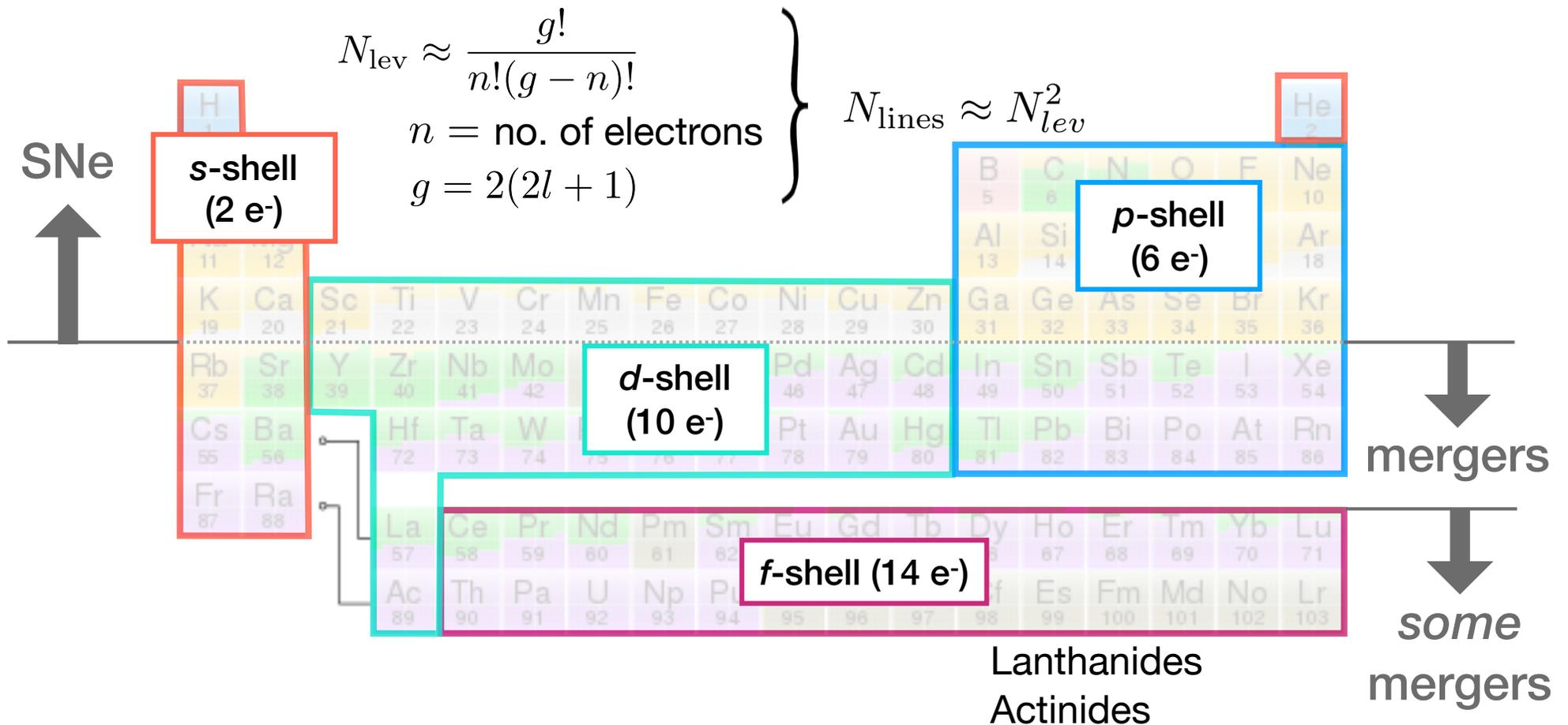
$$Y_e = \frac{p}{p + n}$$

Kilonova opacities depend on composition/nucleosynthesis



Kilonova opacities depend on composition/nucleosynthesis

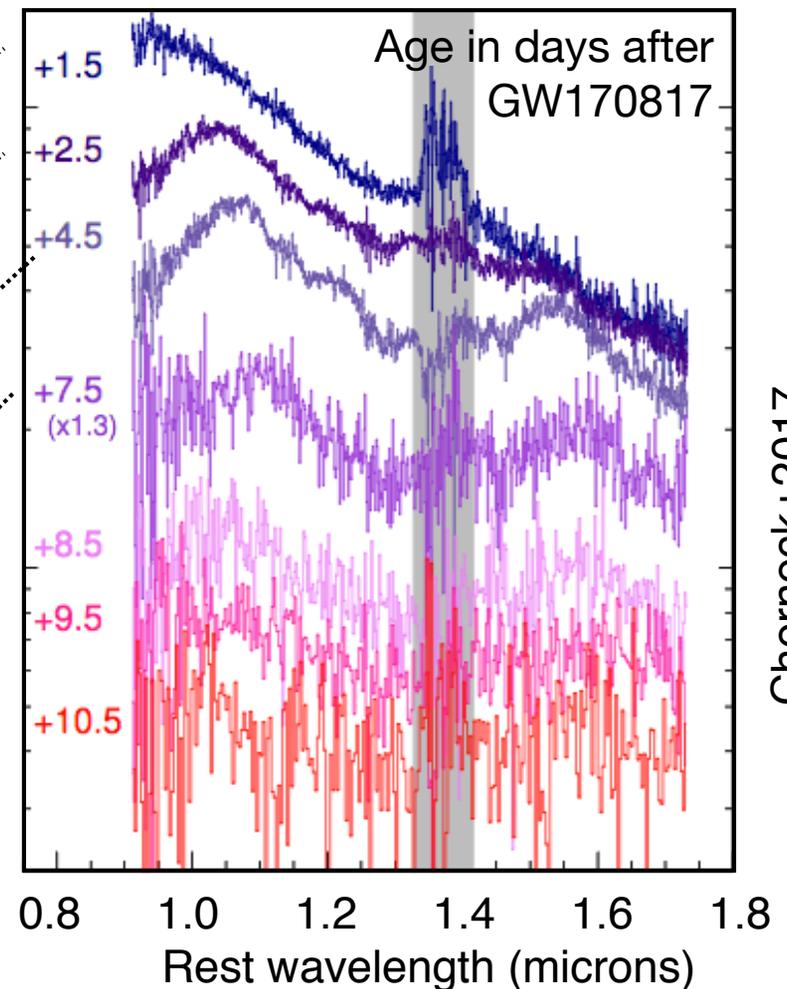
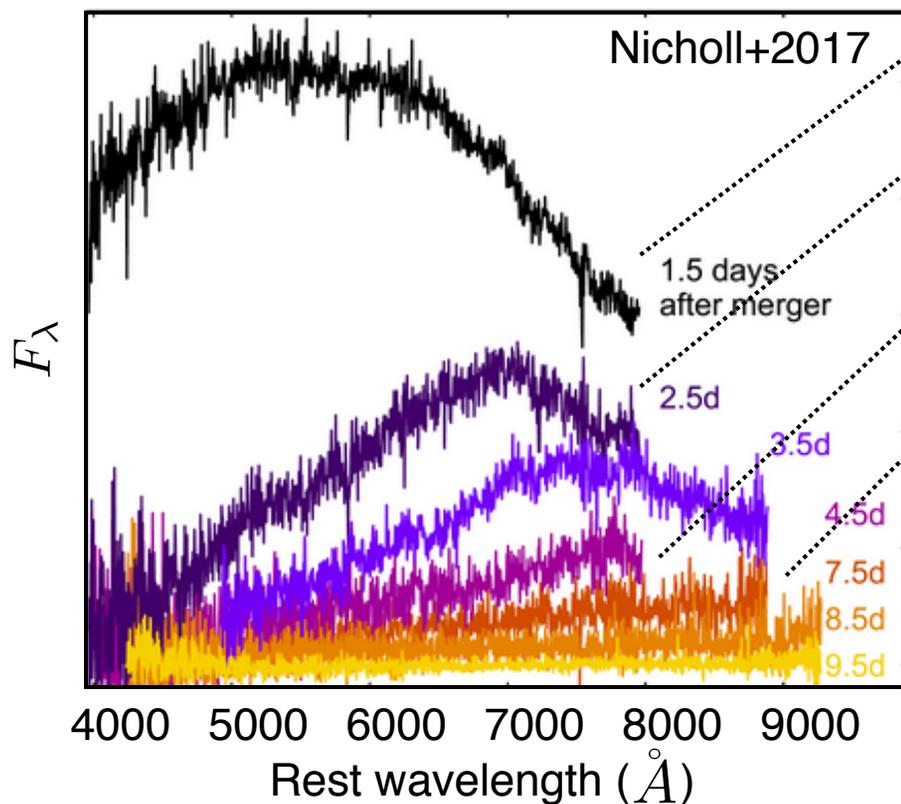
Opacity is correlated with atomic structure



Kilonova opacities depends on composition/nucleosynthesis

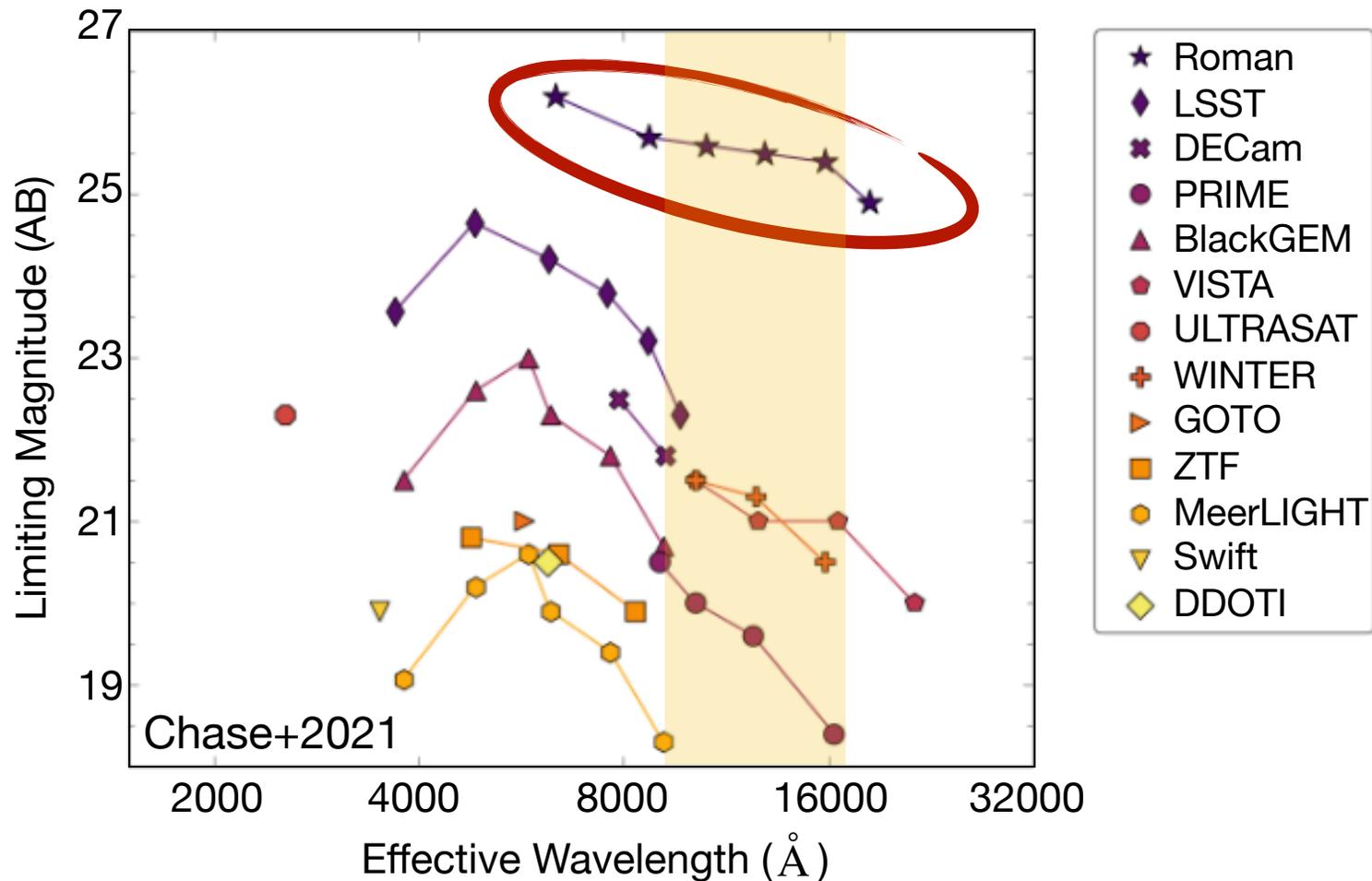
GW170817 showed evidence of multiple outflows with distinct patterns of nucleosynthesis

UV/optical → NIR



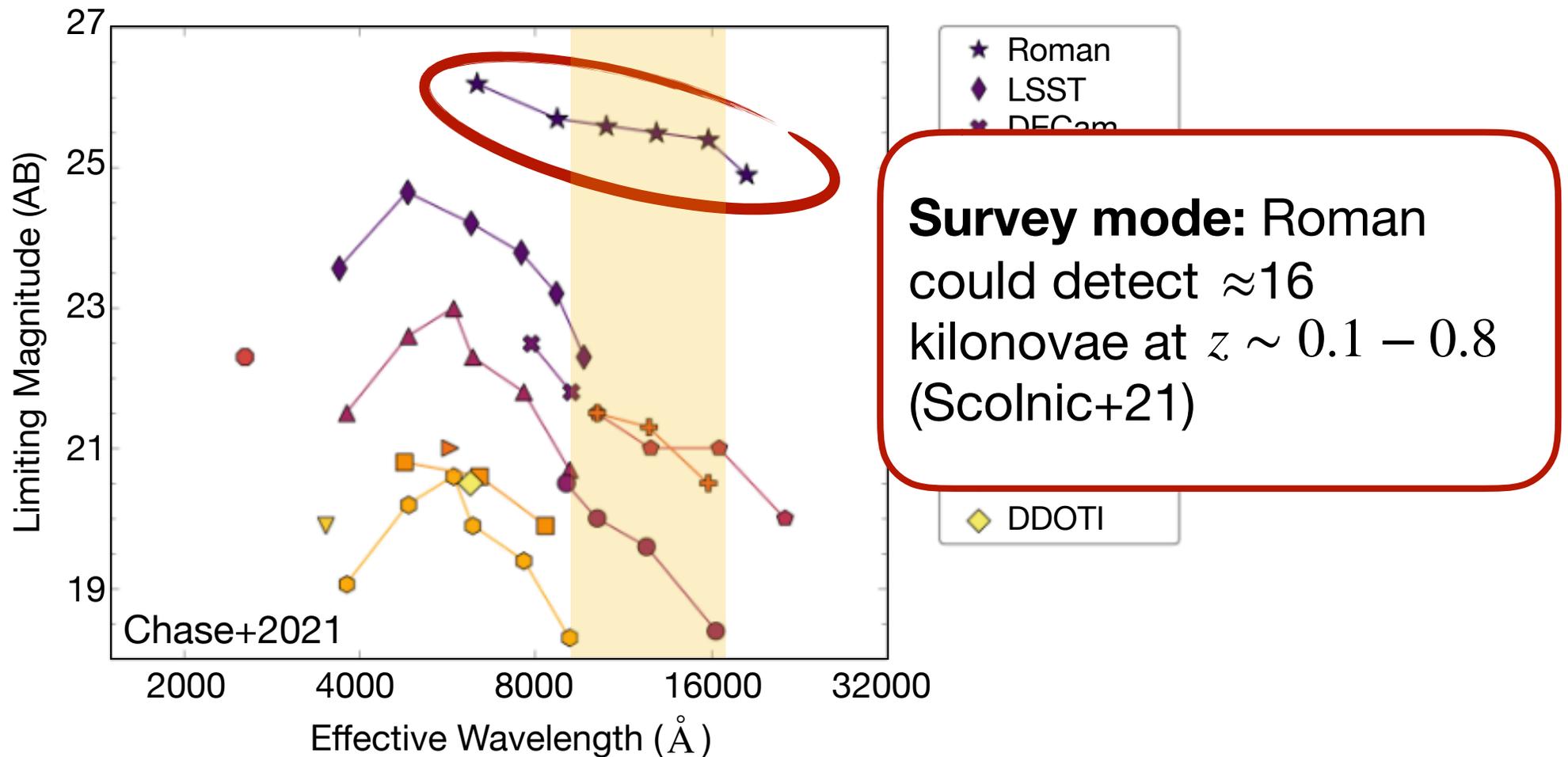
Roman is a powerful tool for kilonova discovery

It is uniquely sensitive in near infrared bands, allowing detections of kilonovae out to high redshifts



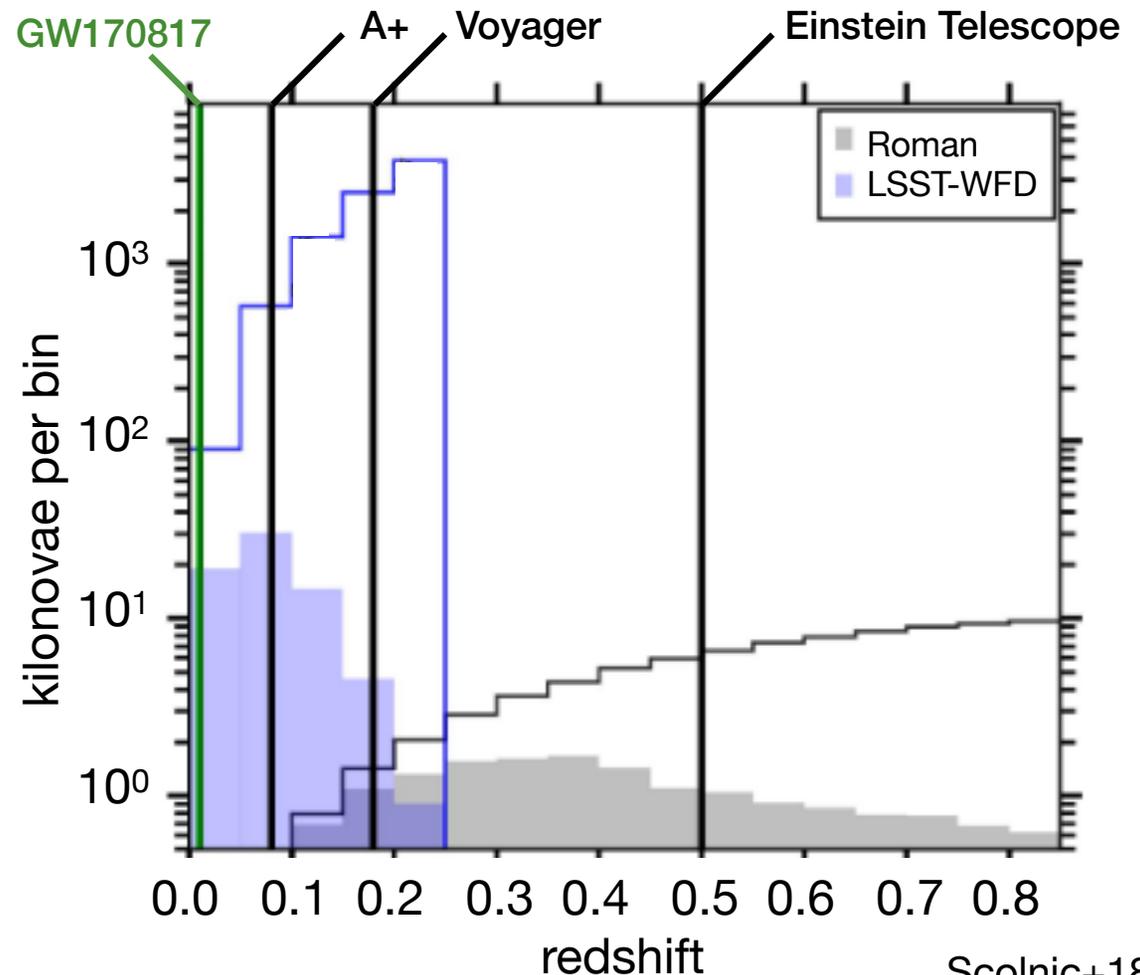
Roman is a powerful tool for kilonova discovery

It is uniquely sensitive in near infrared bands, allowing detections of kilonovae out to high redshifts



Roman is [could be] a powerful tool for gravitational-wave follow-up

- Wide field of view (0.28 deg^2 , 100x Hubble's) can efficiently search large GW sky localizations
- Sensitivity out to red wavelengths is increasingly important as advances on the GW side push the NS merger horizon to higher and higher redshifts



Roman is [could be] a powerful tool for gravitational-wave follow-up

- Wide field of view (0.28 deg^2 , 100x Hubble's) can efficiently search large GW sky localizations
- Sensitivity out to red wavelengths is increasingly important as advances on the GW side push the NS merger horizon to higher and higher redshifts
- Follow-up is not limited to photometry!

