

THE SIMULATED CATALOG
OF OPTICAL TRANSIENTS
AND CORRELATED HOSTS
(SCOTCH)







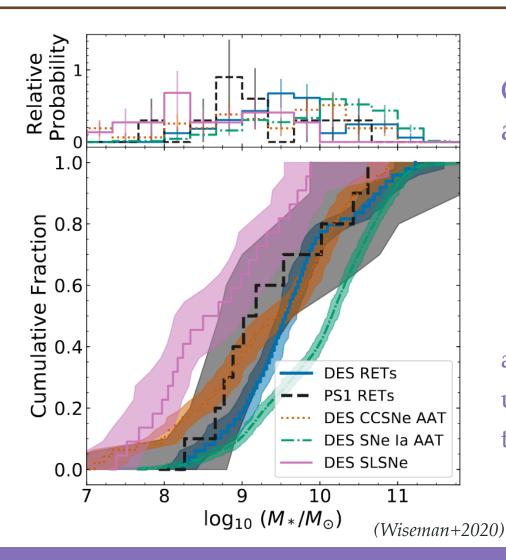
ALEX GAGLIANO
UNIVERSITY OF ILLINOIS/CCA

gaglian2@illinois.edu

MARTINE LOKKEN
UNIVERSITY OF TORONTO
lokken@astro.utoronto.ca

Collaborators: G. Narayan, R. Hložek, R. Kessler, L. Salo, J.F. Crenshaw, M. Vincenzi

### THE VALUE OF HOST-GALAXY INFORMATION IN THE TIME-DOMAIN



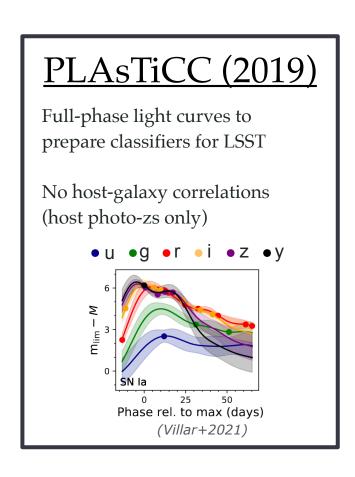
#### Contextual information aids in

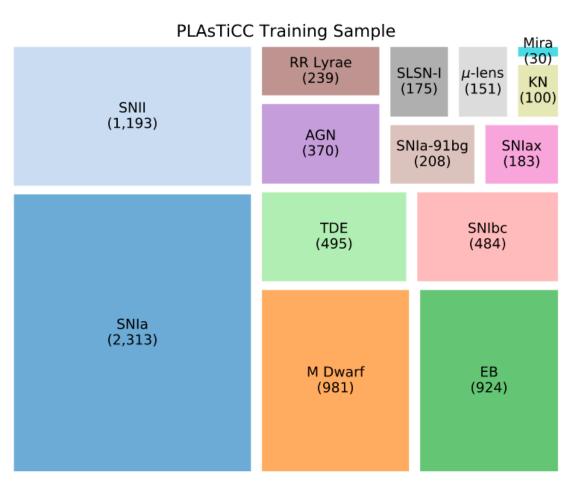
- a. Classification
- b. Progenitor studies
- c. Cosmological constraints
- d. Fast follow-up of relevant events

and will be crucial for understanding Rubin and Roman transients.

Background & Motivation 2

### EXTENDING THE PLASTICC CHALLENGE (DEC. 2018 - FEB. 2019)





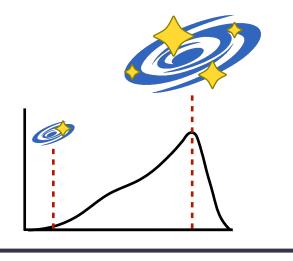
Background & Motivation

## INCREASING THE REALISM OF HIGH-Z TRANSIENT SIMULATIONS

#### **SCOTCH (2022)**

True photometry for 13 extragalactic transient classes (z<3)

Host association dependent on galaxy photometry (griz), color,  $M_*$ , SFR

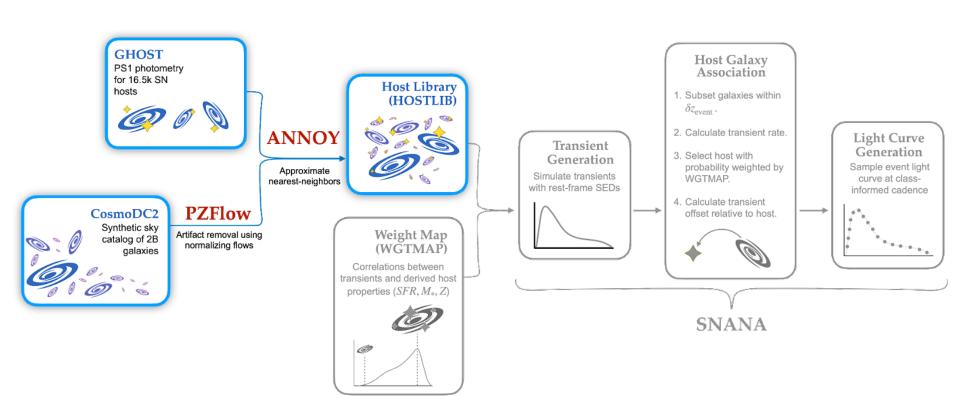


SN Ia (2M)

SN II,IIn (2M)

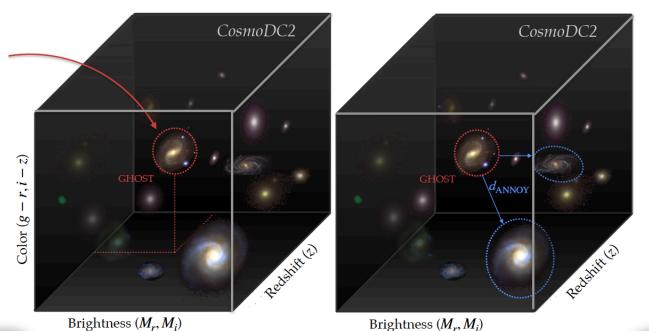
SN IIb	SN Ib	SN Ic	SN II,IIn (100k)	SN Iax
(100k)	(100k)	(100k)		(100k)
SLSN-I	KN	AGN	TDE (100k)	SN 91bg
(100k)	(100k)	(100k)		(100k)

# COMBINING OBSERVED SAMPLES WITH SIMULATIONS



Background & Motivation

#### GENERATING A LIBRARY OF 5M TRANSIENT HOST GALAXIES



PS1 photometry for 16.5k observed SN host galaxies (Gagliano+2021)



(Lokken, **Gagliano**, et al., 2022)

Galaxies drawn from *CosmoDC2* matching photometry of observed *GHOST* galaxies.

2B galaxies in 440 sq. deg simulation (Korytov+2019)

Methodology 6

#### A LIST OF HOST LIBRARIES BY TRANSIENT CLASS

SN Ia Hosts SN II
Hosts
II, IIP, IIL, IIn

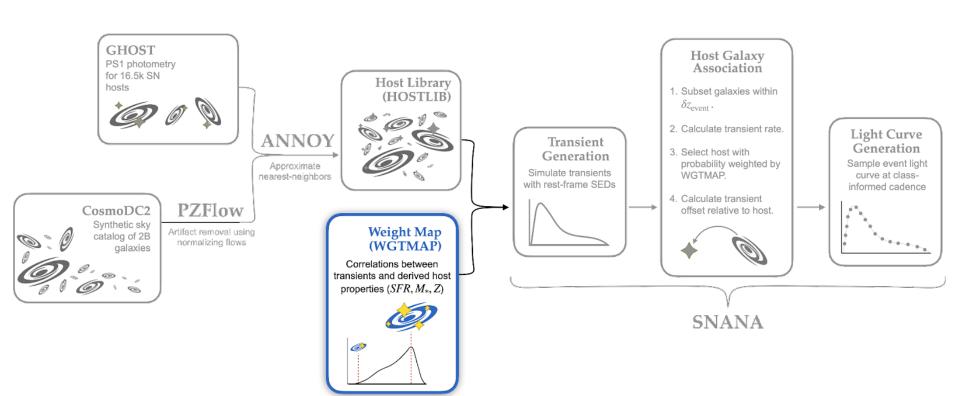
H-poor SN Hosts

Ib, Ic, Ic-BL, SLSN-I, IIb Random DC2 Subset

AGN, TDE, KN

Methodology

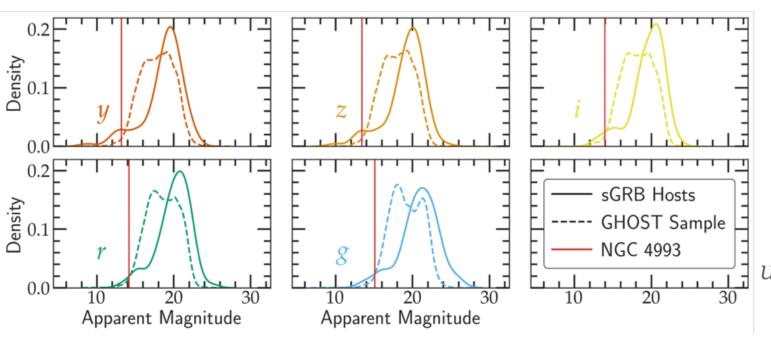
## CORRELATING TRANSIENTS WITH DERIVED HOST-GALAXY FEATURES



Methodology 8

### LEVERAGING SGRB SAMPLES TO PREPARE FOR THE NEXT KILONOVA EVENT

#### *Using sGRBs as a proxy for KNe-host galaxy correlations:*

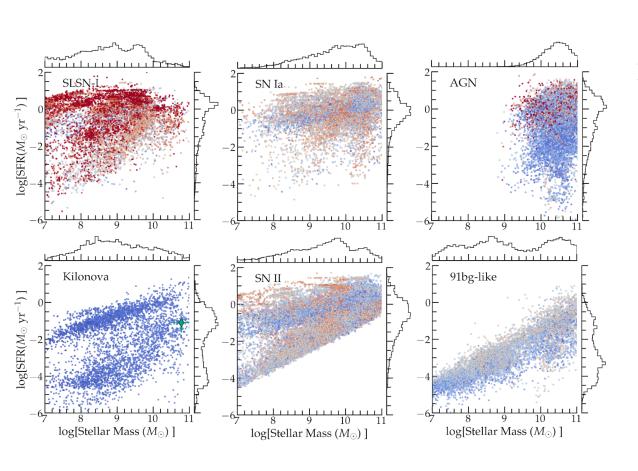




Laura Salo
PhD Candidate,
University of Minnesota

Results & Validation 9

## VALIDATING SYNTHETIC HOST-GALAXY CORRELATIONS: $M_*$ AND SFR

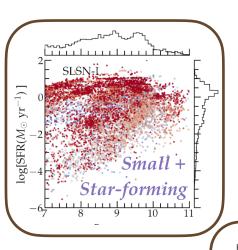


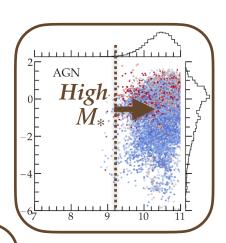
Host selection from host library weighted by class-specific weight map

Weight maps encode **derived** host-galaxy correlations (*M*\*, *SFR*)

Results & Validation 10

### VALIDATING SYNTHETIC HOST-GALAXY CORRELATIONS: $M_*$ AND SFR





Host selection from host library weighted by class-specific weight map

Weight maps encode **derived** host-galaxy correlations (*M*\*, *SFR*)

SLSNe-I found in low-mass, blue galaxies\*, SNe II (core-collapse) in active galaxies\*\*, and AGN in massive galaxies\*\*\*.

Results & Validation 11

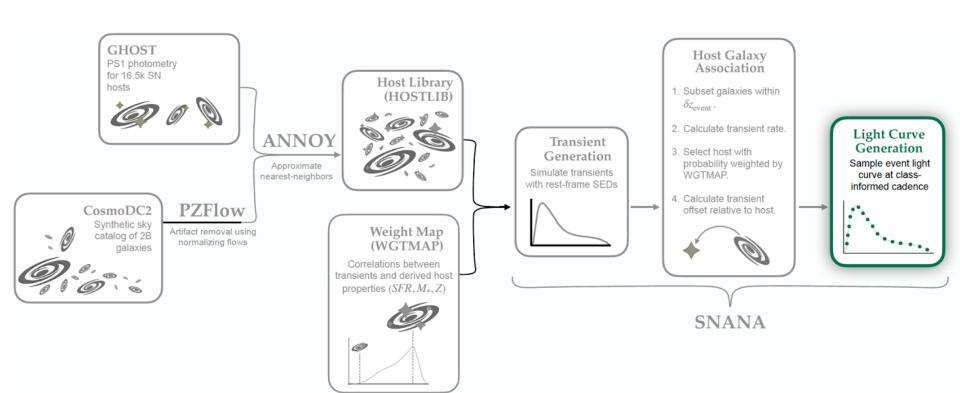
 $\log[\text{Stellar Mass}(M_{\odot})]$ 

<sup>\*</sup>Perley+2016, Wiseman+2020

<sup>\*\*</sup>*Kelly*+2012

<sup>\*\*\*</sup>Kauffmann+2003

#### LIGHT CURVE GENERATION: AN OVERVIEW



Results & Validation 12

### PROPERTIES OF IDEALIZED TRANSIENT LIGHT CURVES

#### Transient photometry is

#### High-cadence:

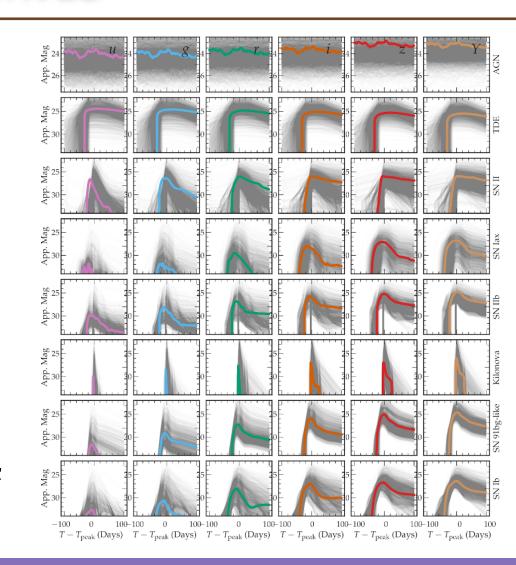
Regular 2-day for most classes Variable for rapidly-evolving transients (KN)

#### *Top-of-the-galaxy:*

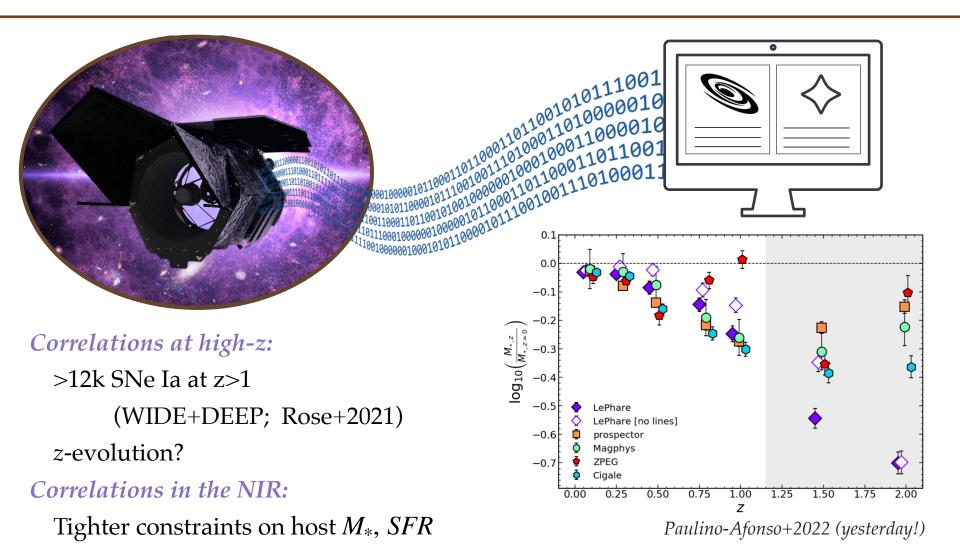
No atmosphere No Galactic extinction

#### Host-extincted

Transients also placed at realistic offsets from their host galaxies.



## EXPLORING THE TRANSIENT (+HOST) UNIVERSE WITH NANCY GRACE ROMAN



### CONCLUSION: SCOTCH FOR UPCOMING TIME-DOMAIN SURVEYS

Catalog of 5M optical transients of 13 extragalactic classes (z<3) with realistic host-galaxy properties. Paper in collaboration-wide review!

MNRAS 000, 1-21 (2021)

Preprint 2 February 2022

Compiled using MNRAS LATEX style file v3.0

#### The Simulated Catalog of Optical Transients and Correlated Hosts (SCOTCH)

Martine Lokken<sup>1,2,3</sup>\*, Alex Gagliano<sup>4,5,6</sup>, Gautham Narayan<sup>4,5</sup>, Renée Hložek<sup>1,3</sup>, Rick Kessler<sup>7,8</sup>, John Franklin Crenshaw<sup>9</sup>, Laura Salo<sup>10</sup>, Catarina Alves<sup>11</sup>, Deep Chatterjee<sup>4,5</sup>, Maria Vincenzi<sup>12</sup>

<sup>1</sup>David A. Dunlap Department of Astronomy and Astrophysics, University of Toronto, 50 St. George Street, Toronto, Ontario, M5S 3H4 Canada

<sup>2</sup>Canadian Institute for Theoretical Astrophysics, University of Toronto, 60 St. George St., Toronto, ON M5S 3H4, Canada

<sup>3</sup>Dunlap Institute of Astronomy & Astrophysics, 50 St. George St., Toronto, ON M5S 3H4, Canada

<sup>4</sup>Department of Astronomy, University of Illinois at Urbana-Champaign, 1002 W. Green St., IL 61801, USA

<sup>5</sup> Center for Astrophysical Surveys, National Center for Supercomputing Applications, Urbana, IL, 61801, USA
<sup>6</sup> National Science Foundation Graduate Research Fellow

<sup>7</sup>Department of Astronomy and Astrophysics, University of Chicago, Chicago, IL 60637, USA

<sup>8</sup>Kavli Institute for Cosmological Physics, University of Chicago, Chicago, IL 60637, USA

<sup>9</sup>Department of Physics, University of Washington, Box 351560, Seattle, WA 98195

<sup>10</sup> School of Physics and Astronomy, University of Minnesota, 116 Church Street S.E., Minneapolis, MN 55455, USA
<sup>11</sup> Department of Physics & Astronomy, University College London, Gower Street, London WC1E 6BT, UK

Department of Physics, Duke University Durham, NC 27708, USA

Accepted XXX. Received YYY; in original form ZZZ

(Lokken, **Gagliano**, et al., 2022)

- 1. Simulations are **survey-agnostic**\*

  \*except for the LSST ugizy bands
- 2. Host libraries, weight maps, and associated software **open-source**
- 3. Same host-association used for **ELASTICC**, with LSST-specific exposure time, footprint, and cadence.

Conclusions 1