Transient alerts and brokers, a perspective from **F**(**NK**

Exploring the Transient Universe with the Nancy Grace Roman Space Telescope

Anais Möller February 10/11 2022



Acknowledging the traditional owners of the land, Wurundjeri People of the Kulin Nation



VERA C. RUBIN OBSERVATORY LSST a deep "video" of the sky

- world's largest CCD camera: 3.2 * 10⁹ pixels
- 6 optical bands: *ugrizy* (320-1050nm)
- 1,000 images/night = 15 TB/night

5-sigma point source depth

Filter	Single
U	23.9
g	25.0
r	24.7
i	24.0
Z	23.3
У	22.1



























LSST ~ 10 million transient alerts per night



+ connecting with other telescopes and data?



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Rubin community brokers

promising candidates for <X science>

+ connecting with other telescopes and data?



Rubin community brokers



In 2019....

- 14 Letters of Intent
- A lot of work
- Decision on August 2021!

Full stream brokers

- Will receive from Rubin ALL transient candidates within 37 seconds of detection
- Will do this for the 10 years of operations

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+ connecting with other telescopes and data





- A community driven effort, open to anyone
- A community of >40 researchers and engineers >7 countries and growing!
- Open source data + source code



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- Selected by Rubin Operations to receive the full alert stream for 10-years



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- First publication A. Möller, J. Peloton, E.E.O. Ishida et al. MNRAS 2020, arXiv: 2009.10185
- Funding to deploy at CC-IN2P3 for LSST

French computing centre where half of the LSST data will be processed

Fink in a nutshell









LSST ~ 10 million transient alerts per night ~10,000 every 37s



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Since Nov 2019 > 120 million alerts collected & processed

Reduced streams:

- Supernovae
- Microlensing
- Variable stars
- Solar System objects
- GRB afterglows
- Kilonovae
- Early type la supernovae
- And more...



Rubin & Roman

Photometry

Rubin public optical photometry from Difference Imaging to complement Roman NIR data + contextual information + features



Photometric detection @ Rubin



+ shall we trigger follow-up?+ what could it be?



Photometric detection @ Rubin

Spectroscopic follow-up







Spectroscopic resources are limited optimisation needed!

We will not be able to follow-up all candidates in the Rubin era...

Kilonovae candidates



Also see Jielai Zhang's talk on Day 2

Early supernovae candidates

https://fink-portal.org/ZTF21abyfxqr

ZTF21abyfxqr



Sep 19

Observation date

Sep 26



Combining these scores and cross-matches with catalogues we obtain a filtered stream of SN candidates

Early SN Ia Leoni et al. 2021 SN Ia SuperNNova (AM et al. 2020) SN SuperNNova (AM et al. 2020)

Sep 12

2021

Early supernovae candidates

Reported in average 7 days before max

Reported by Fink 11/2020 to 11/2021	847
Spectroscopically confirmed SNe	552 from this <i>446 SNe Ia</i> (>80%)
Spectroscopically confirmed non-SN	1

Combining these scores and cross-matches with catalogues we obtain a filtered steam of SN candidates

Publicly communicated



Other candidates in real-time to trigger follow-up

- Kilonovae
- Svom .
- GRB (orphans, on-axis/off-axis)
 - Neutrino
 - Fast transients
 - Anomalies
 - Supernovae
 - SNe + contextual info
 - PISN
 - ... more to come!

Also Fink candidates are available for fup with integration



Automatically

reported in real-time





Obtaining large samples without spectroscopic follow-up



Obtaining large samples without spectroscopic follow-up: type la supernovae



Select a sample using a probability threshold

+ selection cuts

Obtaining large samples without spectroscopic follow-up: type la supernovae



SNIa vs non-la SNe classification in the Dark Energy Survey using light-curves & host-galaxy spectroscopic redshifts with SuperNNova

Accuracies >98% (Vincenzi+ 2021, AM+2022)

Contamination ~2% (Vincenzi+ 2021)

Also see Helen Qu's talk on Day 1

Obtaining large samples without spectroscopic follow-up: type la supernovae



Spectroscopic samples of hundreds can become photometrically identified samples of thousands!

Obtaining large samples without spectroscopic follow-up

Spectroscopic samples of hundreds can become photometrically identified samples of thousands! NIR data from Roman for better calibration and information ...

Going beyond our current samples with a rich dataset!



Kilonovae J. Vlieghe, GRANDMA

- Microlensing E. Bachelet, M. Moniez •
- GRB (orphans, on-axis/off-axis, SVOM) D. Turpin, J. Bregeon, O. Godet, G. Ducoin, R. Le Montagner, L. ٠ Bouchet, M. Llamas
- **SSO** S. Karpov, B. Carry, R. Le Montagner,
- Neutrino KM3NET: D. Dornic, G. Vannoye, V. Kulikoskiy •
- Supernovae M. Leoni, T. Allam, U. Burhanudin, J. Maund •
- **PISN** Blondin, Pruzhinskaya ٠
- Fast transients Biswas
- Anomalies Pruzhinskaya, Kornilov, Russeil, Beschastnov

And many others: Arnault, Hrivnac, Pateyron, Boutigny, Hernandez, Gangler, Russeil, Nebot, Pineau





Rubin community brokers allow harnessing the power of optical time-domain data

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- Built on state-of-the-art technologies
- Is already processing ZTF data stream (MoU 2020).
- First science modules deployed: SNe, GRB, microlensing, ...



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Transient science can only gain with synergies between Roman's, Rubin's and other multi-wavelength and messenger!

- Complementary photometry
- Multi-wavelength/messenger optical counterparts
- Catalogue information for transients
- Early classification for follow-up coordination & analysis
- Large photometrically selected samples
- Automated reporting + real-time streams for follow-up

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