

The background image shows a night sky filled with numerous white, streaking star trails, indicating long exposure photography. Below the sky, a coastal town is visible along a shoreline, with lights reflecting off the water. In the distance, a large, dark mountain range is silhouetted against the lighter sky.

International Collaboration in Astronomy for Asia

Strategy of Subaru Telescope

Nobuo ARIMOTO
Subaru Telescope

© H.Fujiwara (2015)

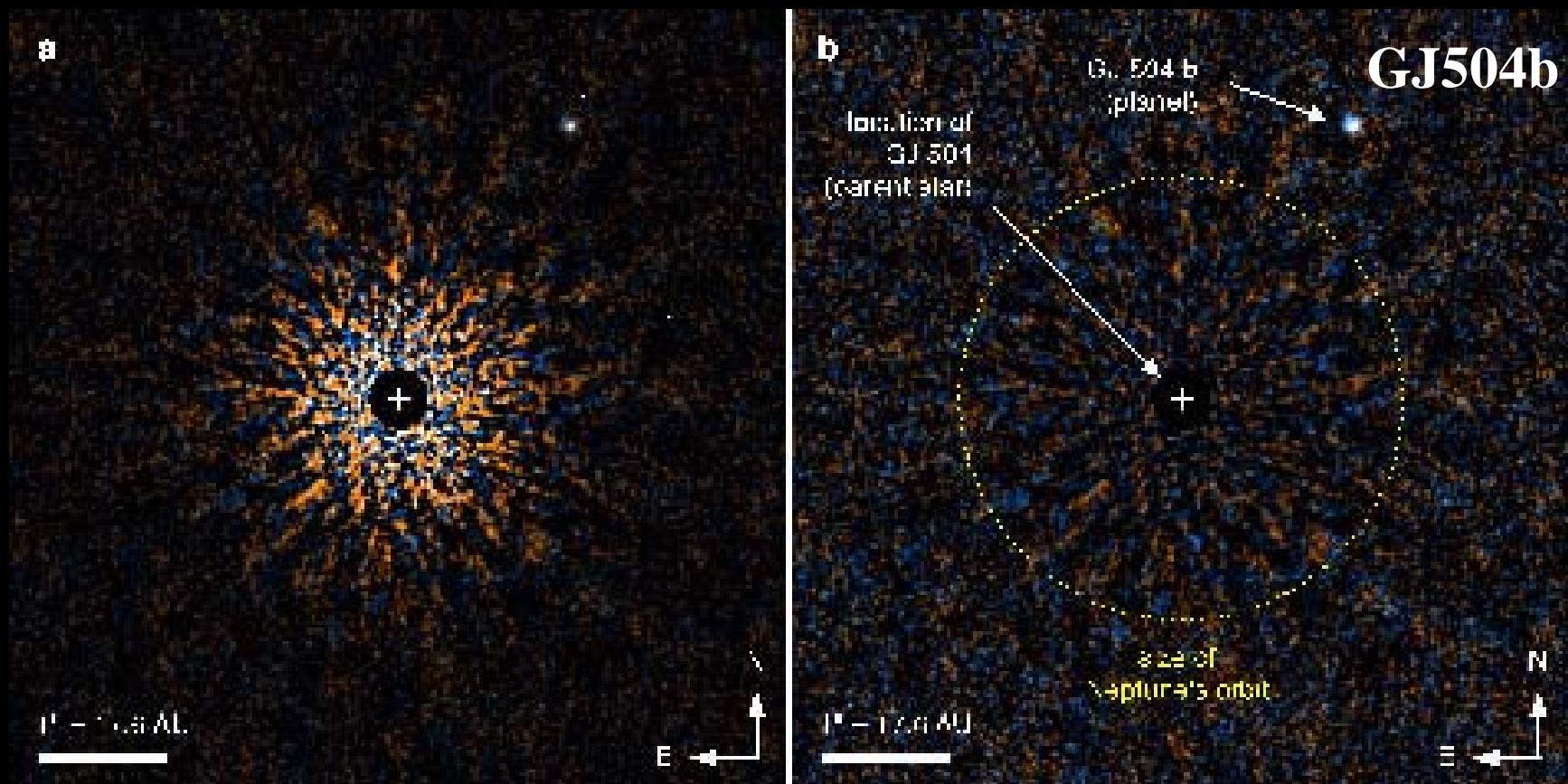


Subaru Telescope Science Topics

Subaru Telescope

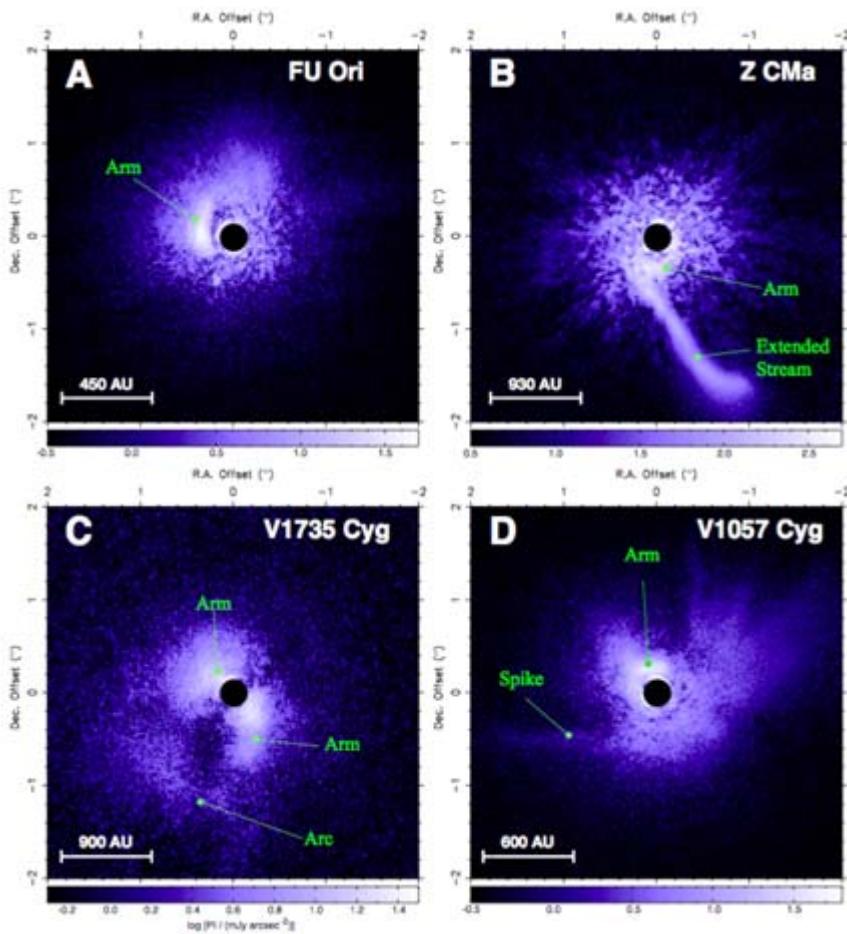
Direct Discovery of a Second Jupiter Shows the Power and Significance of the SEEDS Project

(Kuzuhara et al. 2013)

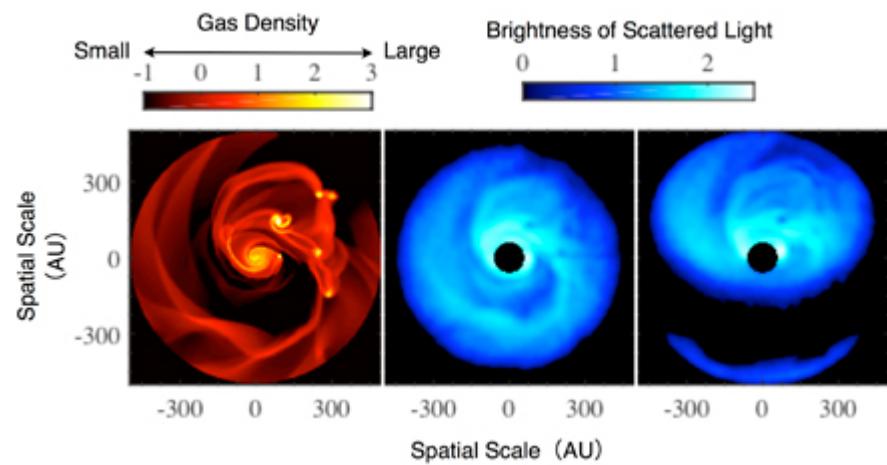


Subaru Telescope

Subaru-HiCIAO Spots Young Stars Surreptitiously Gluttonizing Their Birth Cloud (Liu et al. 2016)



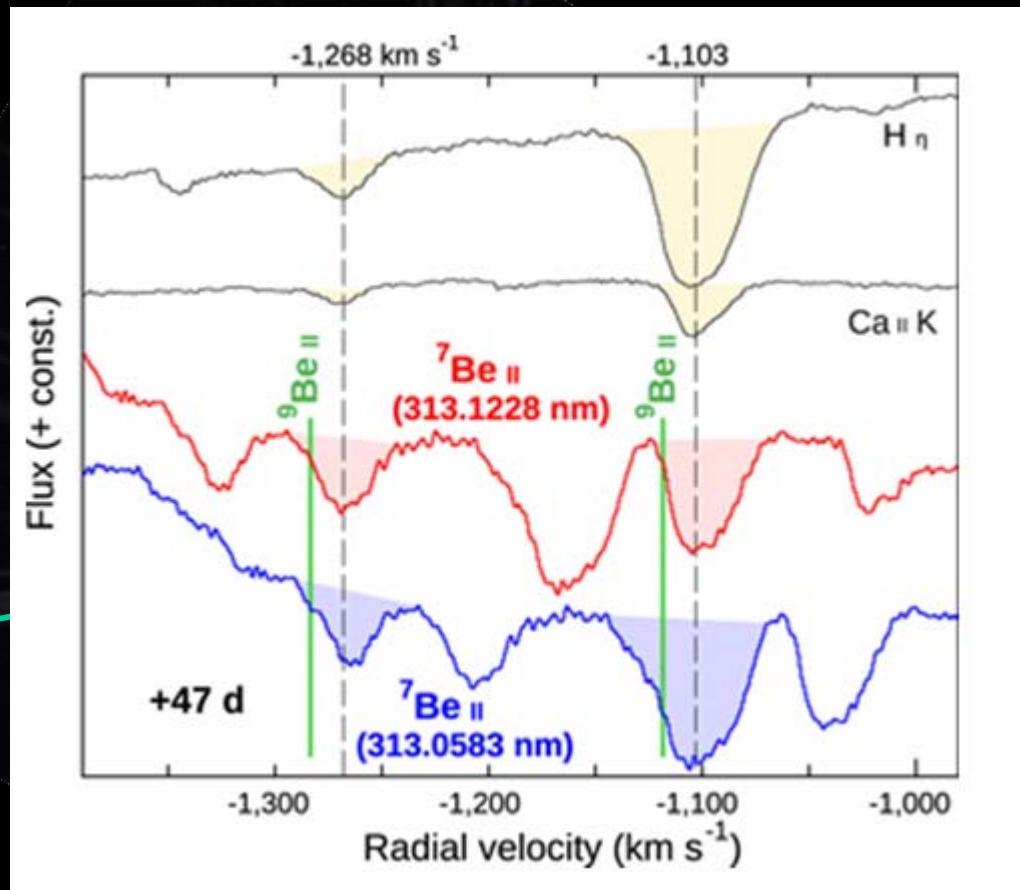
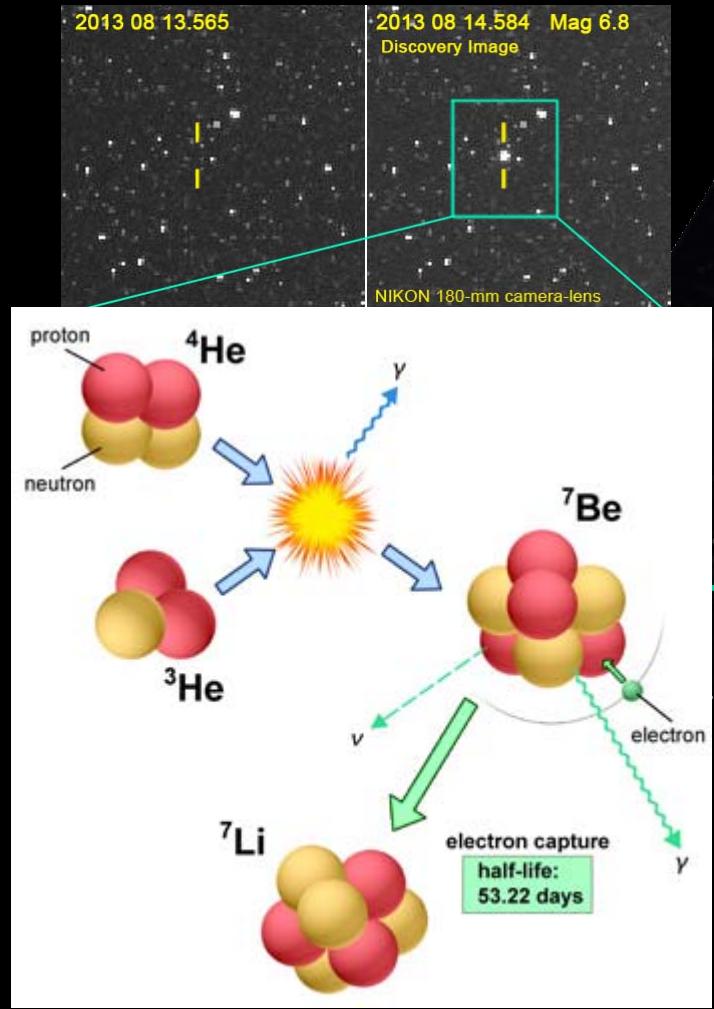
Images made from computer simulations based on one theory for violent growth of a star. (Left) Simulations of the motion of circumstellar materials falling onto a baby star. (Middle and right).



Circumstellar structures of young stars

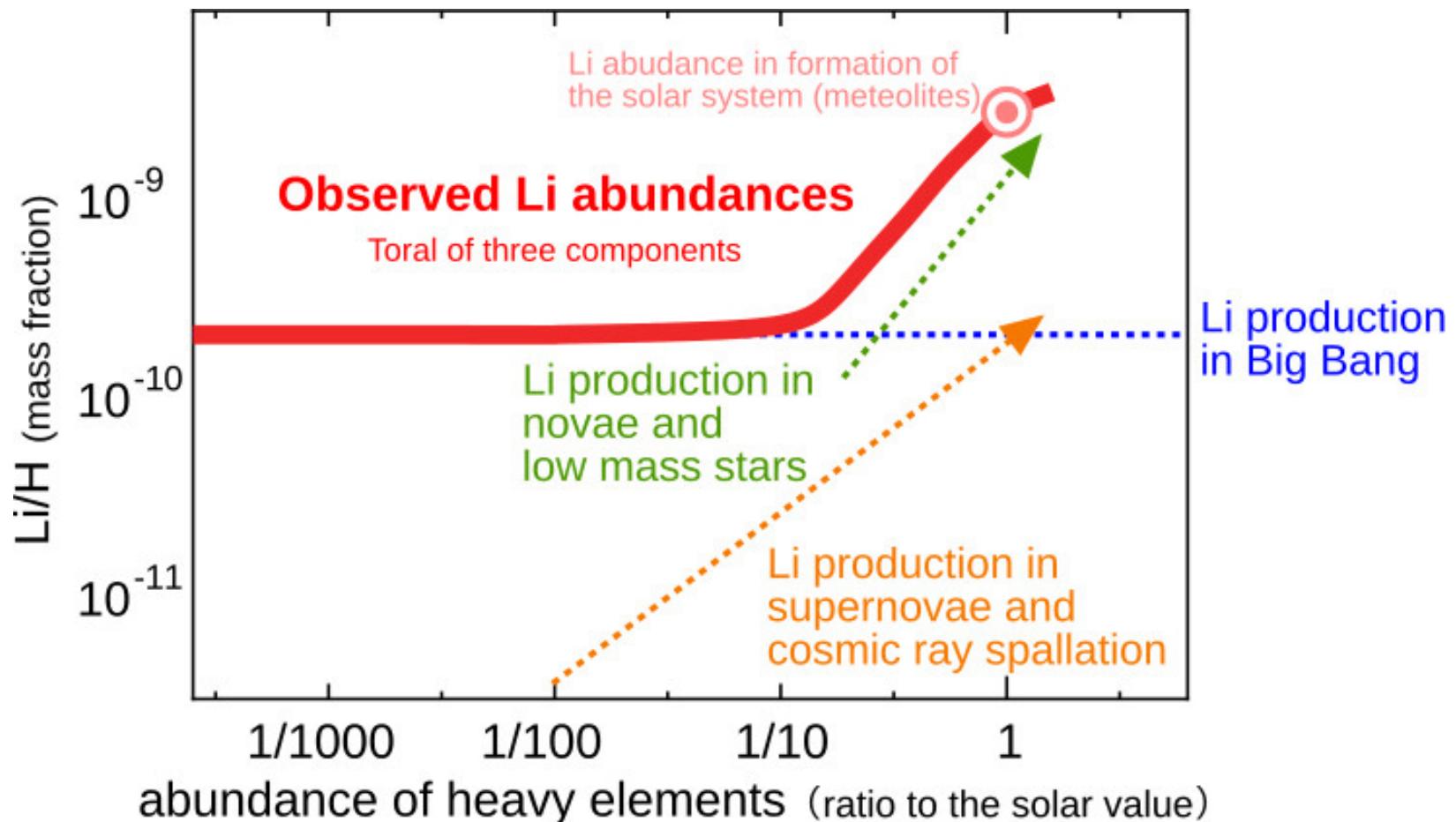
Subaru Telescope

Classical Nova Explosions are Major Lithium Factories in the Universe (Tajitsu et al. 2015)



Classical novae are strong candidates as suppliers of Li in the universe.

Schematic Evolution of Li⁷ in the Universe

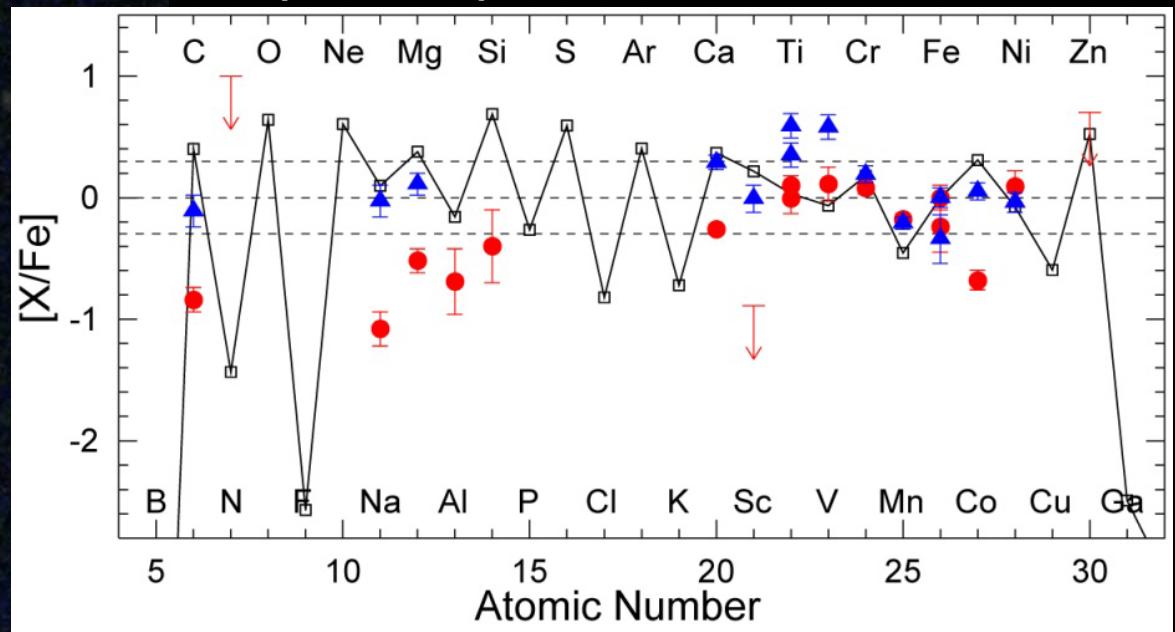


Subaru Telescope

Chemical Signature of First-Generation Very-Massive Stars

The chemical abundance ratios (with respect to iron) of SDSS J0018-0939 (red circles) compared with model predictions for a supernova explosion of a massive star.

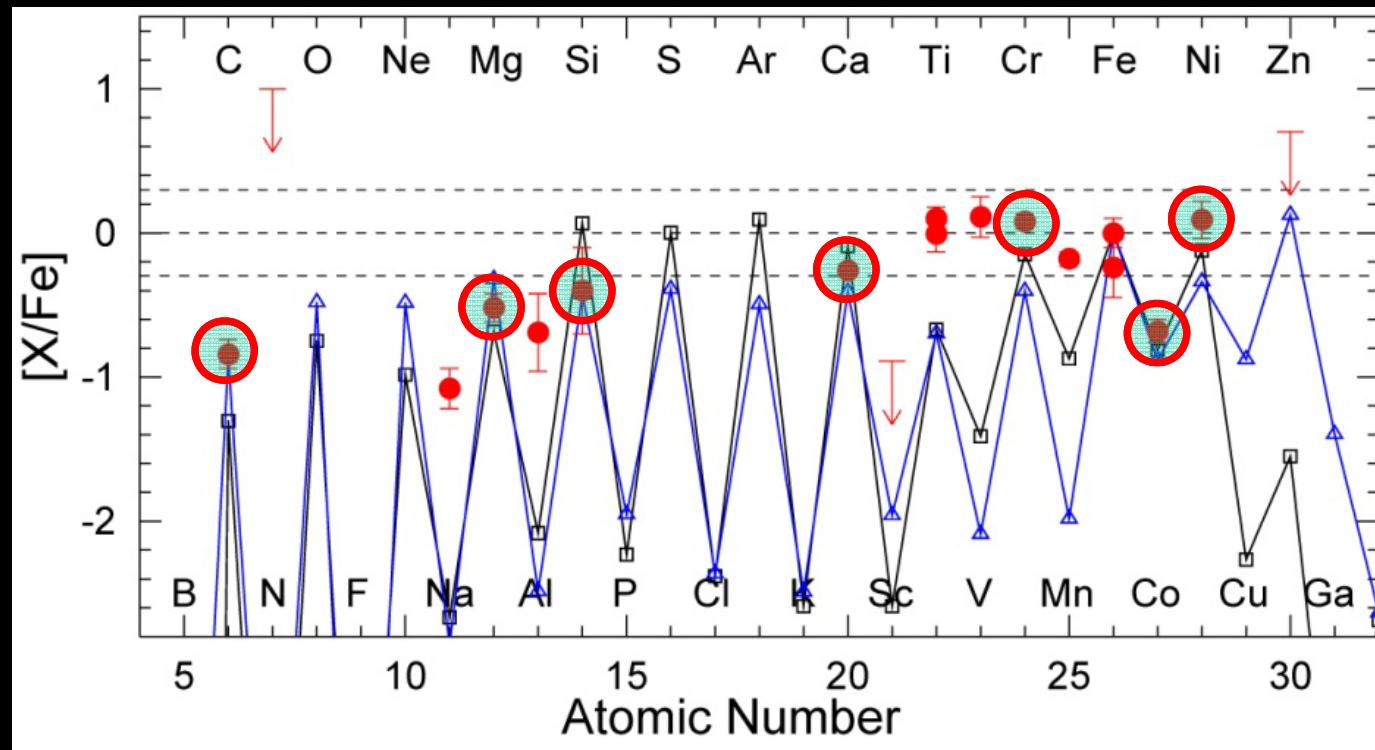
SDSS J0018-0939



Aoki et al. (2014) discover a low-mass star, SDSS J0018-0939, that exhibits the peculiar chemical abundance ratios associated with the process of creating new atomic nuclei (nucleosynthesis) in a first-generation very-massive star. Until now, no observational evidence has supported numerical simulations of the existence of very-massive stars among the first generation of stars formed after the Big Bang.

Subaru Telescope

Chemical Signature of Pair-Instability Supernovae



The chemical abundance ratios (with respect to iron) of SDSS J0018-0939 (red circles) compared with model prediction for explosions of very-massive stars. The black line indicates the model of a pair-instability supernova by a star with 300 solar masses, whereas the blue line shows the model of an explosion caused by a core-collapse of a star with 1000 solar masses.

Subaru Telescope

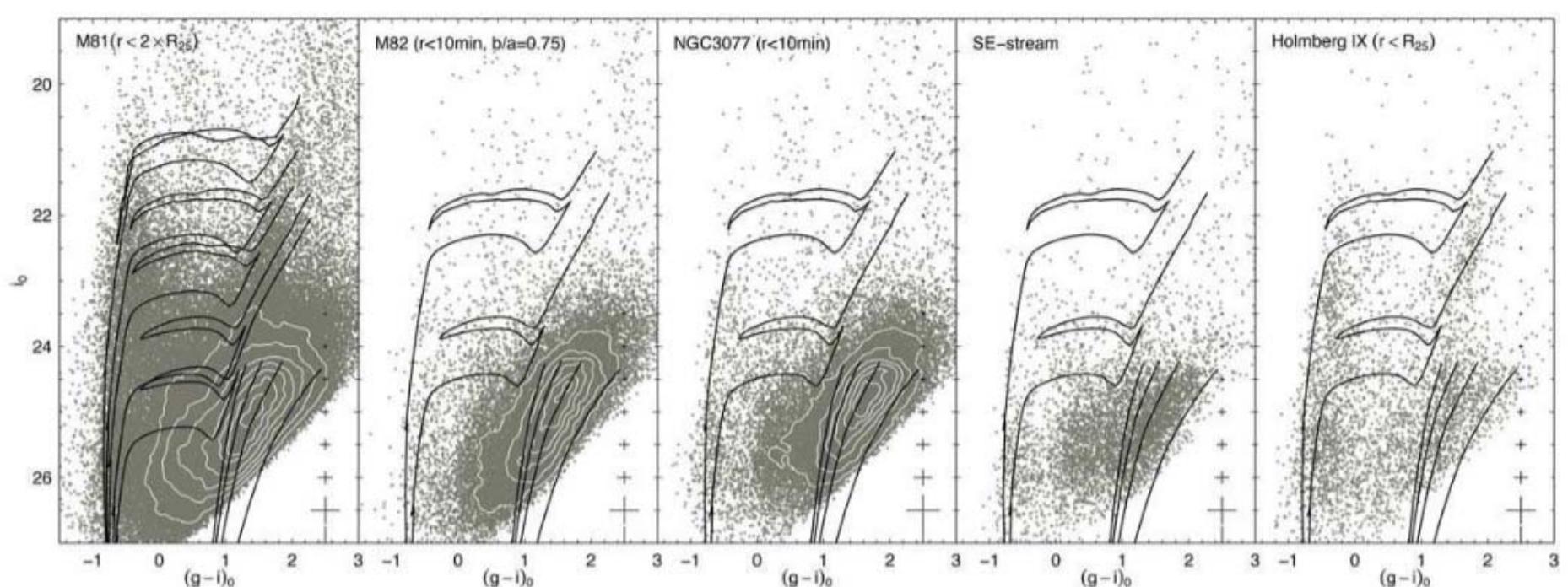
The Ghostly Remnants of Galaxy Interactions Uncovered
in a Nearby Galaxy Group M81

Okamoto et al. (2015)



Subaru Telescope

The Ghostly Remnants of Galaxy Interactions Uncovered in a
Nearby Galaxy Group (Okamoto et al. 2015)



M81

M82

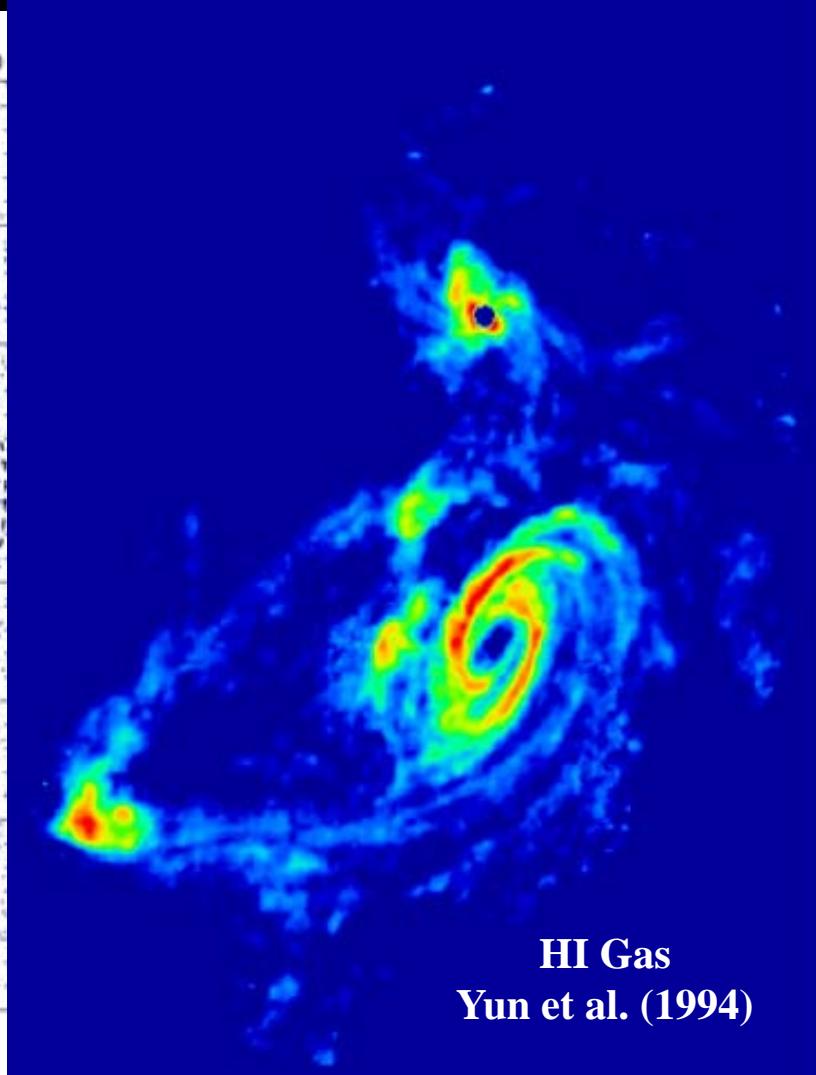
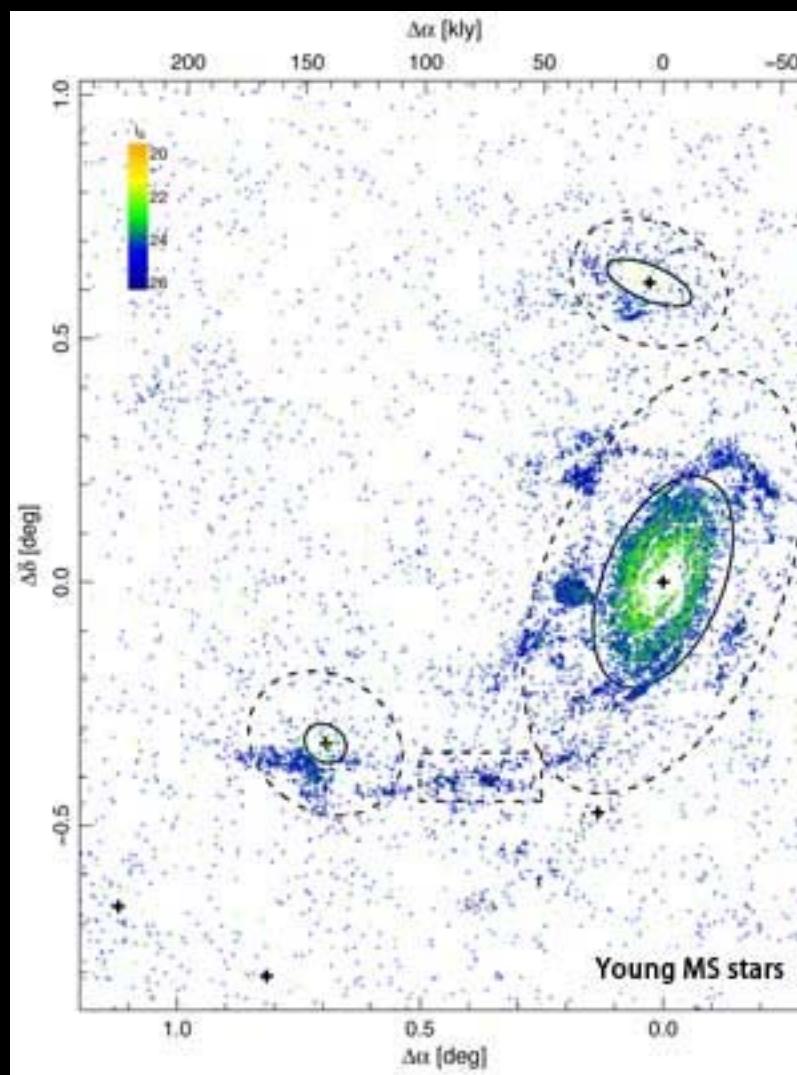
NGC3077

SE-stream

Ho IX

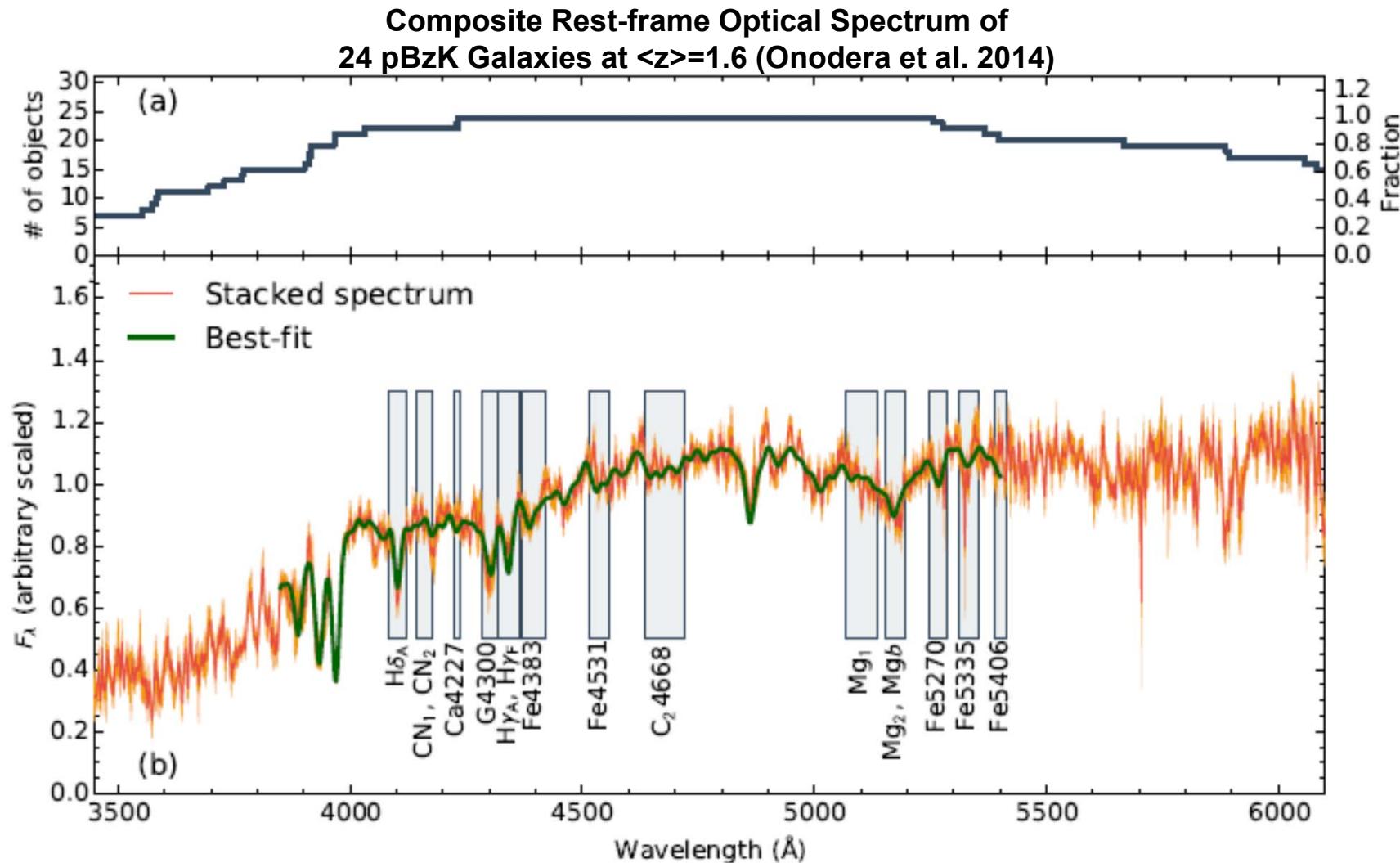
Subaru Telescope

Young main-sequence (MS) stars and red-giant branch (RGB) stars around M81, M82, and NGC 3077



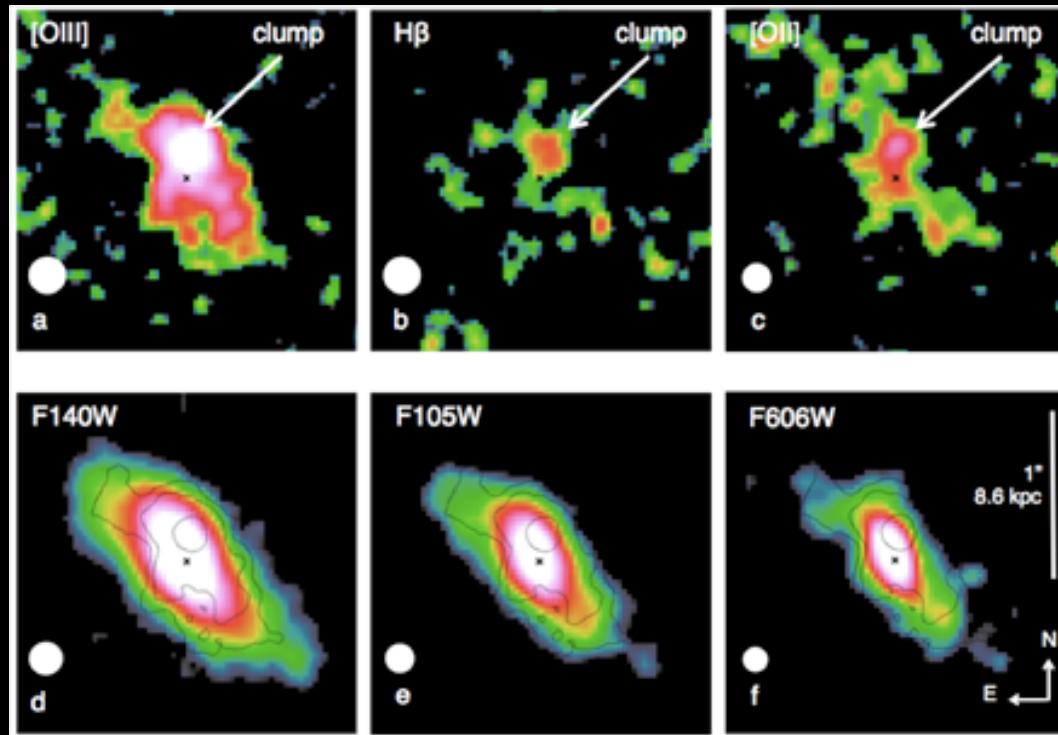
Subaru Telescope

Rapid evolutionary transition revealed in massive quenched galaxies

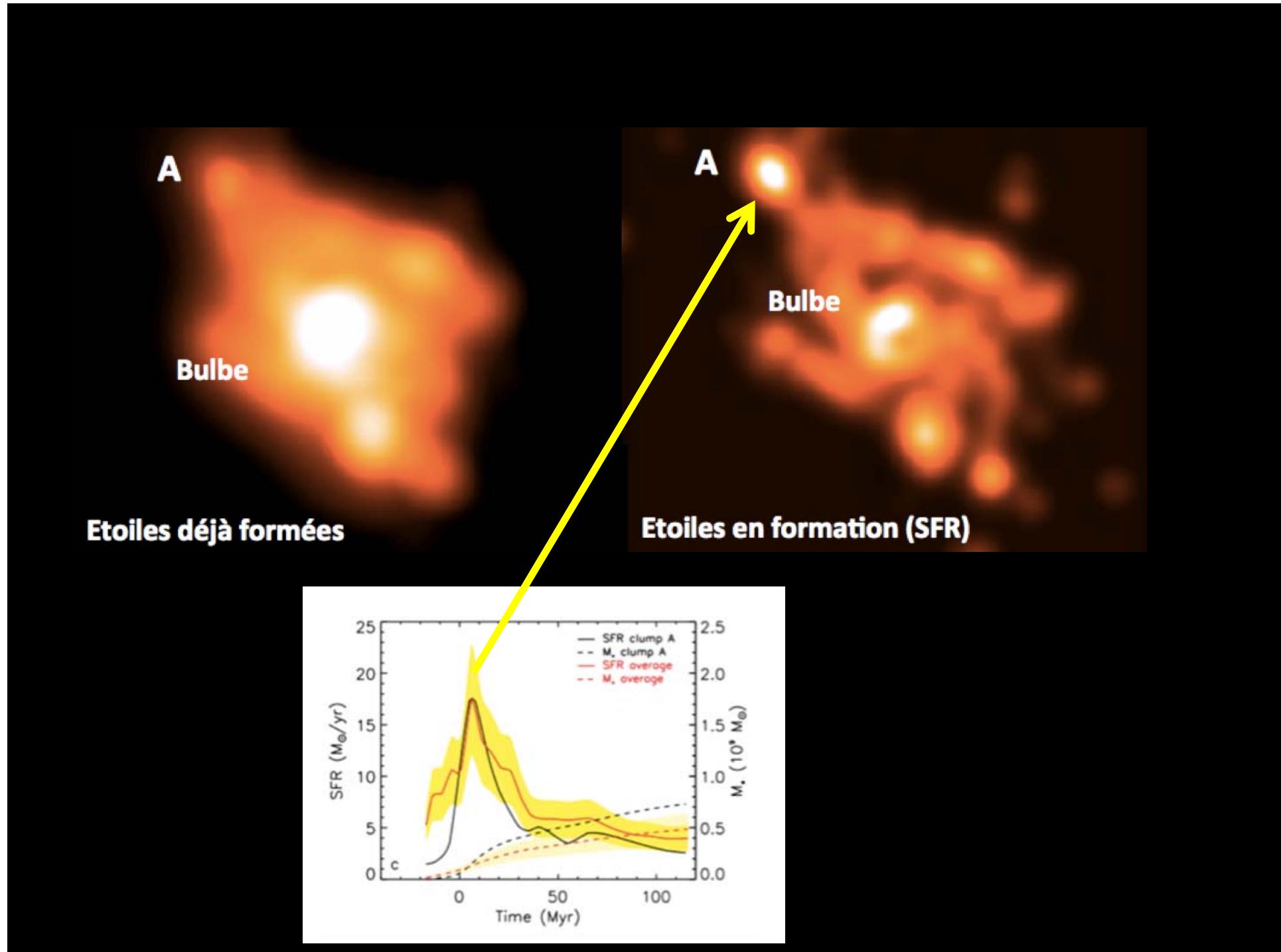


Subaru Telescope

Discovery of an Extremely Young Stellar Clump in the Distant Universe (Zanella et al. 2015)

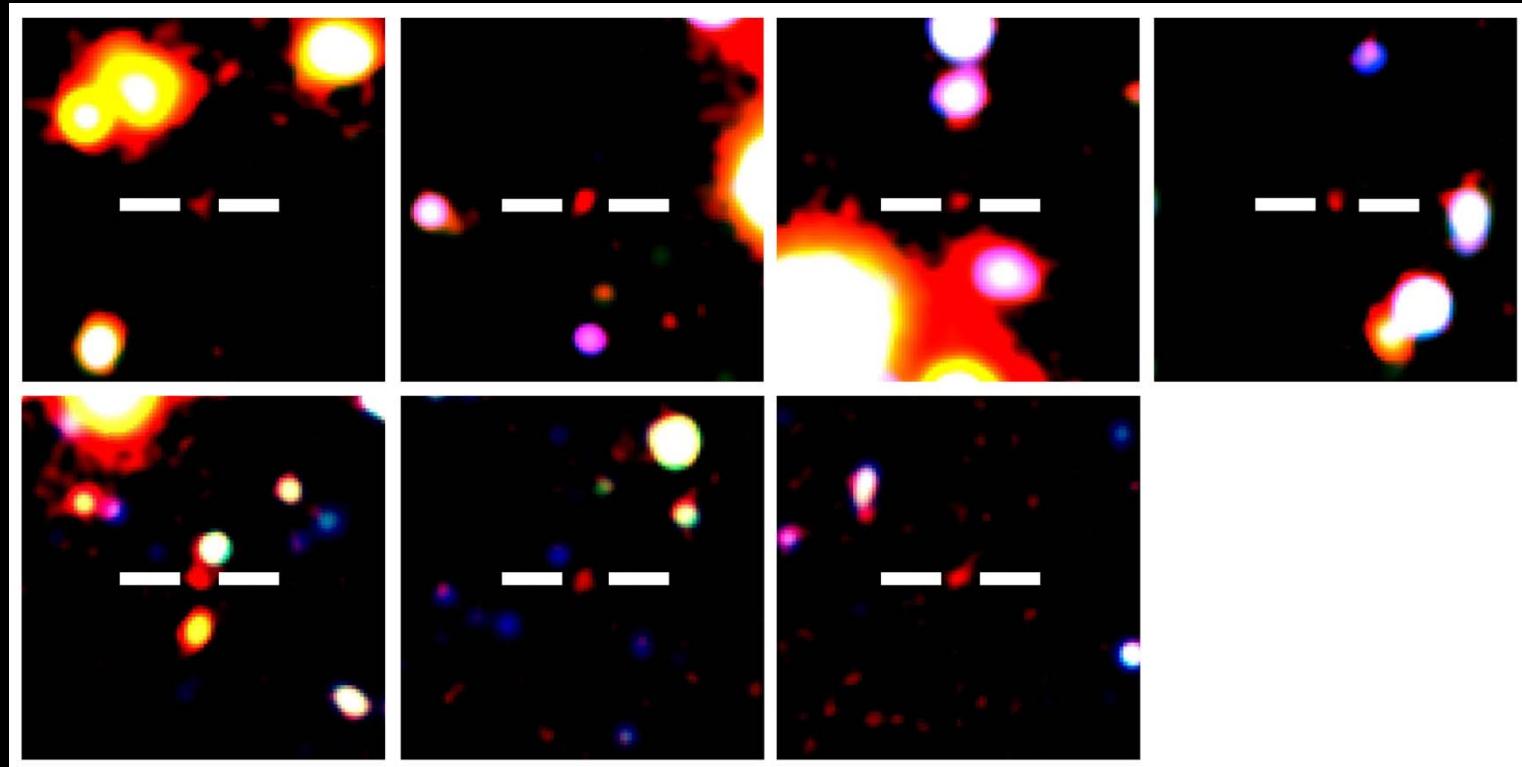


A group of French researchers discovered the birth cry of a massive star-forming clump in the disk of a very distant galaxy. This giant clump is less than 10 million years old, and it is the very first time that such a young star-forming region is observed in the distant Universe.



Subaru Telescope

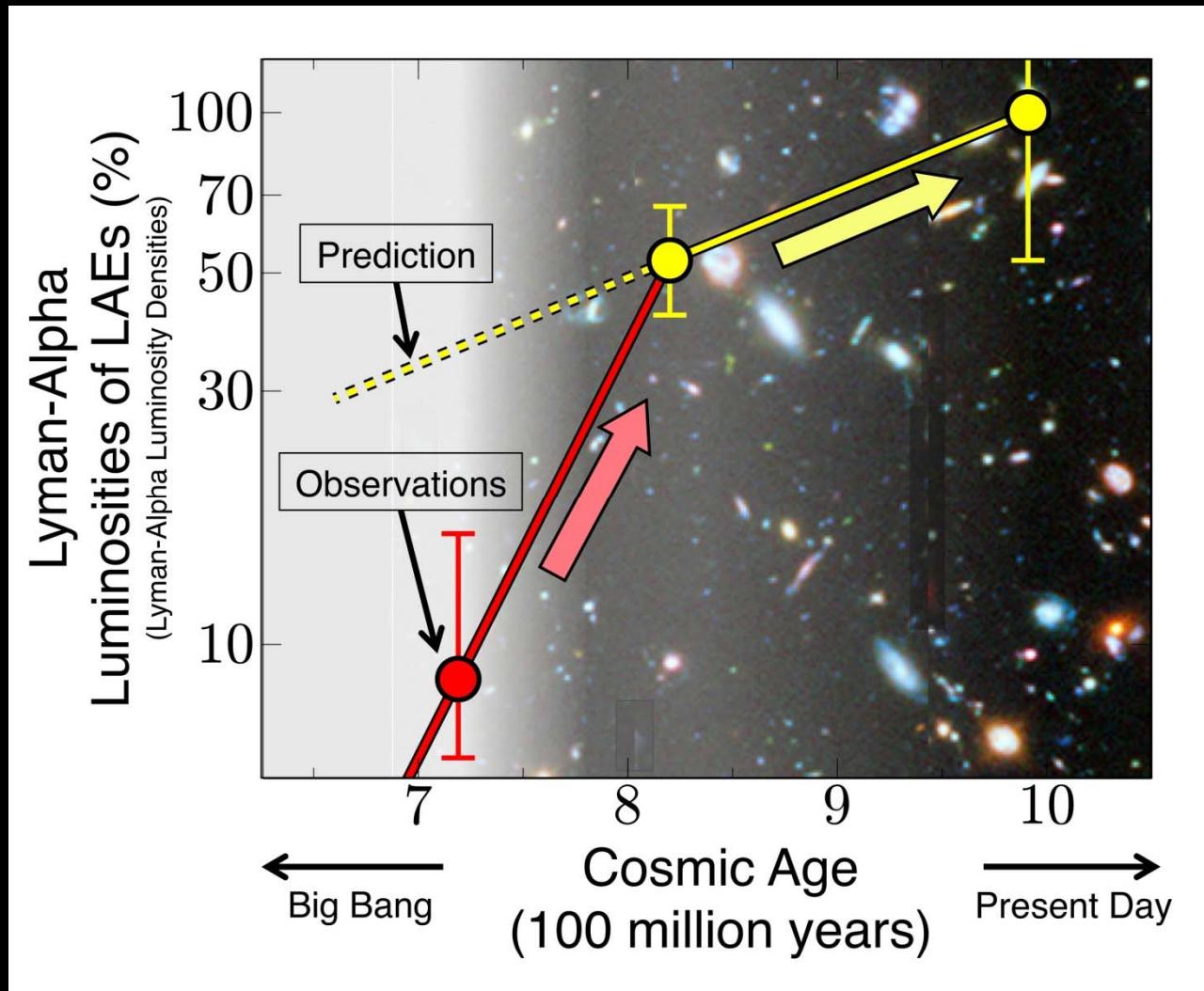
**Subaru Telescope Detects Sudden Appearance of Galaxies
in the Early Universe ($z=7.31$)**



A team of astronomers using the Subaru Telescope's Suprime-Cam to perform the Subaru Ultra-Deep Survey for Lyman-alpha Emitters have looked back more than 13 billion years to find 7 early galaxies that appeared quite suddenly within 700 million years of the Big Bang (Konno et al. 2014).

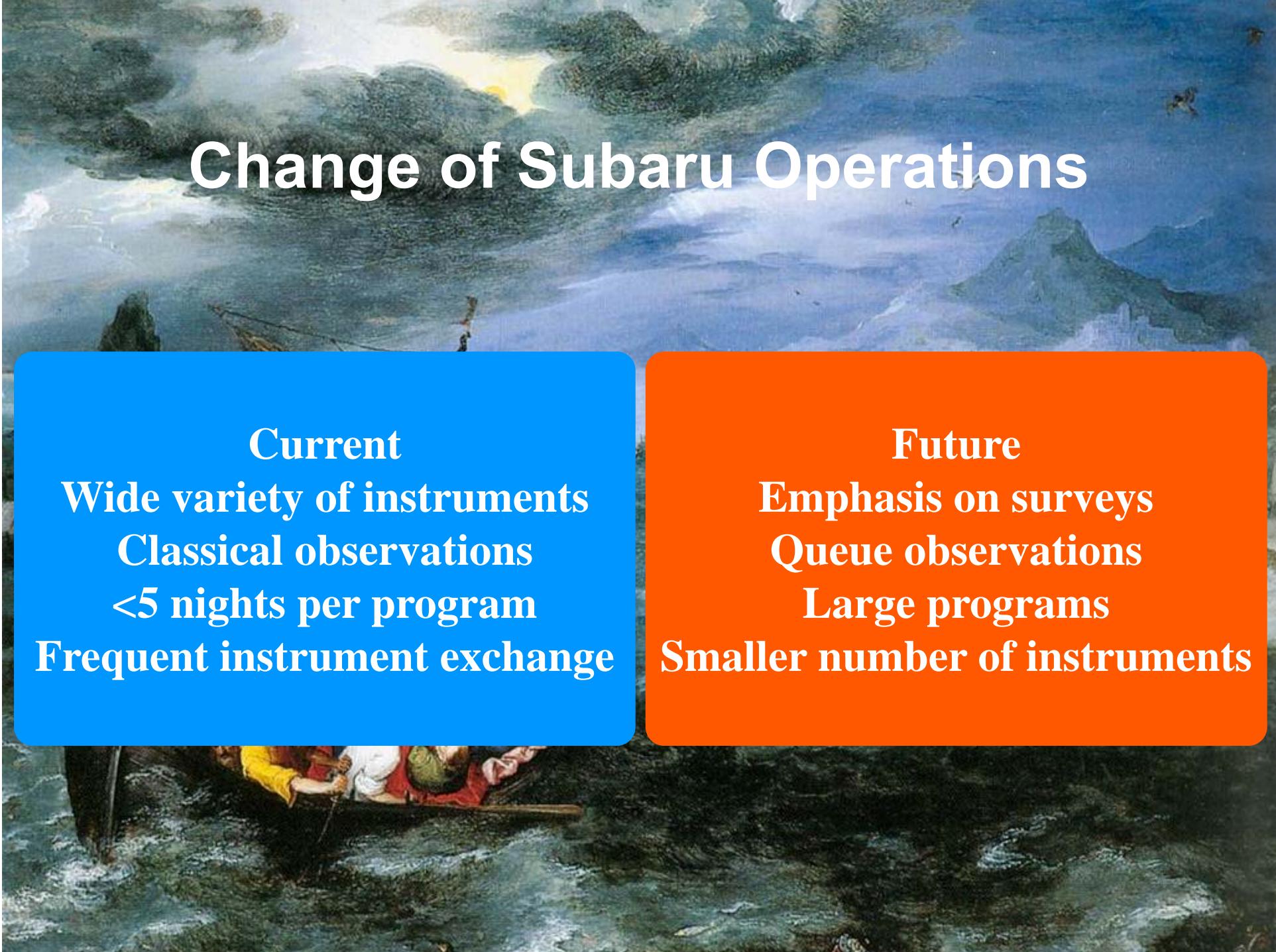
Subaru Telescope

Why Only 7 LAEs?





Subaru Telescope Instruments



Change of Subaru Operations

Current

Wide variety of instruments

Classical observations

<5 nights per program

Frequent instrument exchange

Future

Emphasis on surveys

Queue observations

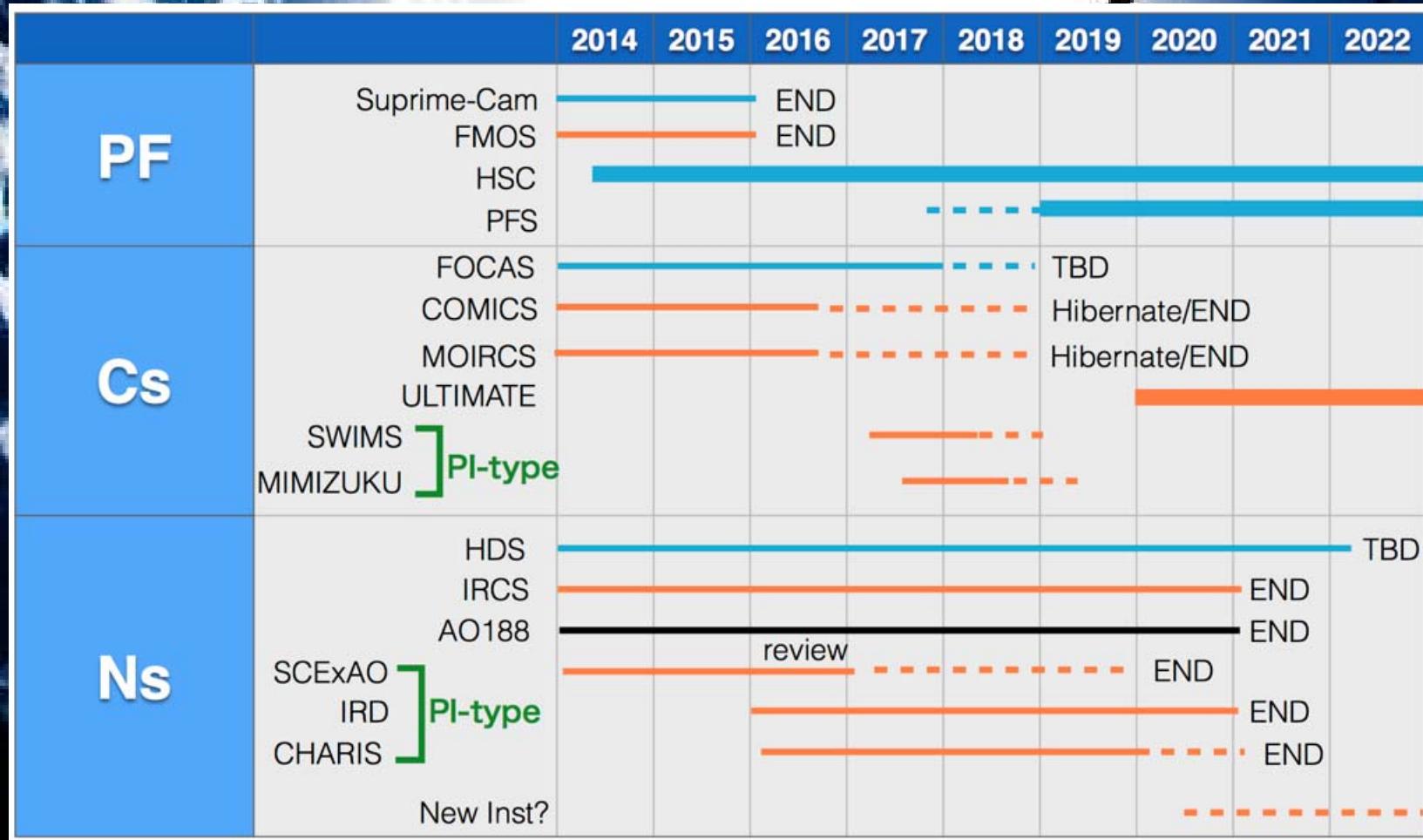
Large programs

Smaller number of instruments

Change of Subaru Operations

- Synergy with TMT - Subaru will conduct large surveys to feed targets to TMT.
- New generation instruments are very large and expensive. We cannot keep whole suite of instruments competitive to newer generation instruments in other telescopes.
- Strong pressure to reduce budget in the era of TMT operation.
- Accepting PI-type instruments
 - Subaru as a development platform
 - Cutting-edge technologies, sharp science cases

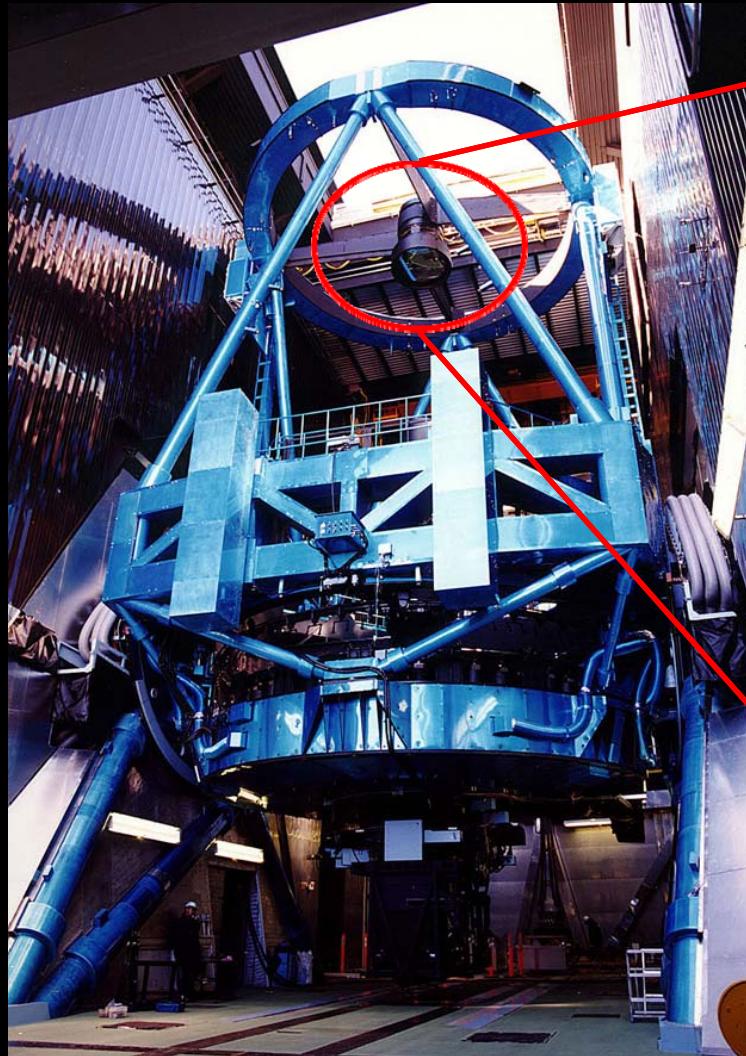
Instrument Timeline (Preliminary Decommission Plan)



**Hyper Suprime-Cam has started
Open-use in March 2014**



Hyper-Suprime-Cam (HSC)



HSC

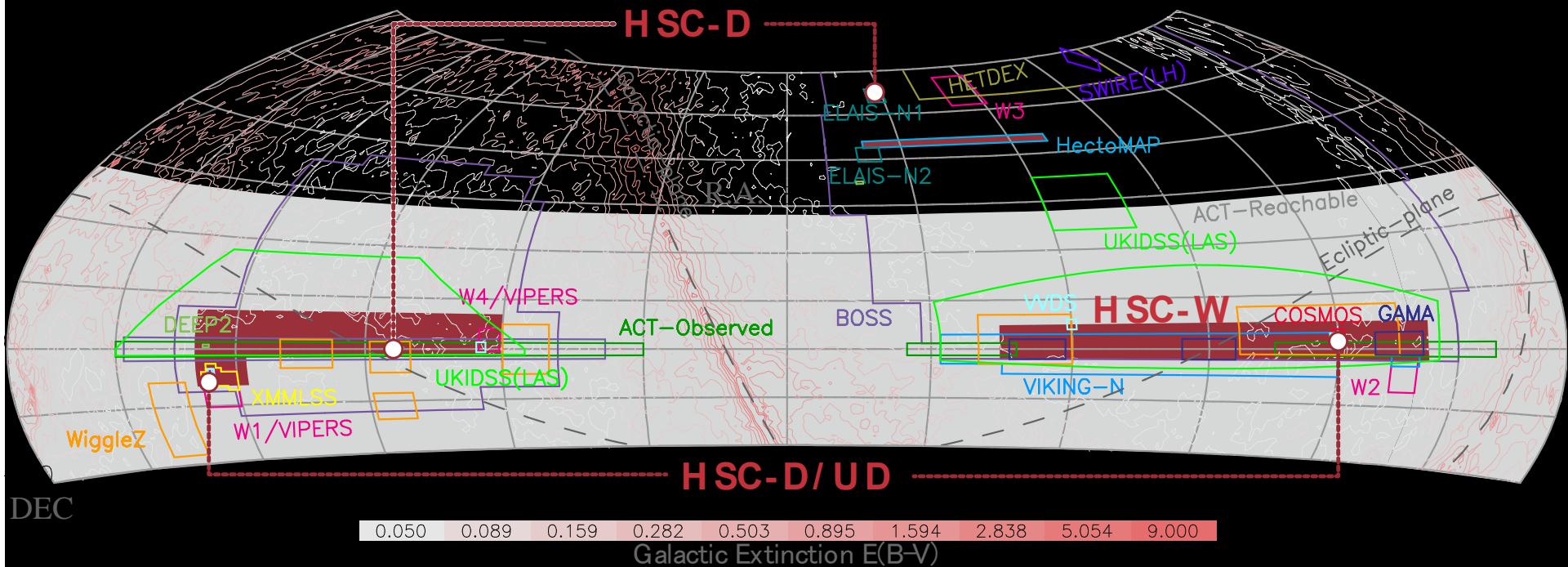
Wide Field

- Quicker, Wider



HSC
Field of View

HSC Survey Area



- **HSC Survey Area**
 - Include the previous surveys
 - Little absorption by dust
 - Observable whole year

**Subaru Strategic Program
2014 – 2020
300 Nights**

Image Quality of HSC

So far ~1.5 nights (Subaru has 8 instruments), but soon ~15 nights

Subaru HSC image (riz: ~2.5hrs)



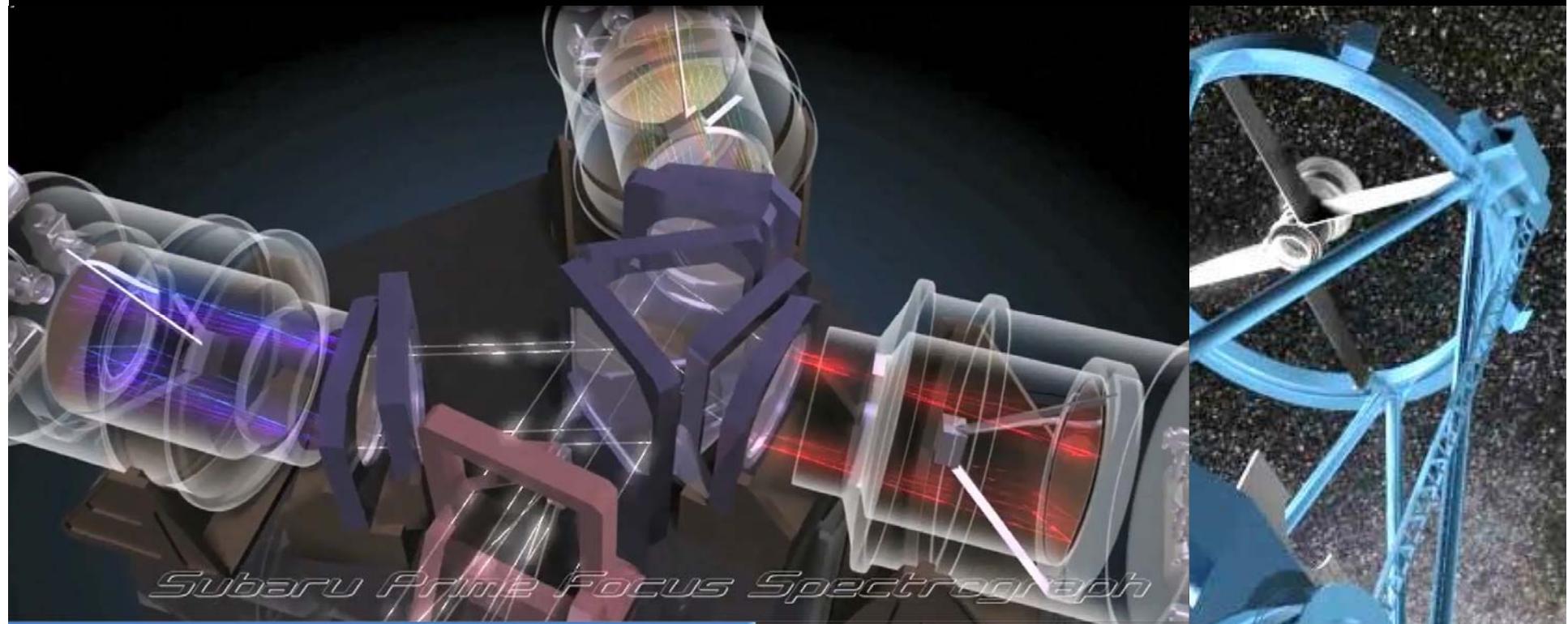
Data reduced by HSC pipeline
(NAOJ, Princeton, Kavli IPMU)

COSMOS HST (640 orbits: ~500hrs)



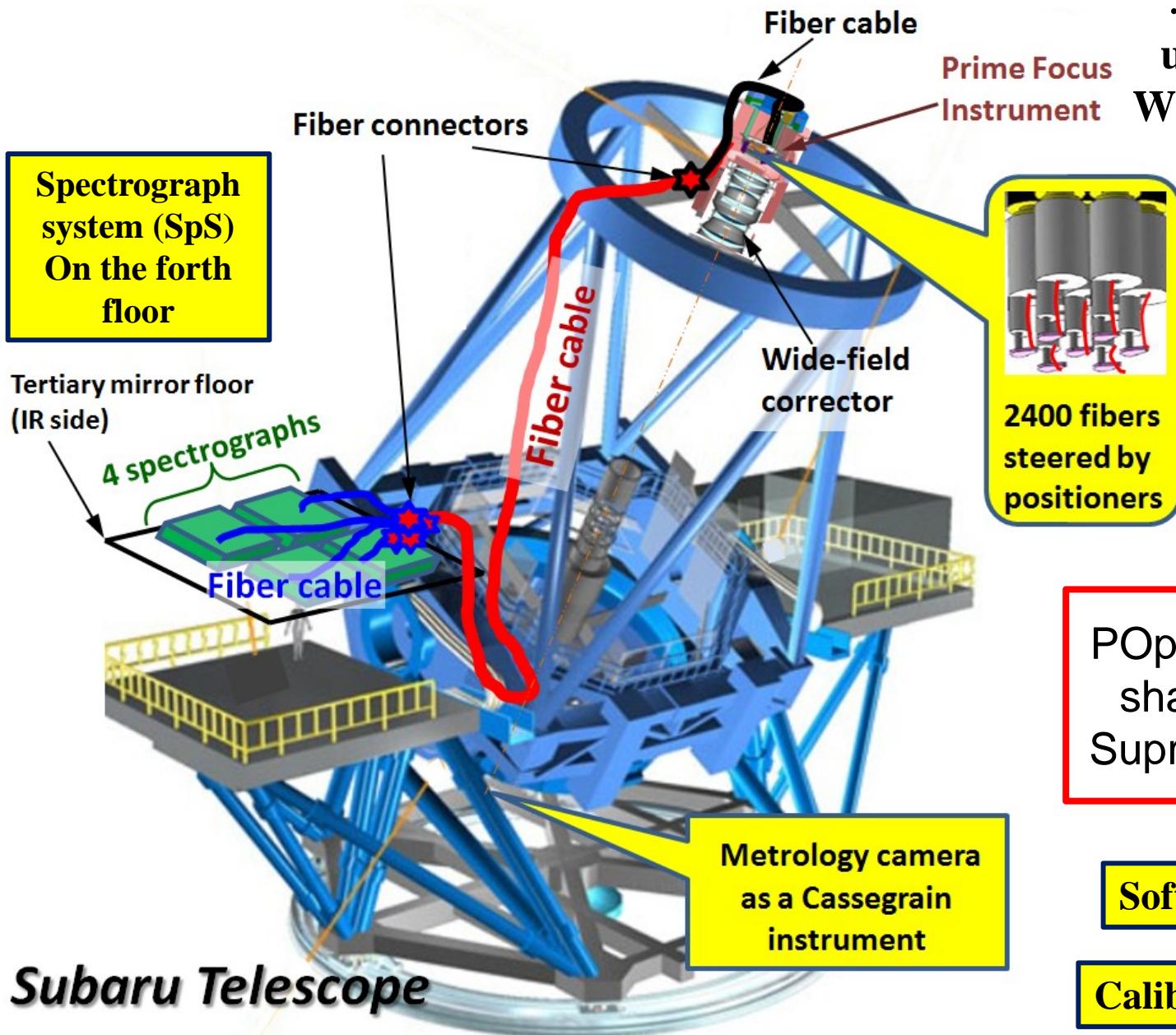
Hubble Telescope

SUBARU PRIME FOCUS SPECTROGRAPH



*Kavli IPMU,
U. Tokyo
PFS project office*





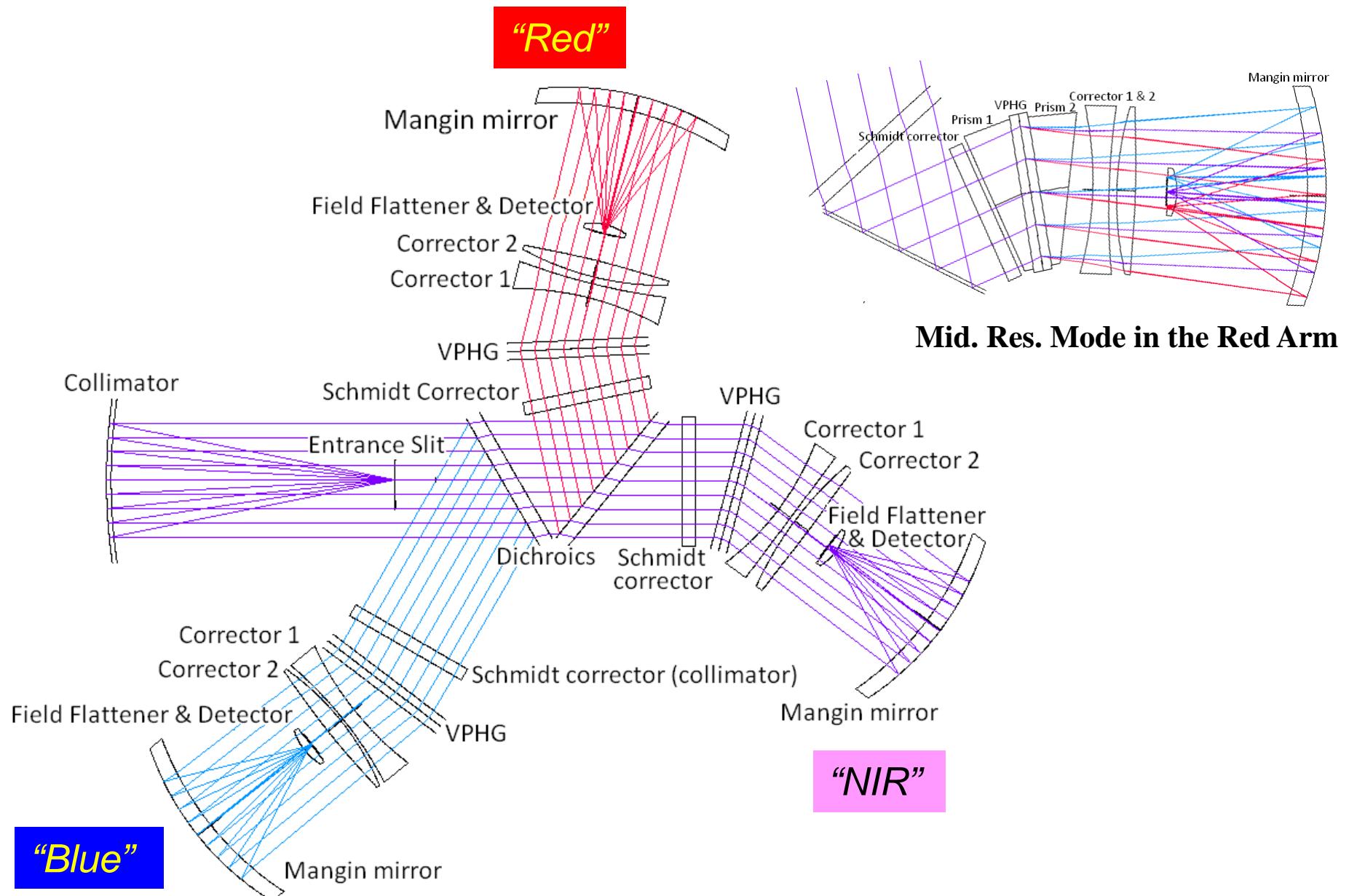
... in Prime focus
unit “POpt2” with
Wide Field Corrector
“WFC”.

POpt2 & WFC will be
shared with Hyper
Suprime Cam (HSC).

Software system

Calibration system

Spectrograph System



	Blue	Red		NIR
		Low Res	Mid Res	
Collimator F/#			2.5	
Camera F/#			1.1	
# of science fibers			597 or 600	
Operating temperature			3 +/- 0.5 degC	
Input fiber core diameter			129um	
Wavelength coverage	380-650nm	630-970nm	710-885nm	940-1260nm
Resolving power	~2300	~3000	~5000	~4300
Detector	Type	CCD (Hamamatsu, with a new blue coating)	CCD (Hamamatsu, with the same coating as HSC)	H4RG (Teledyne, 1.7um cutoff)
	Format	(4K x 2K) x 2	(4K x 2K) x 2	4K x 4K
	Pixel size		15um	
	Readout noise	~4 e-/pix		4 e-/pix
	Dark	~0.4 e-/pix/hour		0.01 e-/pix/s
Thermal background		None		0.006 e-/pix/s

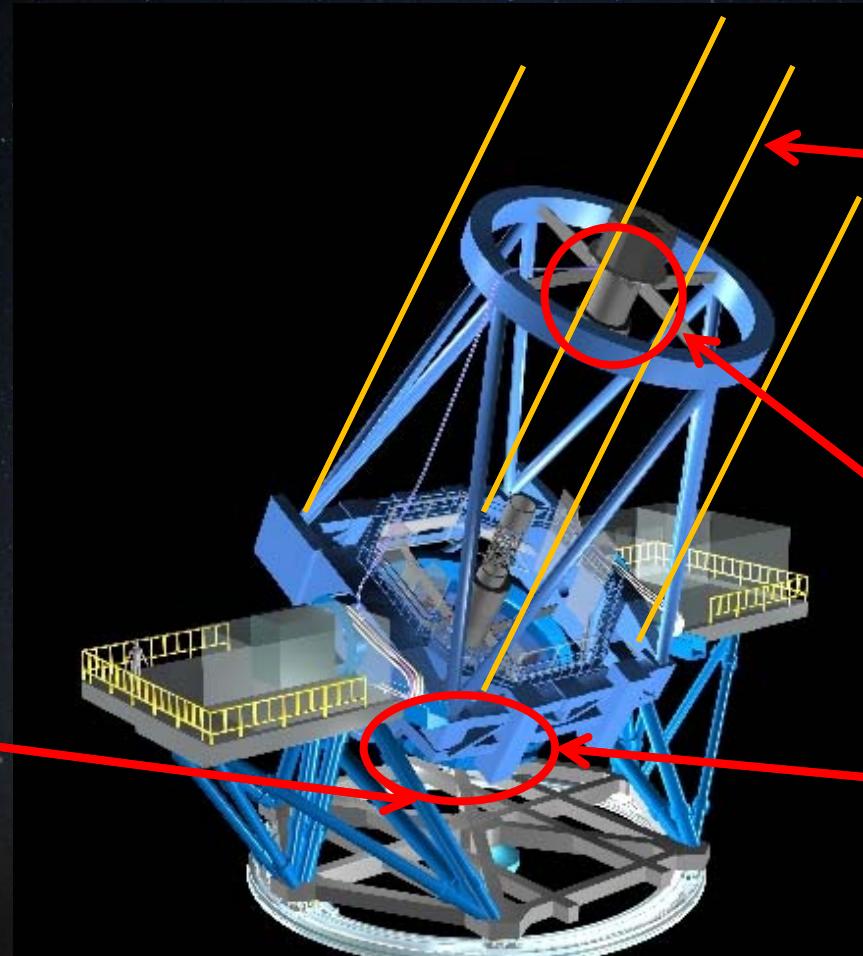
Three Pillars of Subaru/PFS Survey Science

- Cosmology (100 nights)
- Galaxy & AGN evolution (100 nights)
- Galactic Archaeology (GA) (100 nights)

ULTIMATE-Subaru

NIR inst.

14' FoV
Multi-Object IFU
Wide-field Camera



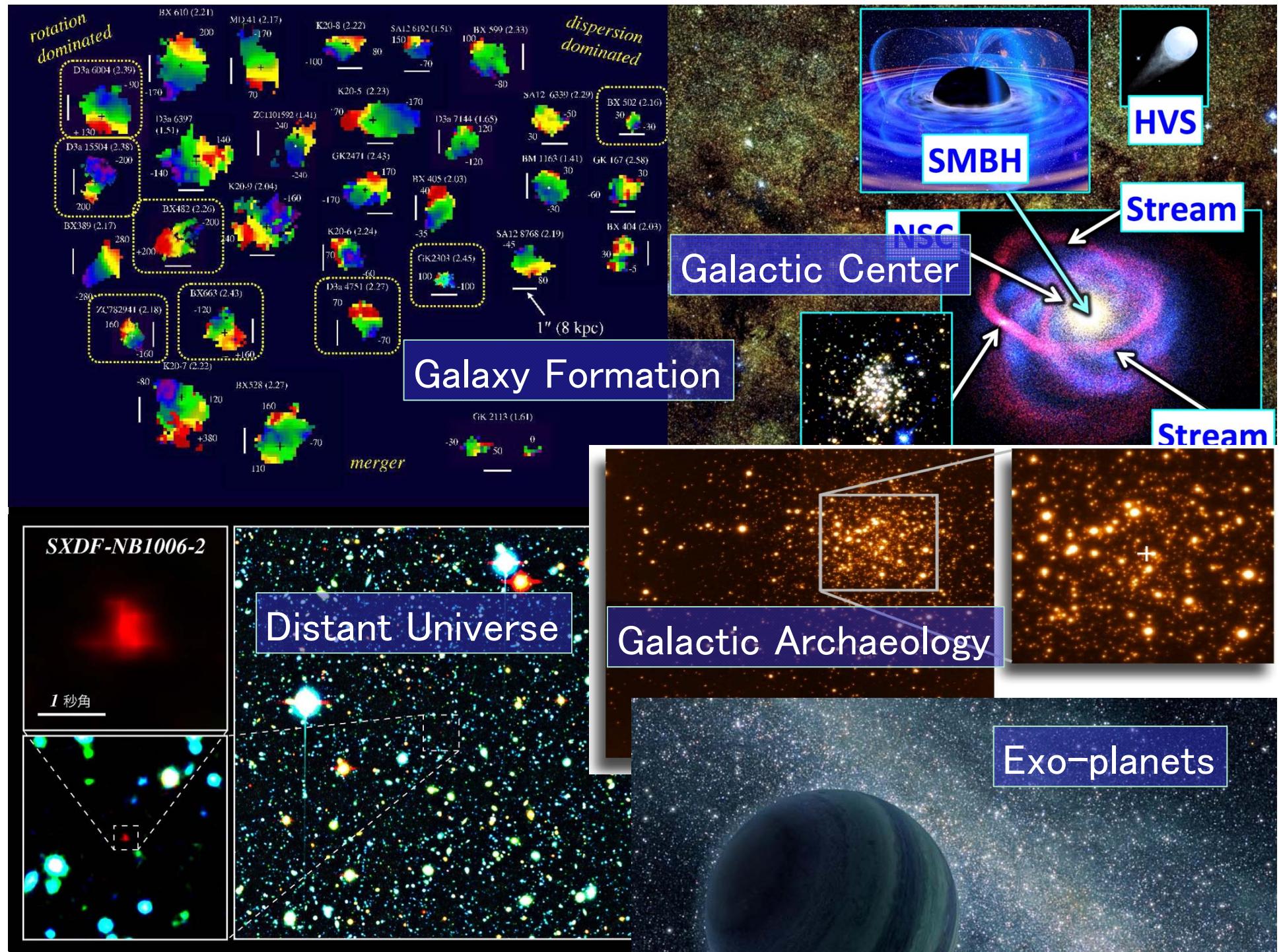
4 Lasers
(side irradiation)

Deformable
secondary
mirror

Wave front
sensor

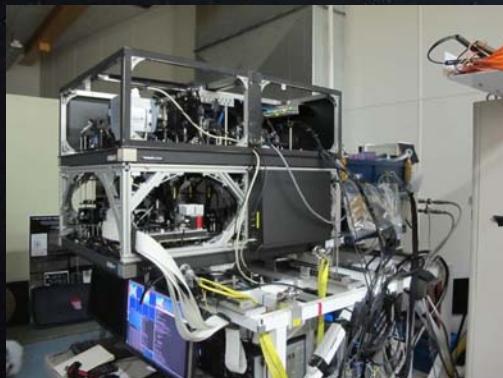
Science of ULTIMATE Subaru

- **Dissect the Galactic Evolution of the Golden Age**
 - Sample of a few 1000 galaxies at $1 < z < 3$ - morphology, dynamical structure, physical parameters, environmental effects, internal motion of stars/gas, AGN contribution, heavy elements distribution
- **Discover Galaxies at the Edge of the Universe**
 - Search for highest-redshift galaxies with highly sensitive narrow band imaging ($z > 10$)
 - Discover galaxies at $z > 7.5$, physical process of cosmic re-ionization
- → **Subaru Original Samples for TMT**
 - Sampling of the most interesting targets cannot be done by HSC + PFS alone.
 - Improvement of Telescope Performance - contributing to various science

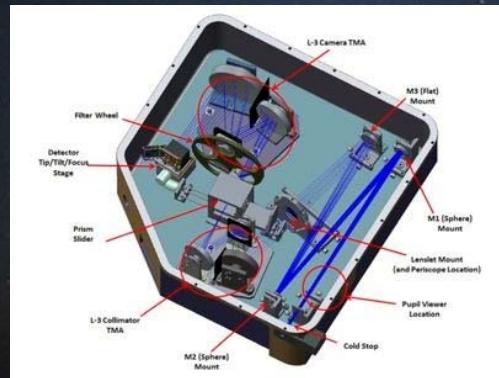


PI-type Instruments for Exoplanets

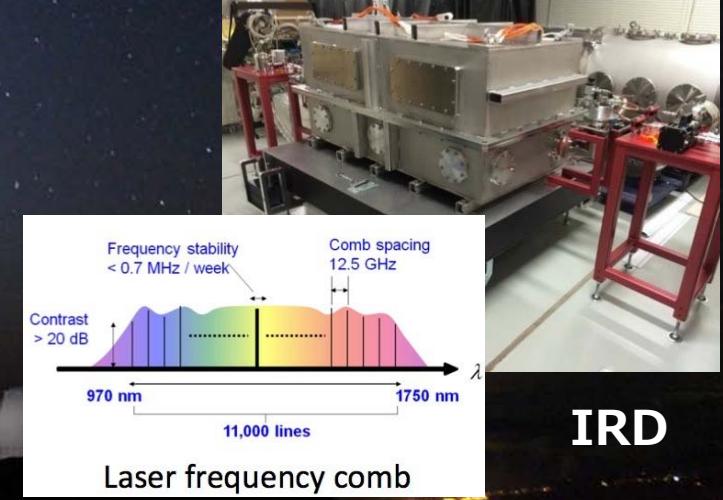
- SCEExAO: Coronagraphic Extreme-AO (direct imaging)
- CHARIS: Integral Field Spectrograph (discover and characterization)
- IRD: Near-IR High-dispersion Spectrograph (Earth-mass planets around M-dwarfs)
- Short time-scale of Cycles (Development-Commissioning-Science) to catch-up this rapidly evolving research field



SCEExAO



CHARIS

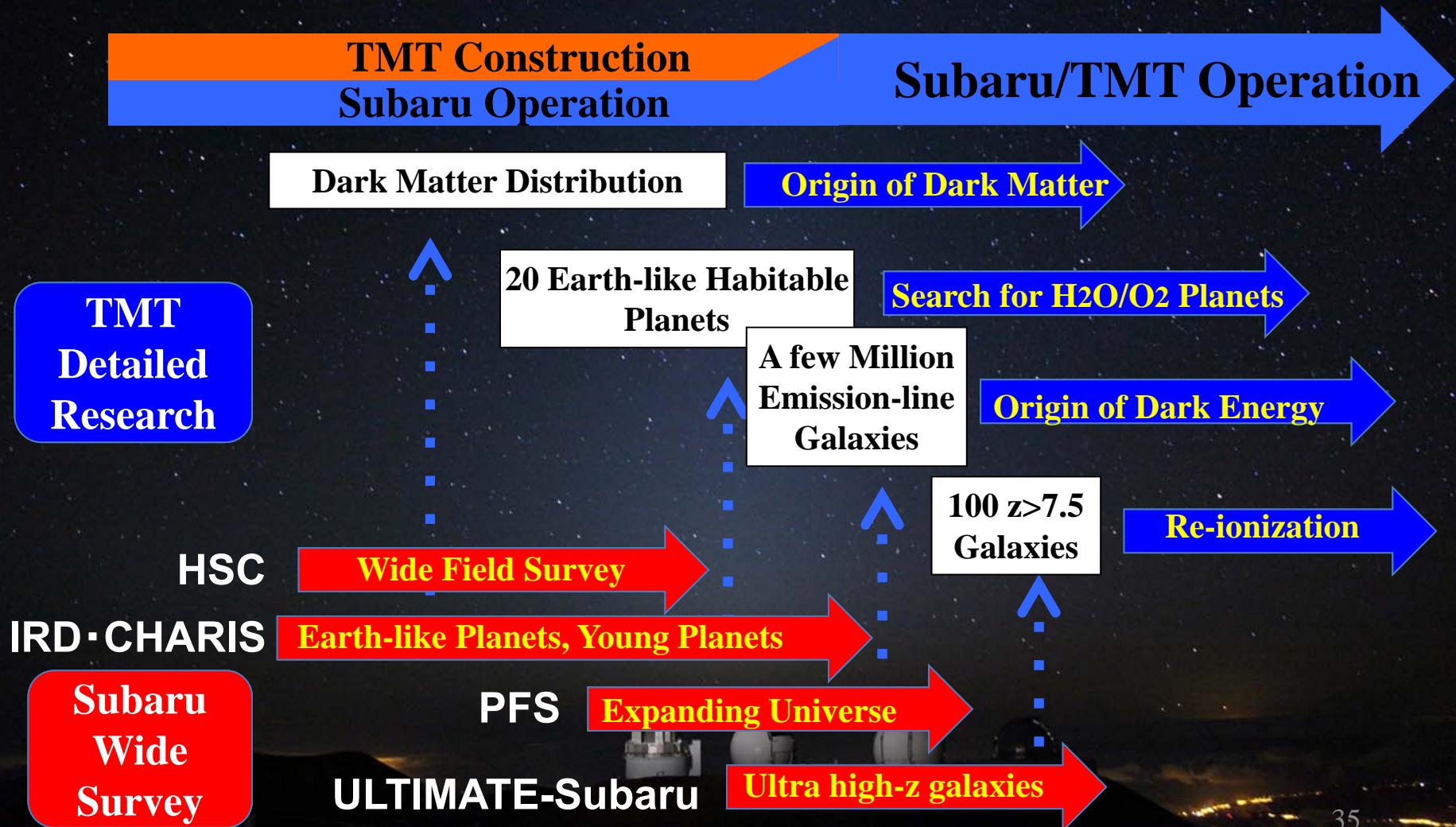


Laser frequency comb

IRD

Strategy from Subaru to TMT

2012 2014 2016 2018 2020 2022 2024 2026 2028





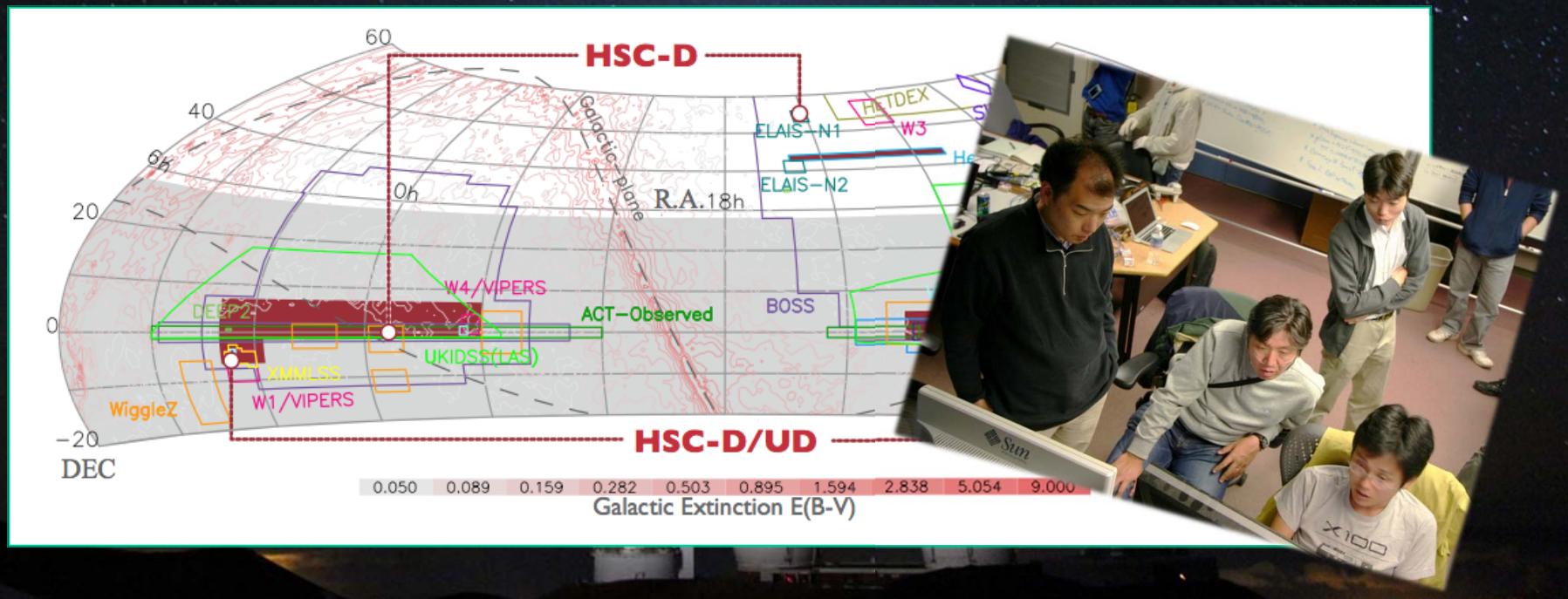
Subaru Strategic Program

Subaru SSP ③

HSC Subaru Strategic Program

“Wide Field imaging with Hyper Suprime-Cam”
“Cosmology and Galaxy Evolution”

S.Miyazaki (PI) 300 nights (2014 - 2018)



Subaru SSP ④

IRD Subaru Strategic Program

“Search for the Earth in Habitable Zone

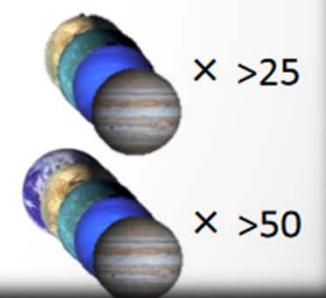
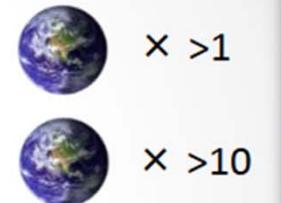
T.Kotani (PI) 150 (?) nights (2017 - 2021)



1. Detection of habitable Earth-like planets around nearby M dwarfs
 - Minimum Success
 - Detection of at least 1, one Earth-mass planet in their HZ
 - Full Success
 - Unveiling frequency and properties of habitable Earth

2. Statistical understandings of planet formation around low-mass stars
 - Minimum Success
 - 25 > Super-Earth - Jupiter-mass planets around low-mass stars
 - Full Success
 - 50 > planets including Earth-mass planets

Number of Detection



Subaru SSP ⑤

PFS Subaru Strategic Program

“Galactic Archaeology, AGN & Galaxy Evolution,
and Cosmology”

H.Murayama (PI) 300 nights (2019 - 2023)



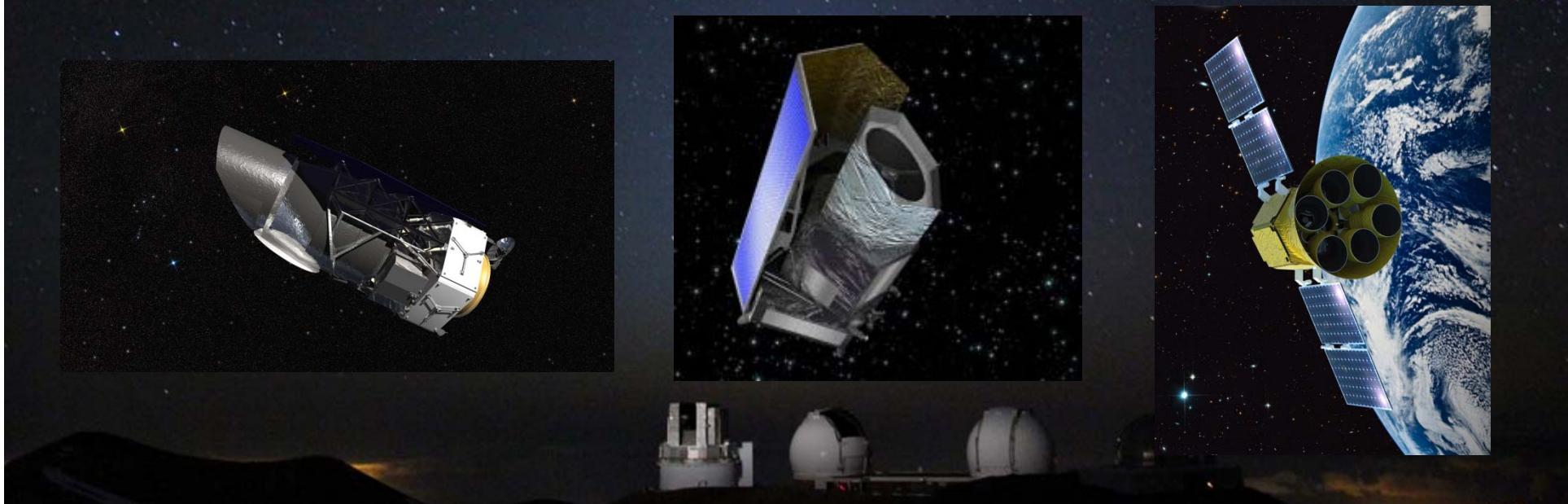
Synergy with Space Missions

Subaru Strategic Program (?)

WFIRST 100-400 nights (2025 – 2030)

Euclid 100 nights (2019 – 2022)

TESS 40 nights (2017 – 2020)





International Collaboration in Astronomy for Asia

- Subaru can offer opportunities for students in Asian countries to have an access to Subaru Data.
- ① Apply for Open Use Program jointly with Japanese PI's.
- ② Contact CoIs of SSP and have an access to HSC/PFS/IRD data
- ③ Use Subaru Archive Data.

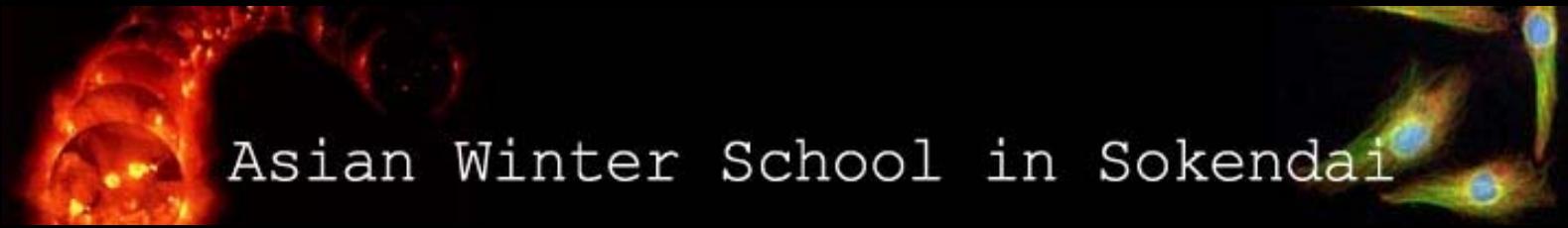
Subaru Winter School in Korea

2014/02/25 – 2014/02/27 (KASI)





SOKENDAI Asian Winter School



Asian Winter School in Sokendai

觀測裝置制御計算機群
Instrument Control Computers

Subaru Internship

