

PROBING ACCRETION PROCESSES THROUGH VARIABILITY

2016 TMT Science Forum

Kyoto, Japan

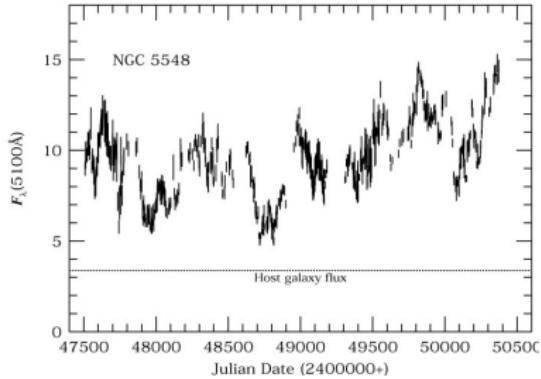
VISHAL PRAMOD KASLIWAL

vishal.kasliwal@gmail.com

Department of Physics & Astronomy
University of Pennsylvania
&
Dept. of Astrophysical Sciences
Princeton University

May 26, 2016

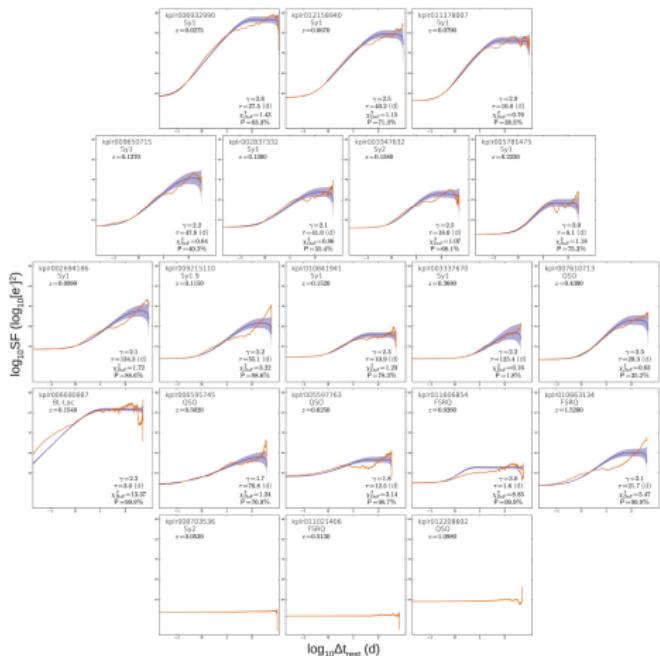
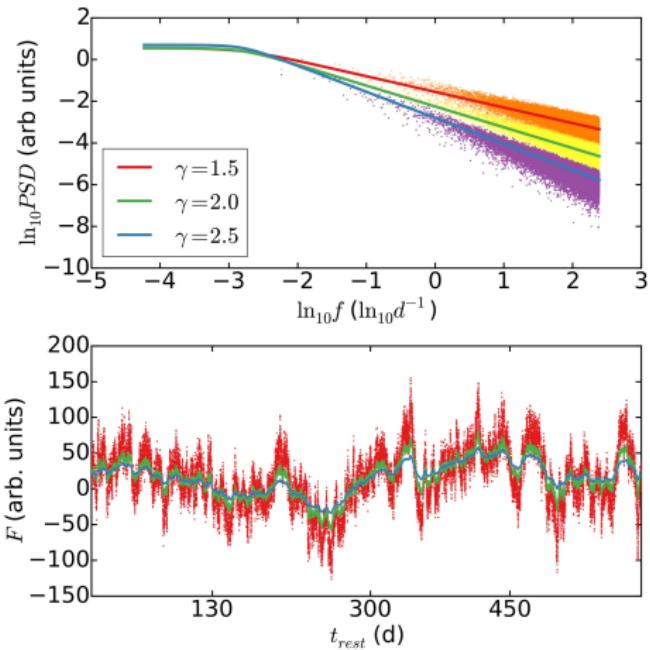
Accretion Characteristic: Rapid, Stochastic, Luminosity Variations (and we do not know why!)



(Peterson et al. 1999)

- ✿ ~ 90 % of AGN vary (Sesar et al. 2007)
- ✿ Pan-spectral: shorter $\lambda \Rightarrow$ stronger variability
- ✿ What can TMT contribute?
- ✿ Stochastic! (Peterson 1997)
- ✿ λ_{long} lag λ_{short} (but sometimes backwards!)

Beyond the Damped Random Walk

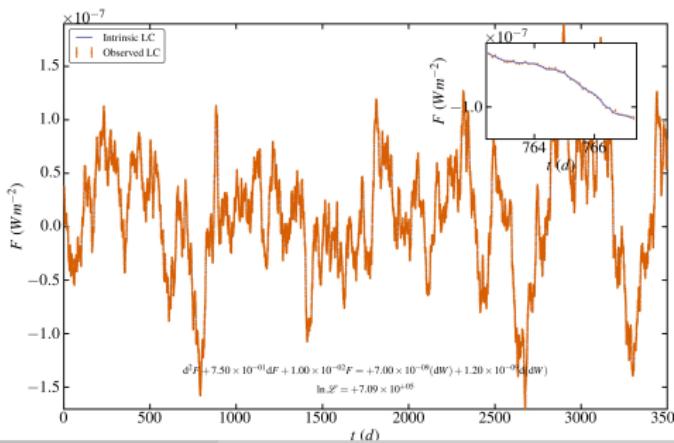


Continuous-time AutoRegressive Moving Average (C-ARMA) Processes

$$dW \sim \mathcal{N}(0, dt)$$

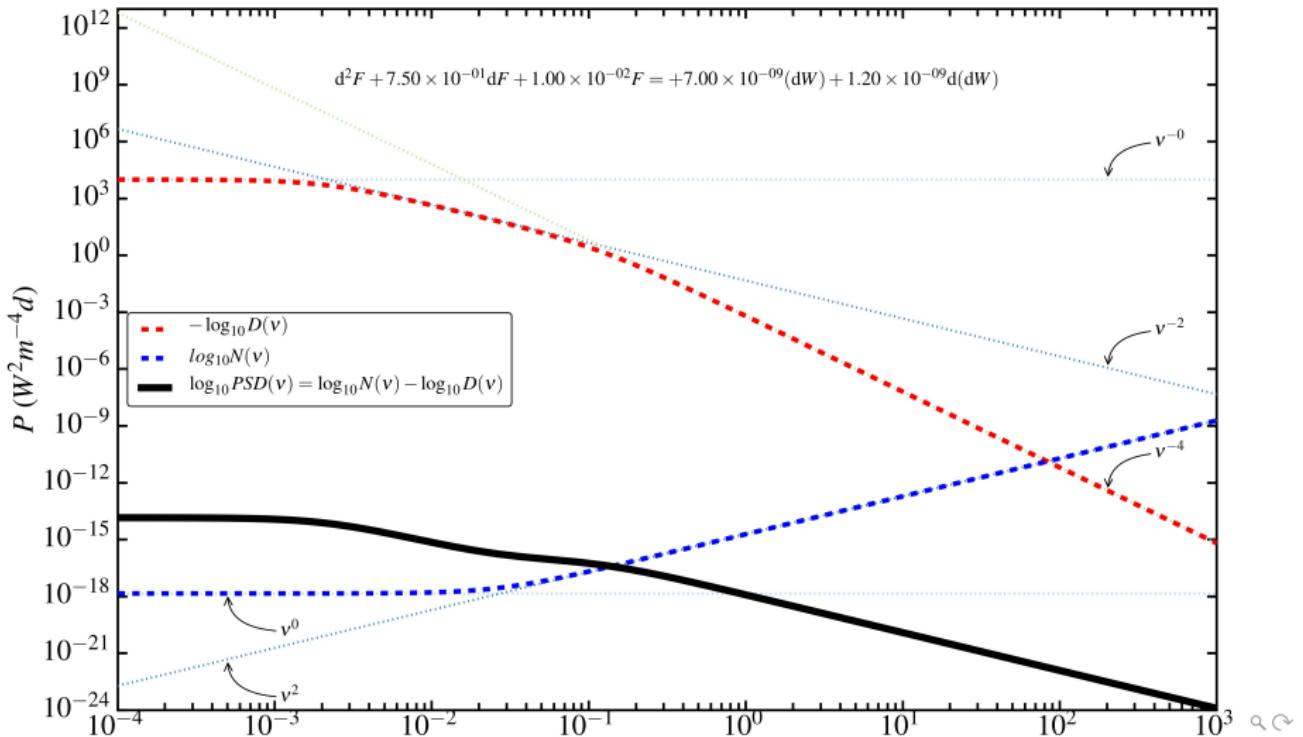
$$d^p x + \alpha_1 d^{p-1} x + \dots + \alpha_{p-1} dx + \alpha_p x = \beta_0(dW) + \dots + \beta_q d^q(dW)$$

- * Uses Itô calculus Davis (2002); Brockwell (2014); Kelly et al. (2014)
- * RHS: Correlation structure of all perturbations.
- * LHS: Evolution of individual perturbations.
- * PSD: ratio of even polynomials of frequency.

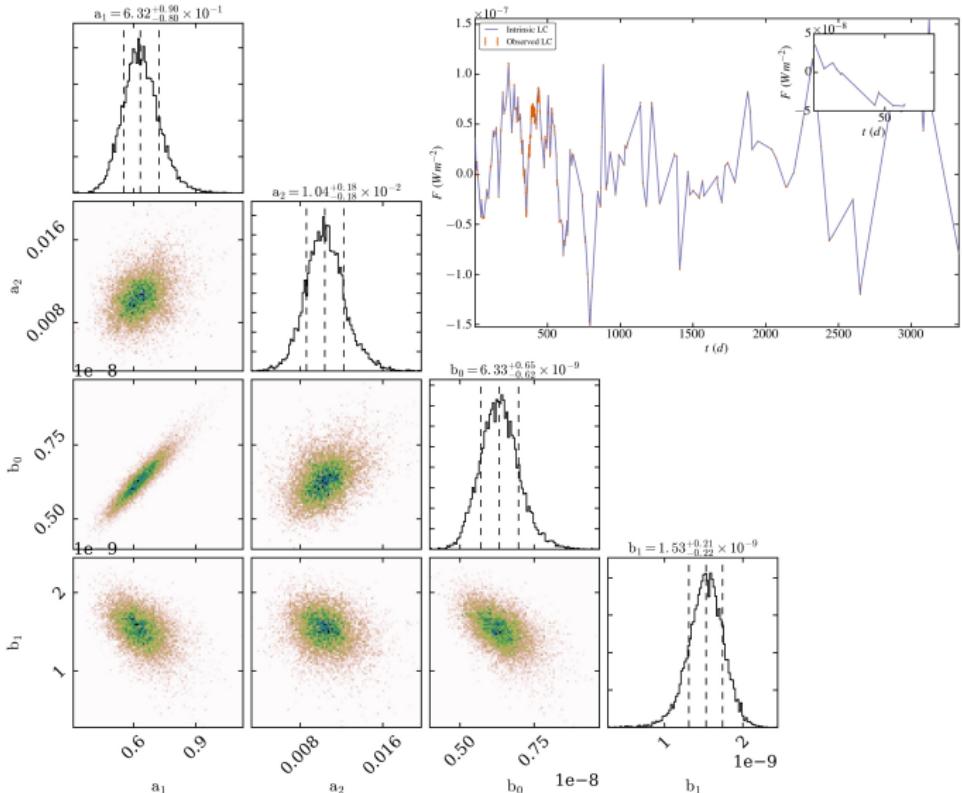


Power Spectral Density

Eg. C-ARMA(2,1)

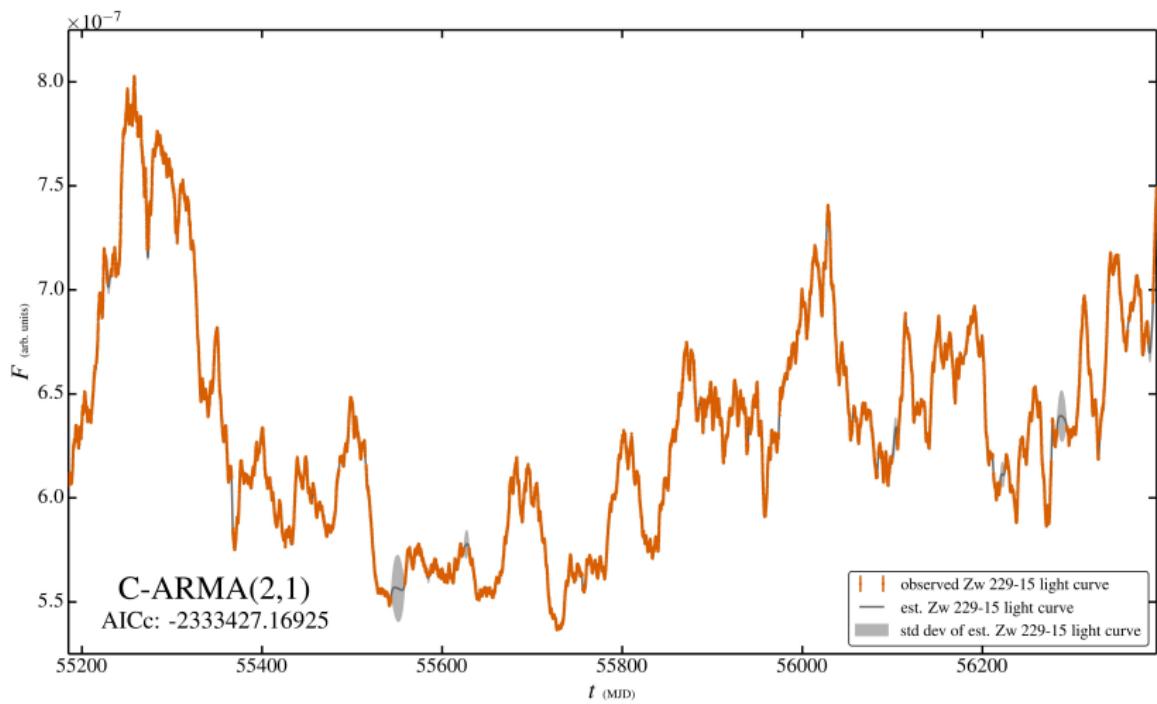


LIBCARMA (<https://github.com/AstroVPK/libcarma>)

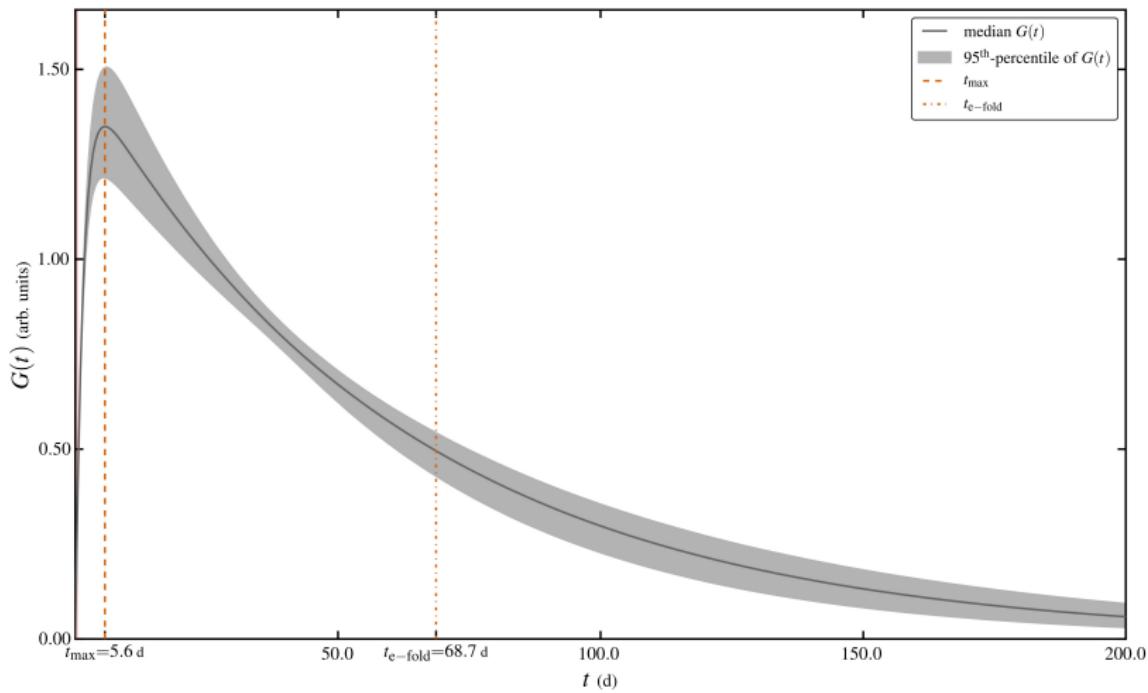


C-ARMA(2,1) model of Zw 229-15

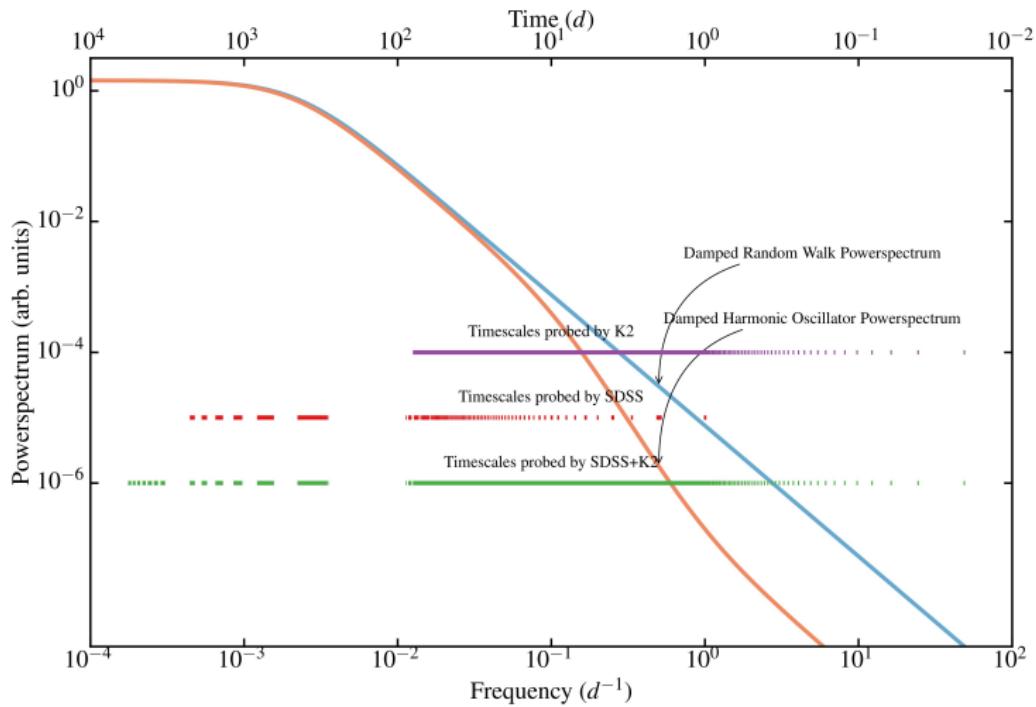
Smoothed light curve



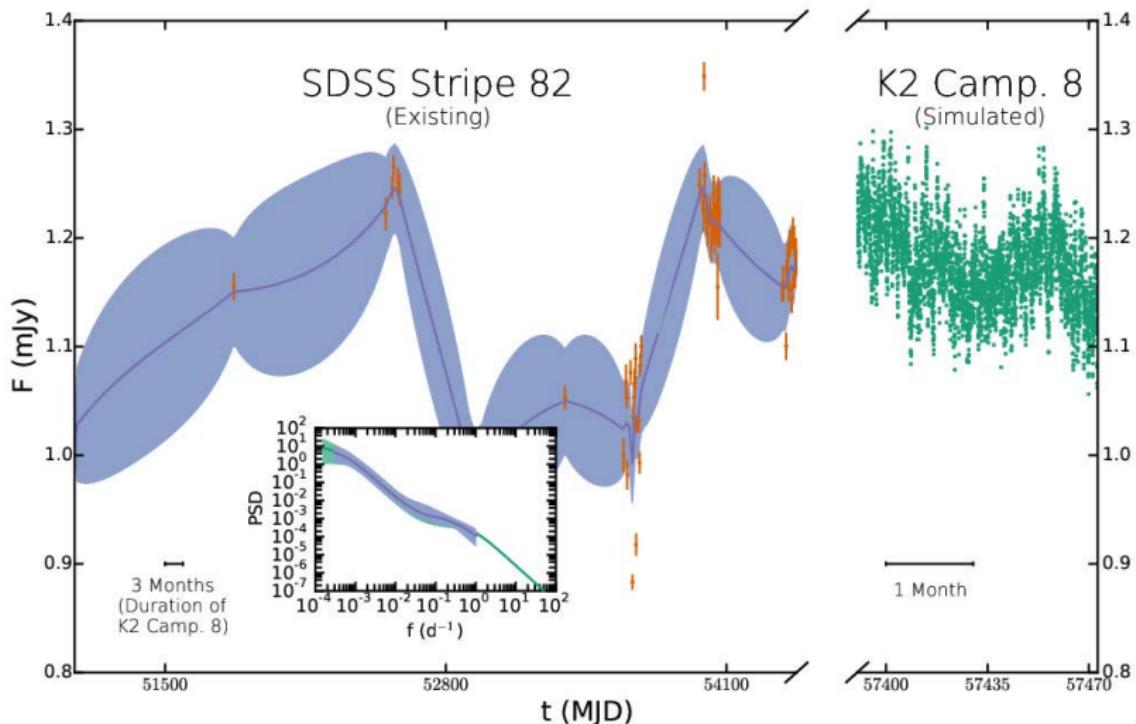
C-ARMA(2,1) model of Zw 229-15 Green's Function



Prototype for LSST+TMT: Combining SDSS with K2



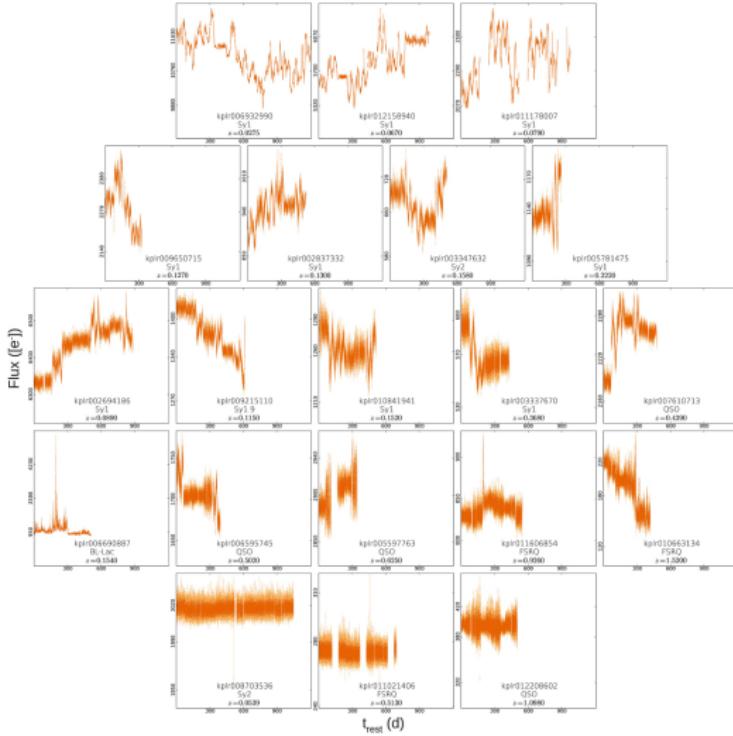
Prototype for LSST+TMT: K2 observations of Stripe 82 QSOs



Conclusions

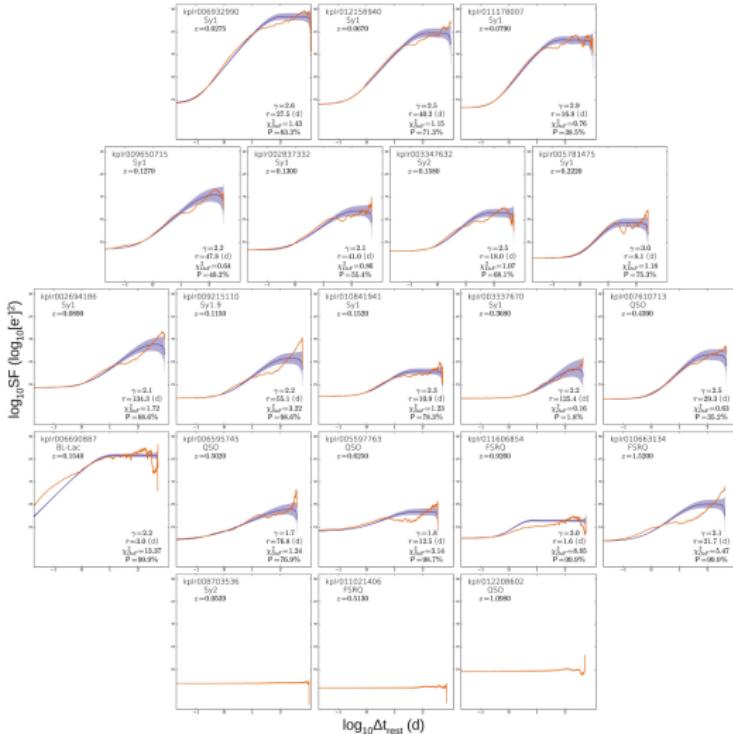
- * We need good data!
- * Good data shows: DRW does **not work** for all AGN
- * Variability \Rightarrow C-ARMA process
- * C-ARMA(2,1) process \Rightarrow Zw 229-15 variability
- * Zw 229-15 \Rightarrow **Damped Harmonic Oscillator + Colored Noise**
- * Future time-domain surveys: **combine multiple data sources**
- * TMT: Rapid sampling + low noise
- * LSST: Long baseline

Full AGN sample



- * $z \sim 0.02\text{-}1.5$
- * $\delta t_{\text{rest}} \sim 14\text{-}28 \text{ min}$
- * $N \sim 16k\text{-}60k$
- * Wide variety of behavior!

Structure Function fits



- * Not all AGN ~ DRW
- * PSD model too simple
- * Variability onsets over ~ 1 hr to ~ 1 d

Kasliwal, Vogeley, & Richards (2015)

Brockwell, P. 2014, Ann. Inst. Stat. Math., 66, 647

Davis, J. H. 2002, Foundations of Deterministic and Stochastic Control
(Birkhäuser)

Kasliwal, V. P., Vogeley, M. S., & Richards, G. T. 2015, MNRAS, 451, 4328

Kelly, B. C., Becker, A. C., Sobolewska, M., Siemiginowska, A., & Uttley, P.
2014, ApJ, 788, 33

Peterson, B. M., Barth, A. J., Berlind, P., et al. 1999, ApJ, 510, 659

Peterson, Bradley, M. 1997, An Introduction to Active Galactic Nuclei
(Cambridge University Press)

Sesar, B., Ivezić, Ž., Lupton, R. H., et al. 2007, AJ, 134, 2236