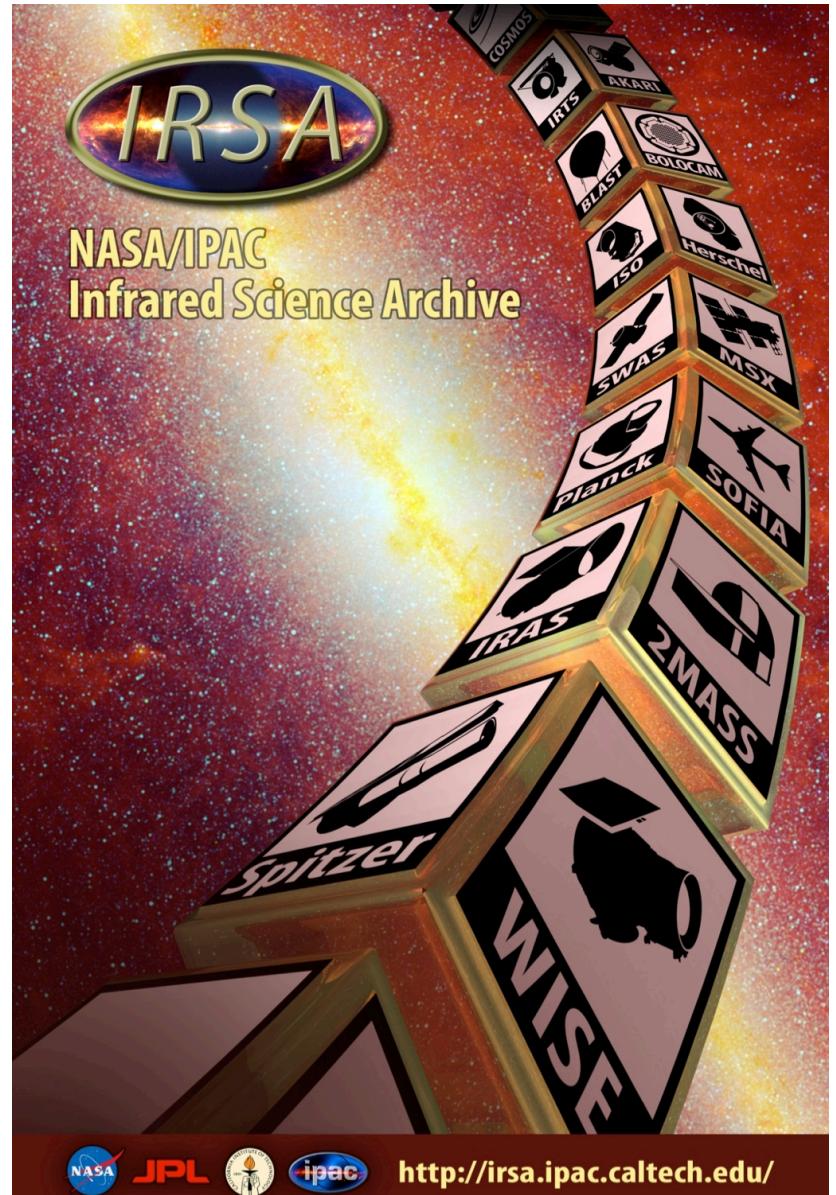
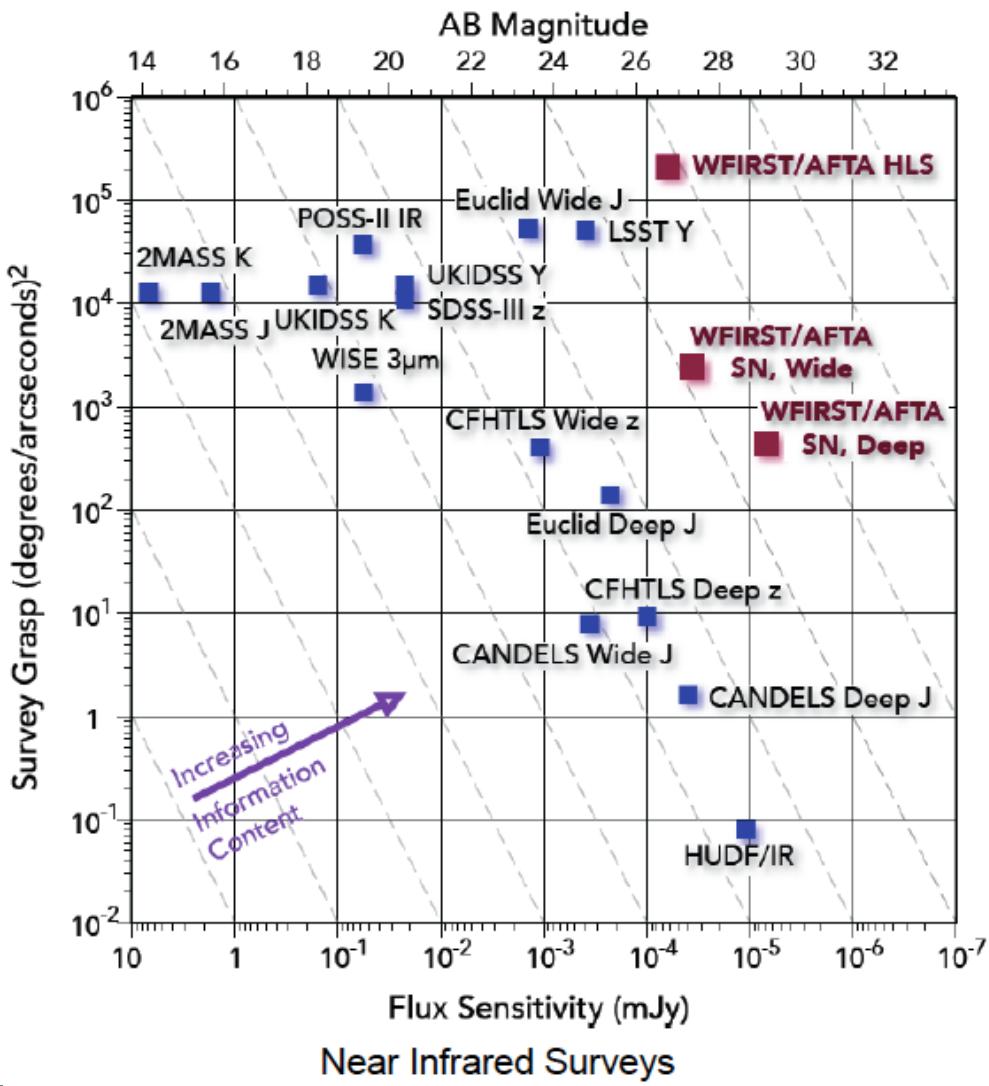


# Community Research Support: Existing IR Surveys

Vandana Desai



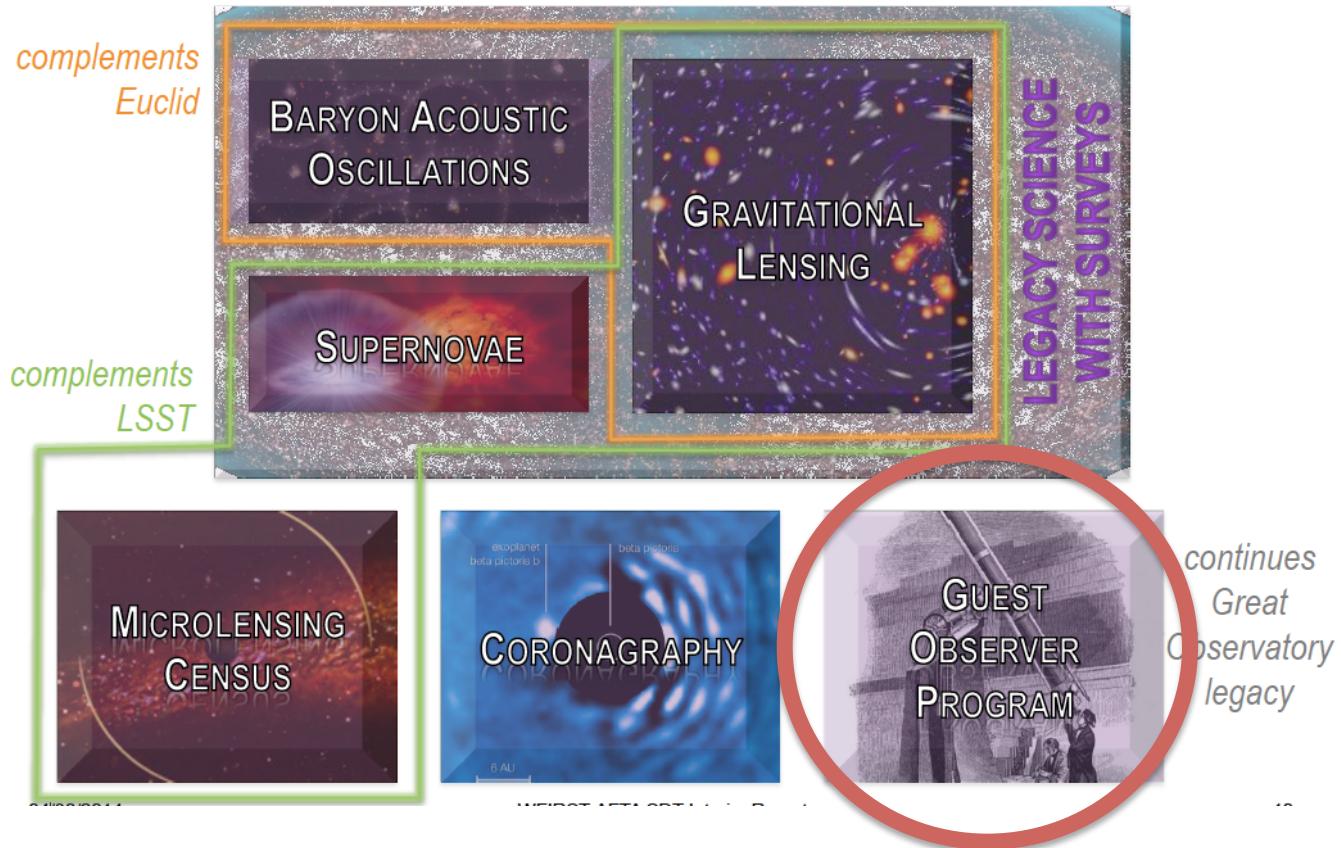


WFIRST will be  
revolutionary.

What is the role  
of existing IR  
surveys?

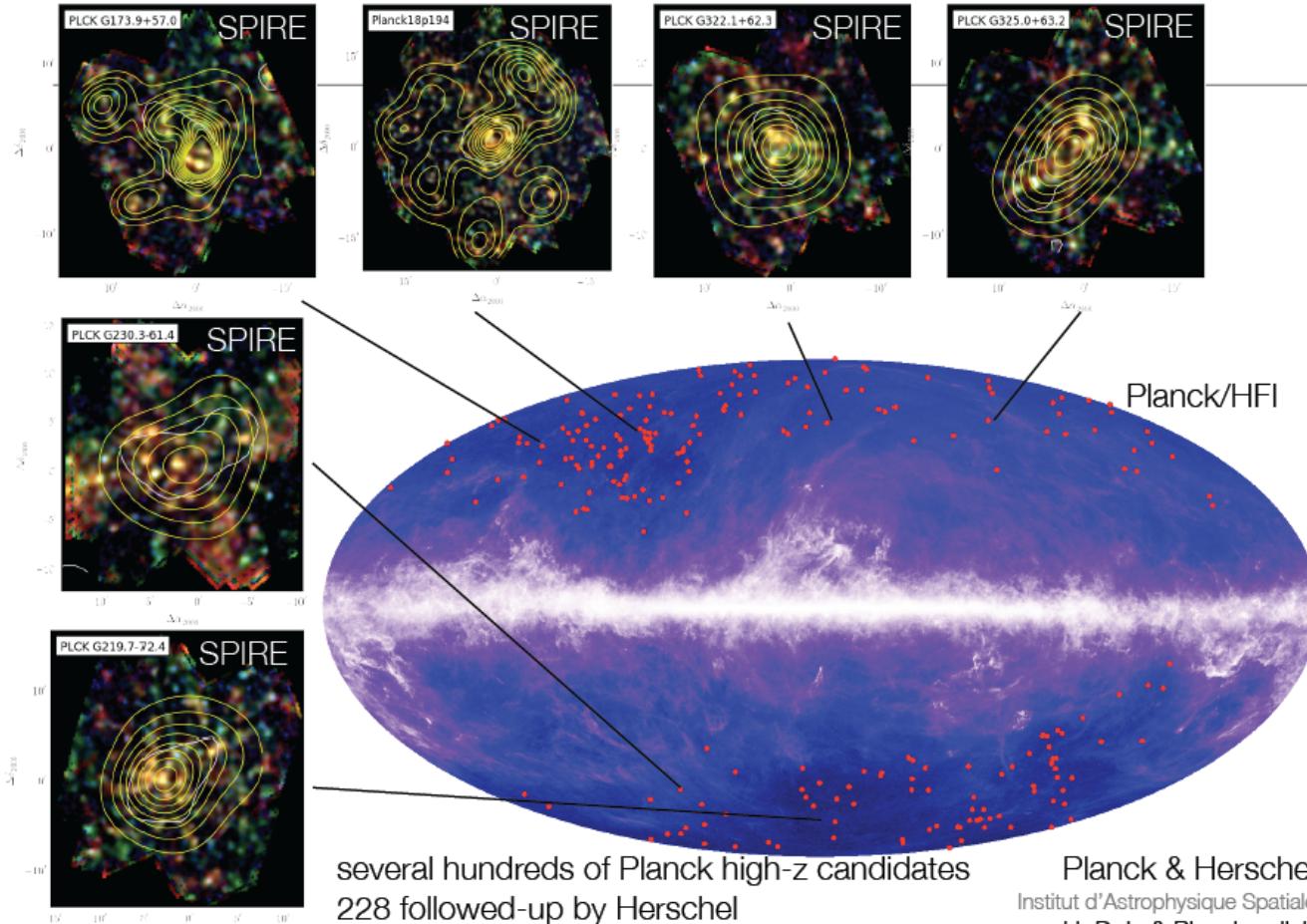


# Existing Surveys will feed WFIRST GO program



- Spitzer is still taking data - Peter Capak's talk
- Harry Ferguson has noted that long wavelength coverage would be a boon to the "for free" beyond the local group science.

# WFIRST GO: $z \sim 2$ clusters selected by star forming galaxies

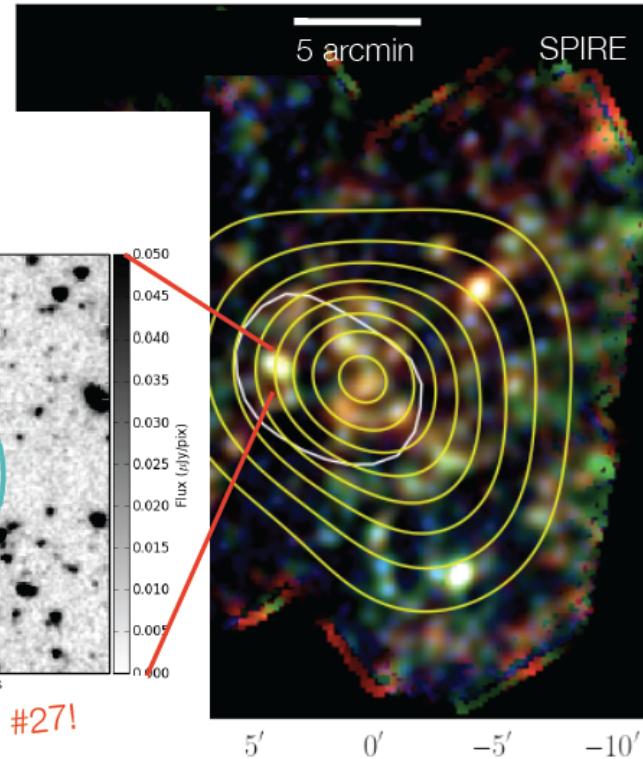
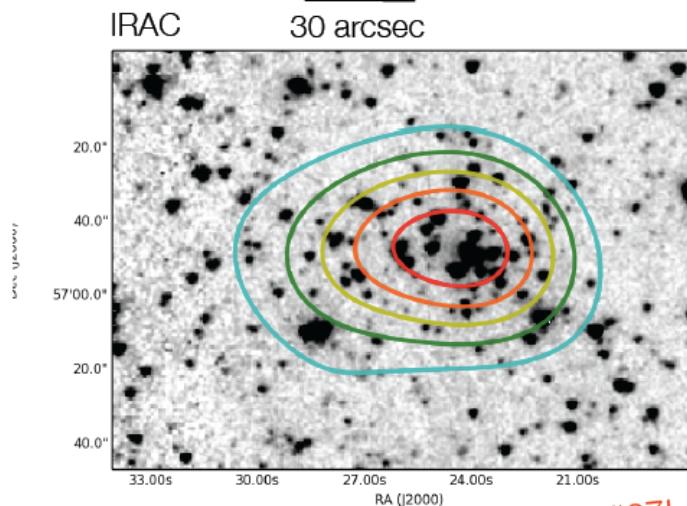


Talk by H. Dole and poster by C. Martinache  
Also related: talk by S. Mei

# WFIRST GO: $z \sim 2$ clusters selected by star forming galaxies

Herschel-SPIRE  
3-color image:  
blue = 250um  
green = 350um  
red = 500um

Euclid will provide this kind of  
sensitivity over the whole sky !  
JWST and WFIRST much  
better, on smaller sky areas !



See Clément Martinache's poster #27!



# WFIRST GO: $z > 1.5$ clusters selected by passive galaxies

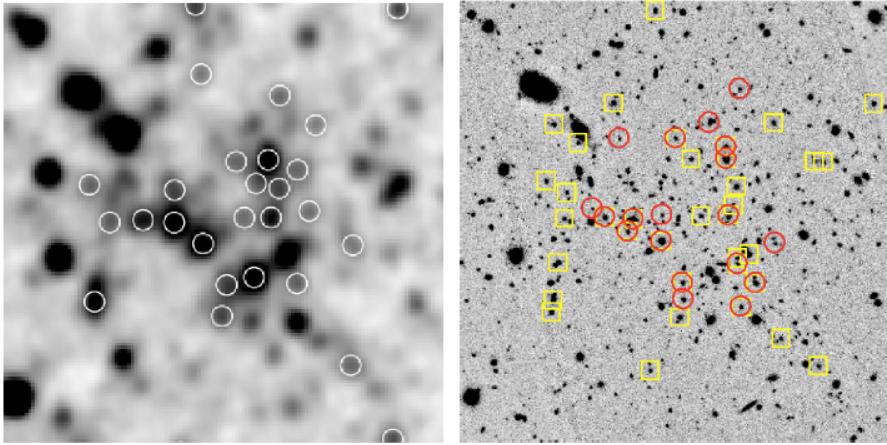
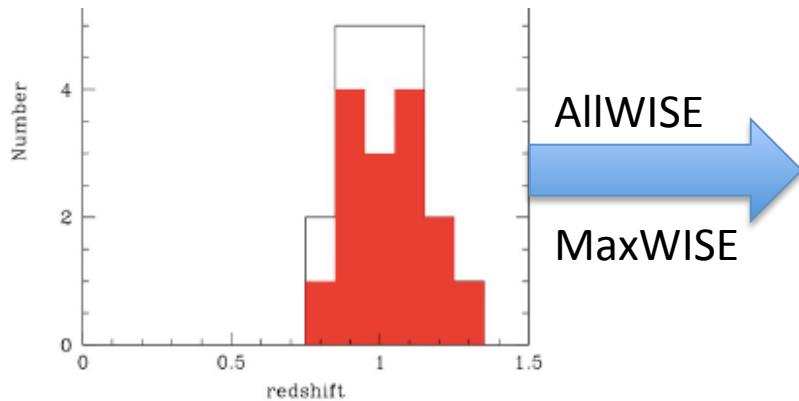


Fig. 1n.— MOO 1514+1346,  $z = 1.06$ , as above; W1 image on the left,  $z$ -band image on the right.



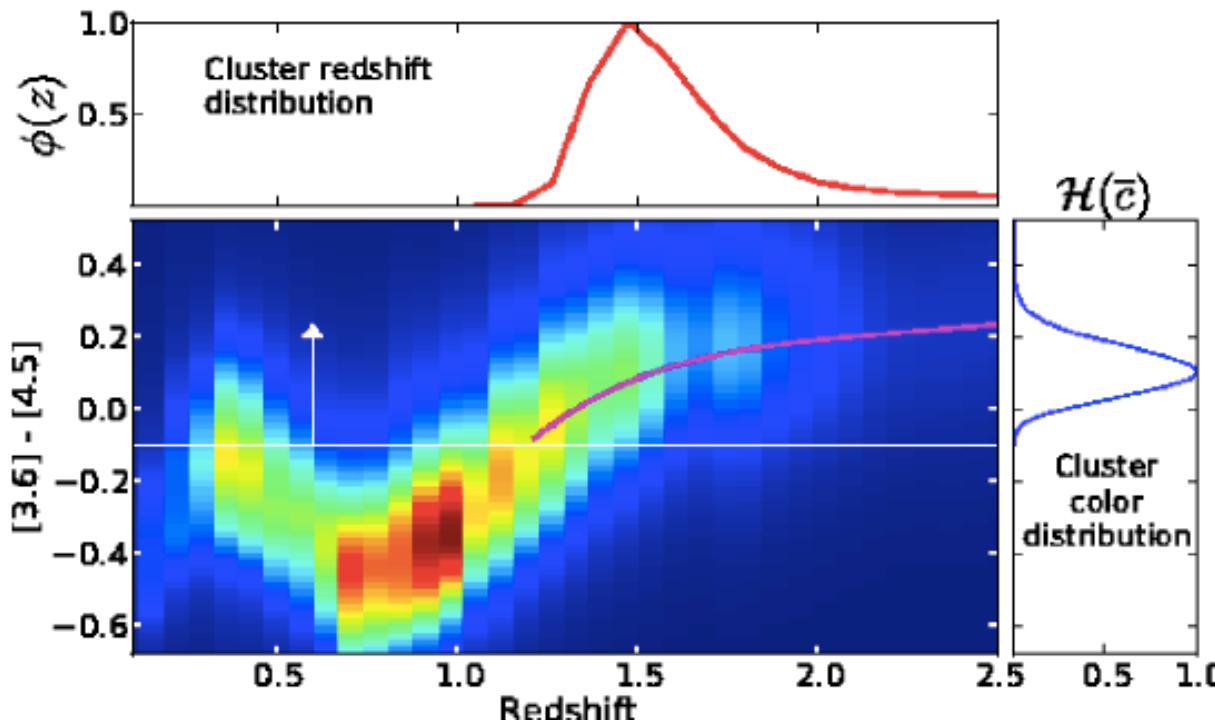
**MaDCOWs:**

- WISE+SDSS
- WISE → AllWISE
- SDSS → DES

Talk by P. Eisenhardt



# WFIRST GO: $z > \sim 1.5$ clusters selected by passive galaxies



Talk by A. Returra

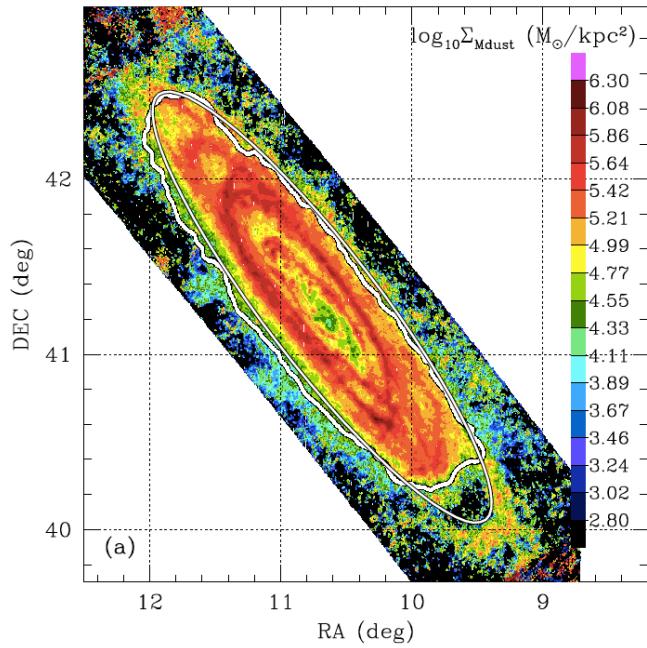


Spitzer Local Volume Survey

# WFIRST GO: Local Galaxies

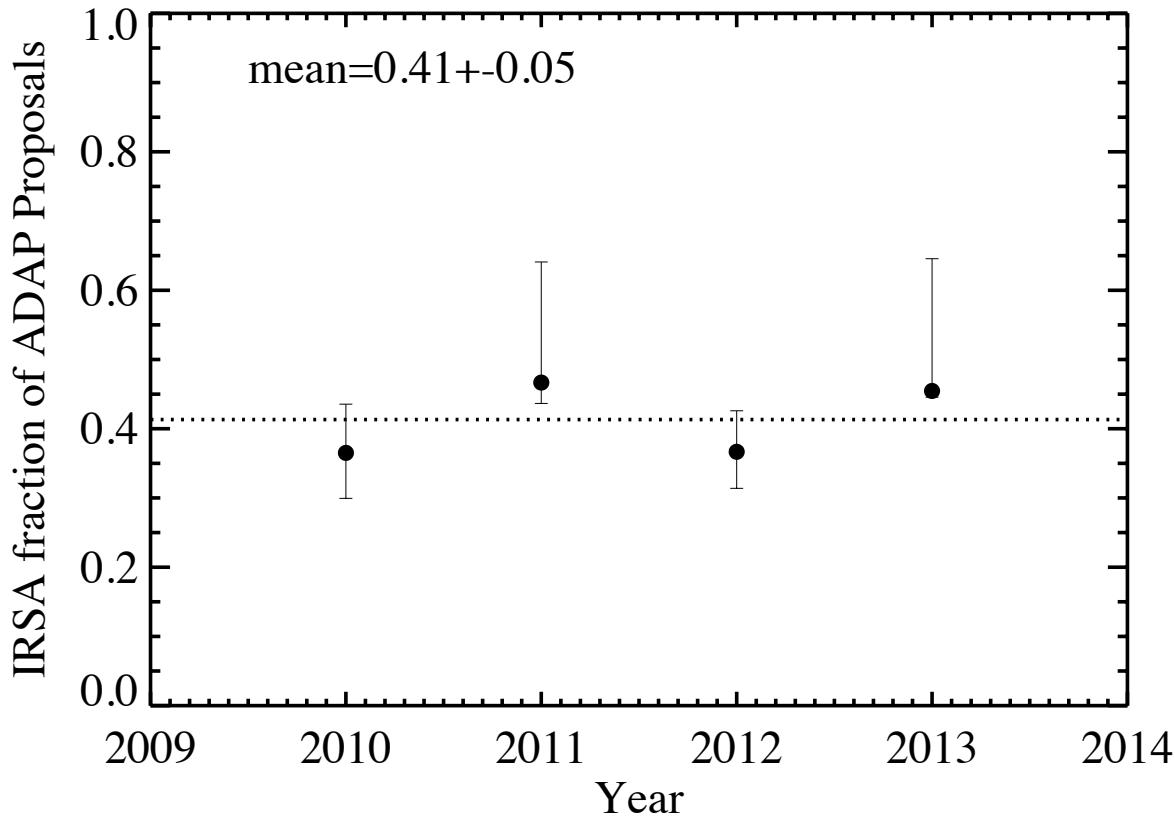
Talk by Dan Weiss

Spitzer, Herschel imaging and spectroscopy provide long wavelength foundation for GO programs.



Draine et al. 2014

# Continued funding for archival research will be important



# Community Needs

- Virtual Observatory
  - Complete metadata
    - Time domain
    - Multiwavelength analyses
  - Supports independent software development
- Grappling with Big Data
  - Visualization of images AND catalogs at the archive
  - Analysis at the archive
  - Personalization
- Supporting time domain/precovery for the future
- Support for Contributed Enhanced Products
- User Support (helpdesk)