

# Young massive evolved stars in the G23.3-0.3 complex

presented by

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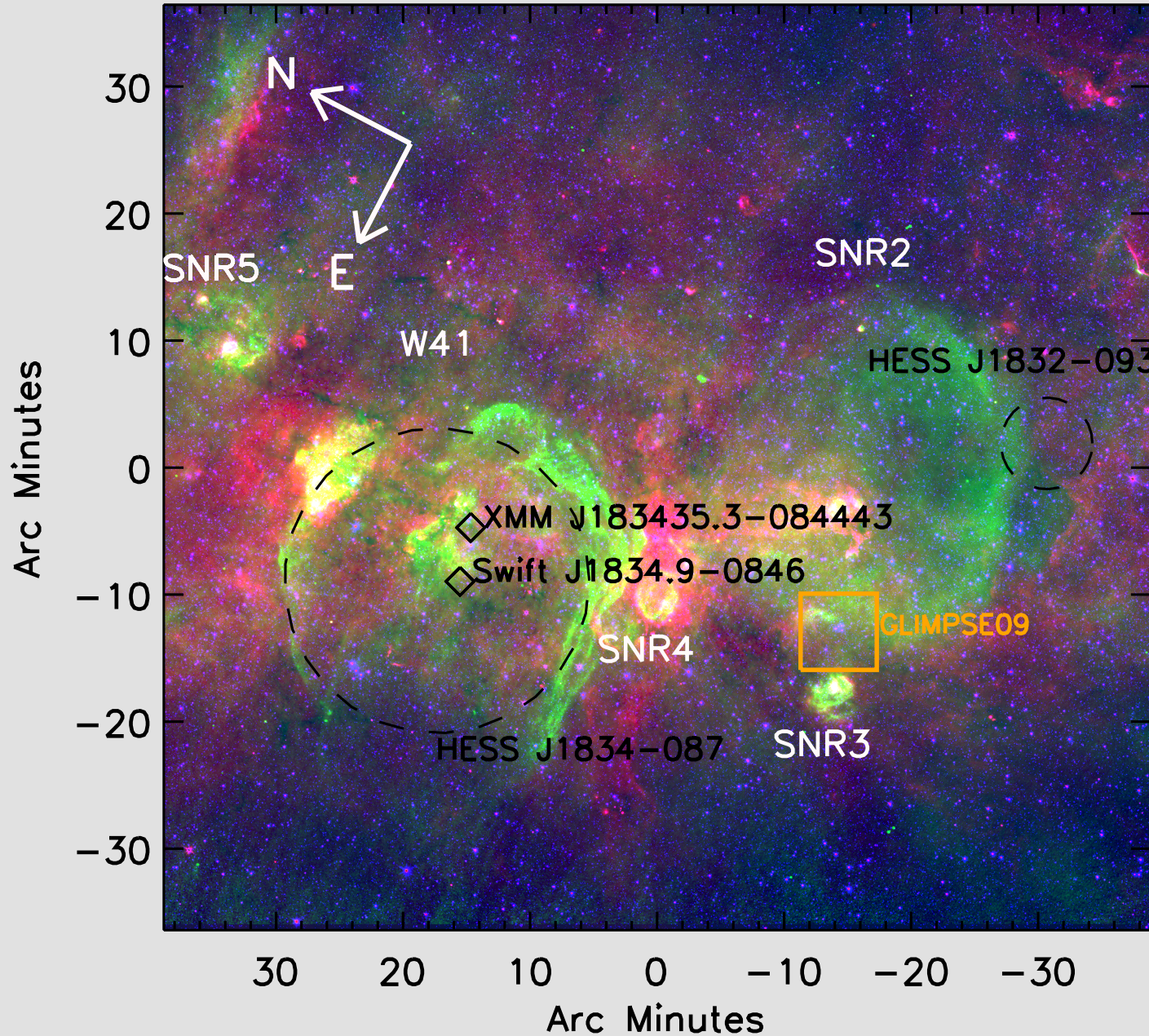
**Pasadena, November 2014**

# Outline:

- A. brief introduction to GMC G23.3-0.3, a beautiful star forming complex
- B. Detected massive stars
- C. Spatial and temporal distribution

Reference: Messineo et al. 2014 A&A 569, 20

# The G23.3-0.3 giant molecular complex



Blue = 3.6 um  
GLIMPSE  
Green = 20 cm  
MAGPIS  
Red = 8 um  
GLIMPSE

SNRs from:  
Green 1991  
Helfand et al. 2006  
1=W41  
2=G22.7-0.2  
3=G22.7583-0.4917  
4=G22.9917-0.3583  
5=G23.5667-0.0333  
Albert et al 2006  
Yang Su et al. 2014

High-energy  
Aharonian et al. 2005  
Laffon et al. 2011  
Mukherjee et al. 2009  
Kargaltsev et al. 2012

Center: Longitude 22.99 Latitude -0.19



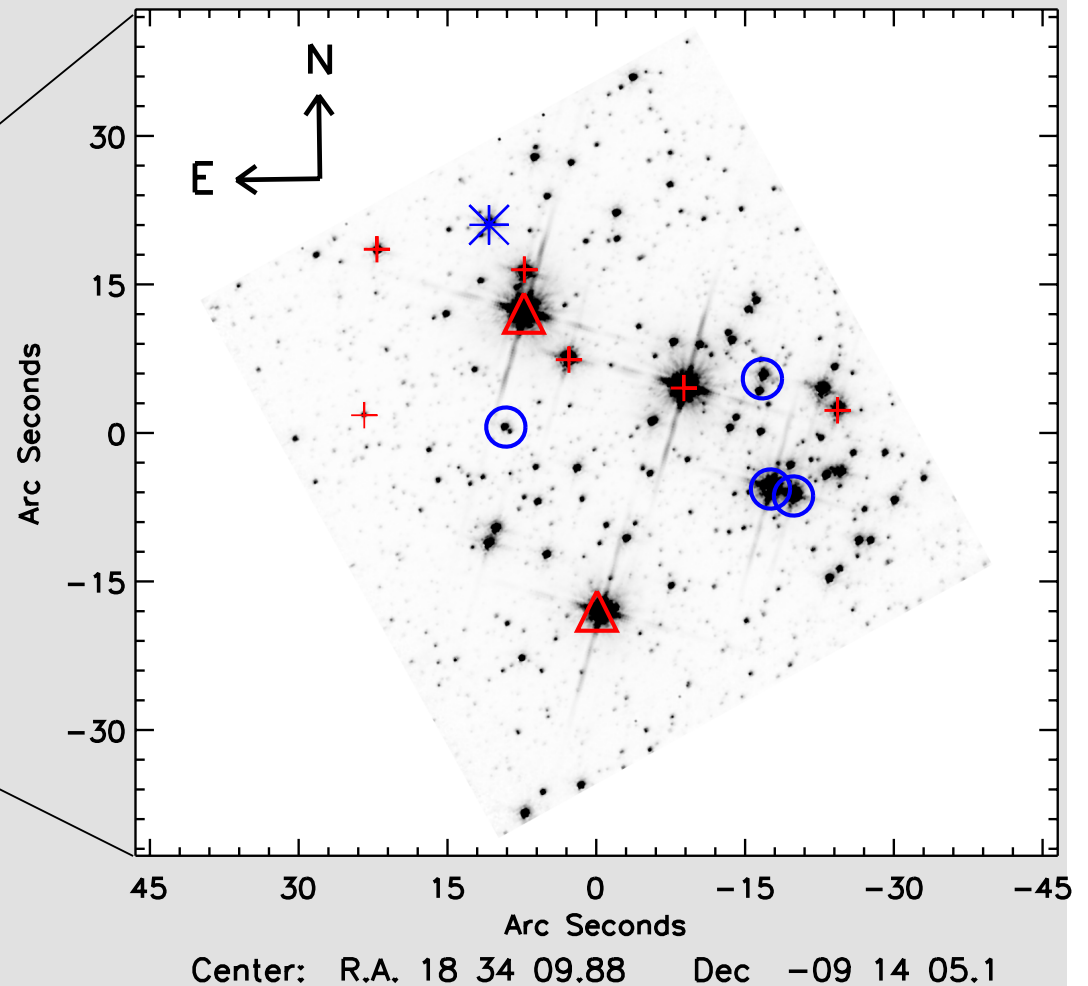
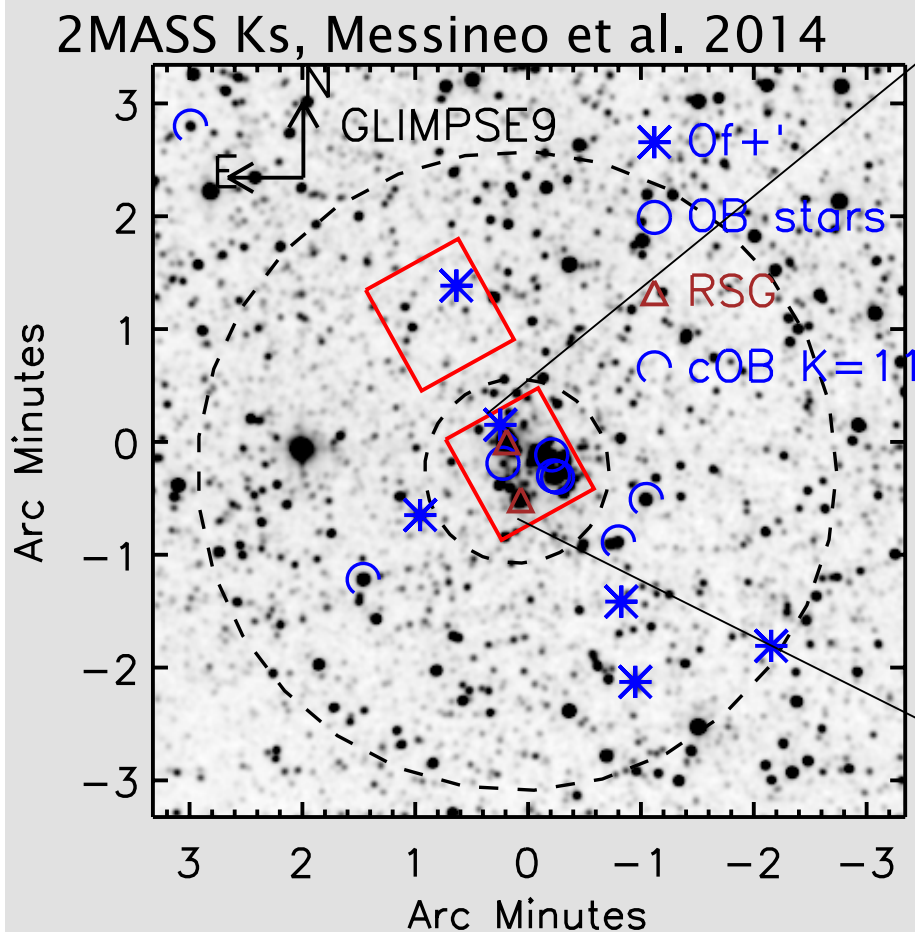
## Spectrophotometric distance

<b>Spec</b>	<b>Nstar</b>	<b>Aks</b>	<b>Mk</b>	<b>DM</b>
O4-6 I	1	1.34	-5.16	13.73
O6-7 I	4	1.73	-5.28	13.48
O7-8.5 I	1	1.62	-5.39	13.65
O9-9.5 I	1	1.68	-5.39	12.89
O6-7 III	3	1.67	-4.84	13.69
O7-8.5 III	1	1.75	-4.66	13.31
O9-9.5 III	1	1.37	-4.47	13.68
B0-3 I	2	1.57	-6.27	13.25

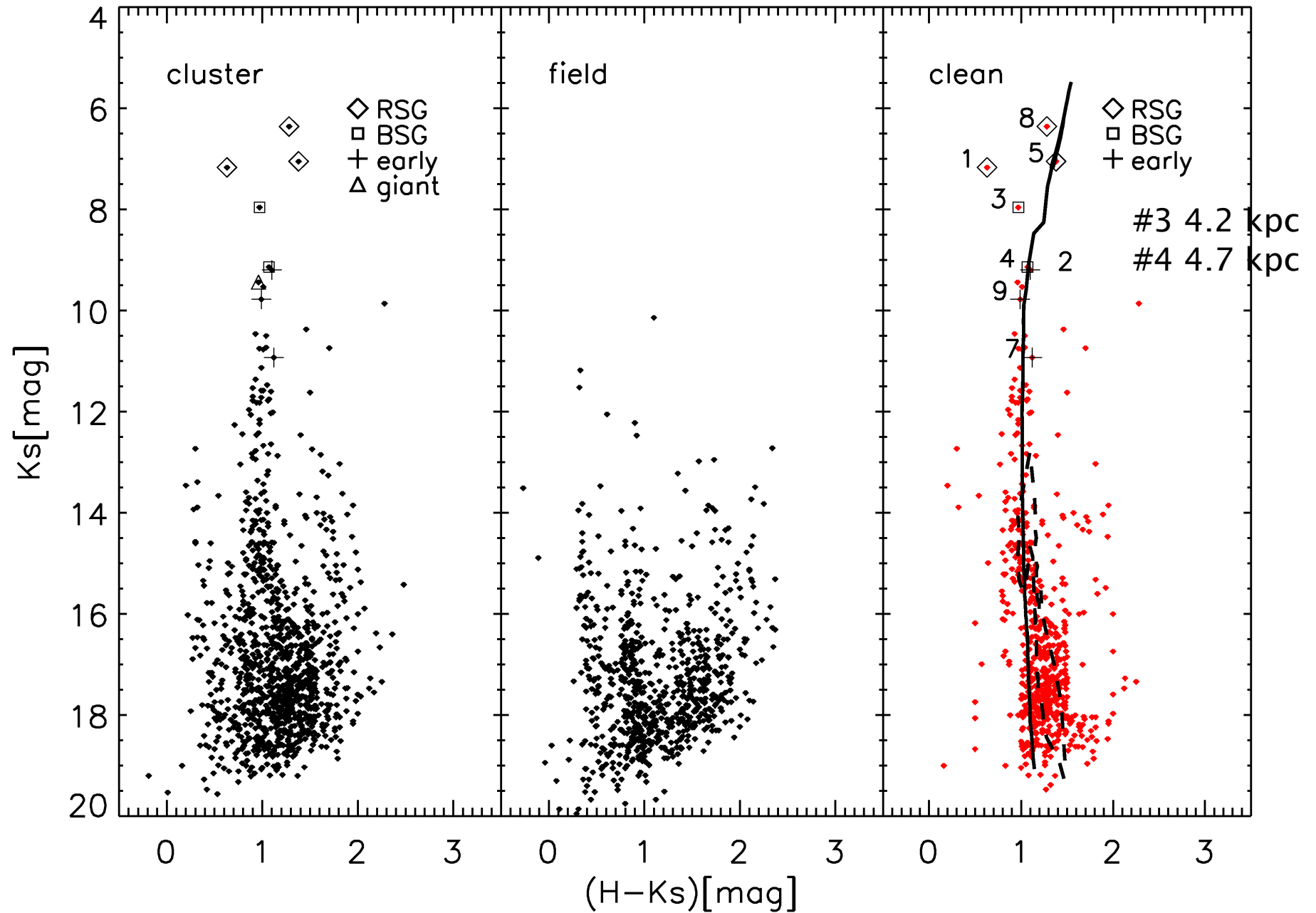
<b>Region/component</b>	<b>Method</b>	<b>DM</b>	<b>References</b>
70.5-82.5	Gas Kinematic	13.19 -- 13.39	Reid et al. 2009 Messineo et al. 2014
G23.01-0.41	Parallax of masers	13.31 pm 0.17	Brunthaler et al. 2009

# GLIMPSE9 and surrounding ( SNR G22.7-0.2 South)

HST/NICMOS F160W, F222M  
f.o.v. = 51.5"x51.5"; pixel scale = 0.2"  
exptime = 19.94s, 55.94s  
(Messineo et al. 2010).

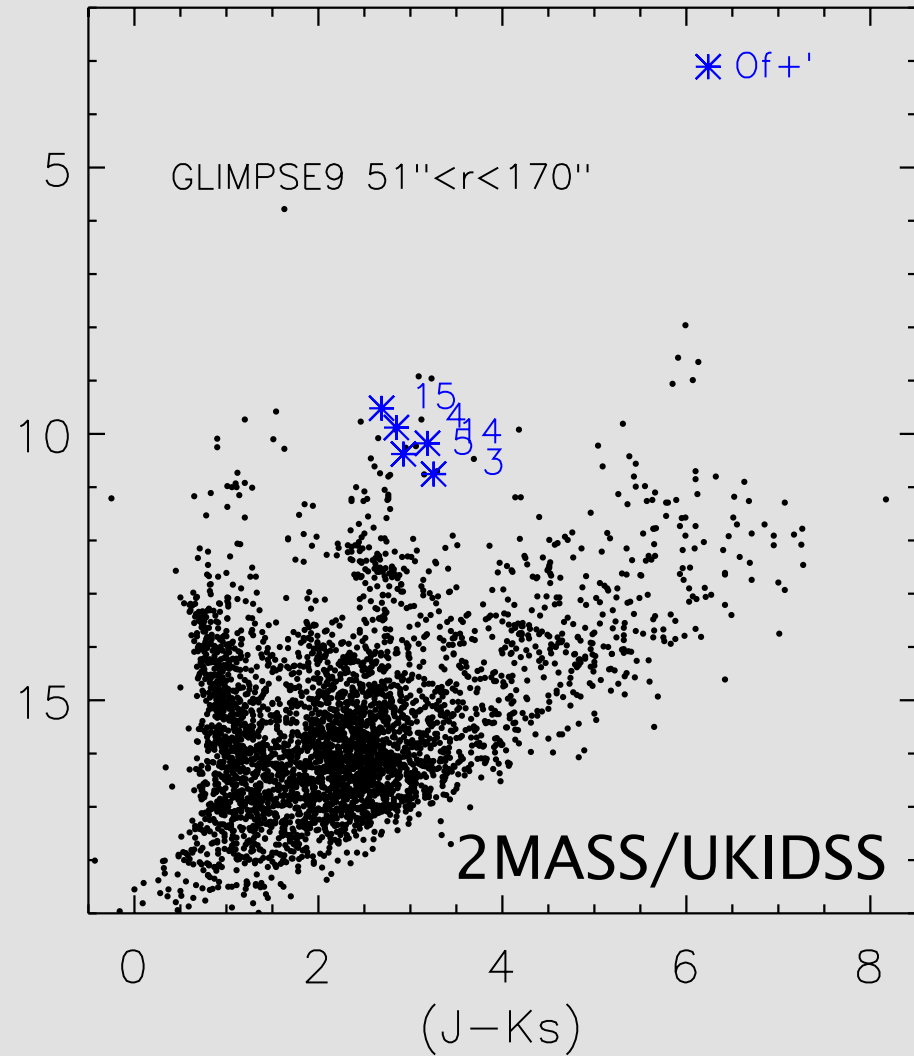
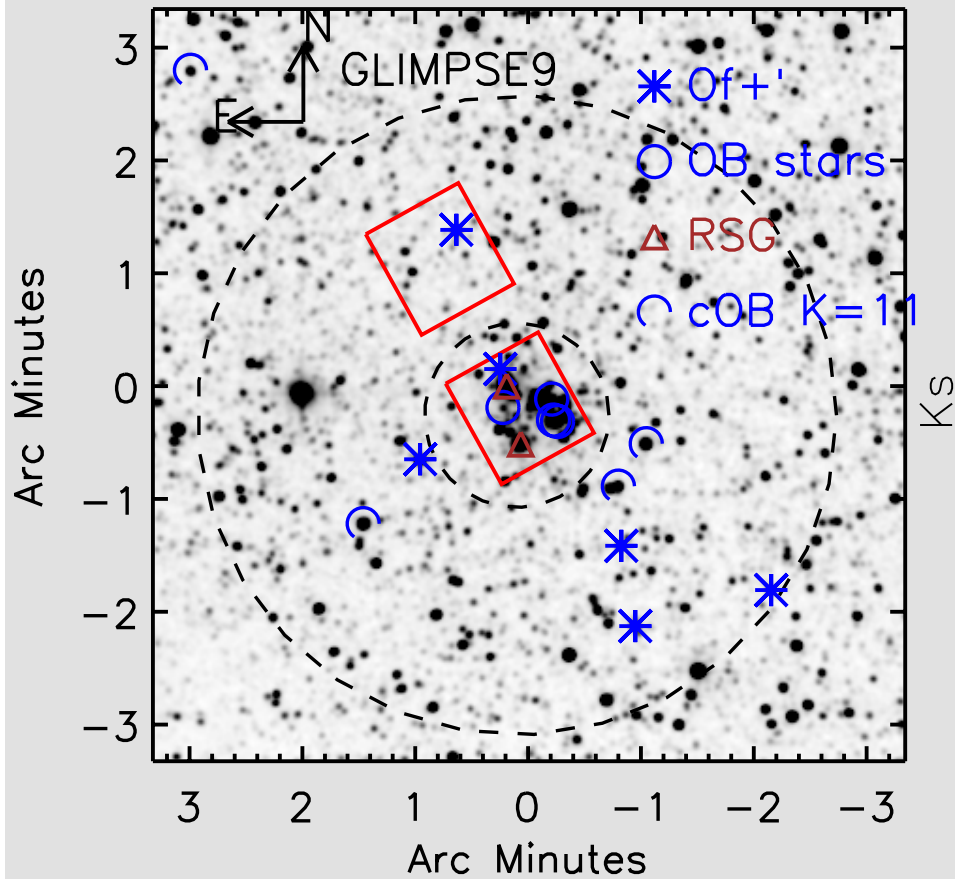


Age = 6-30 Myr (presence of RSGs)  $A_k = 1.6 \pm 0.3$  mag



# GLIMPSE9 and surrounding ( SNR G22.7-0.2 South)

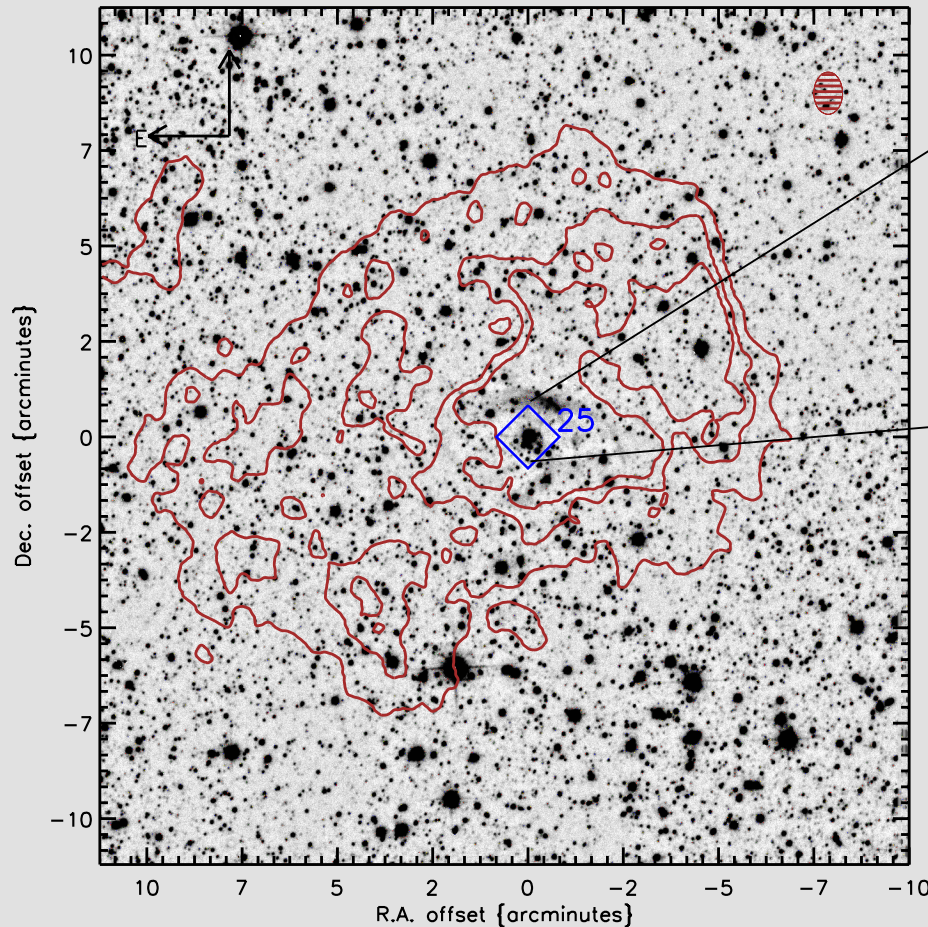
2MASS Ks, Messineo et al. 2014



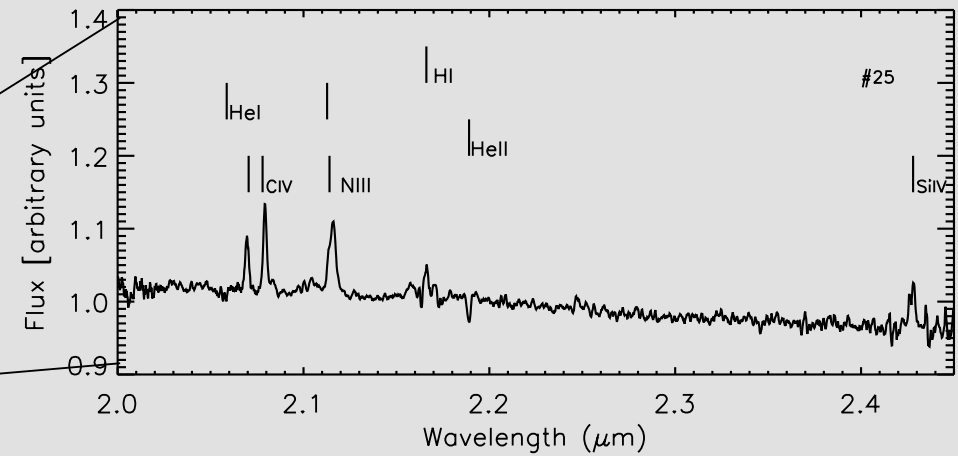


Star #25 in cSNR3-G22.7583-0.4917 (Helfand et al. 2006)  
[BDS2003]117 (Bica et al. 2003)

UKIDSS K, MAGPIS 20cm, 0.2"/pix



ESO-SINFONI



O4-6f+

$T_{\text{eff}} = 38500 \text{ K}$

$K = 9.9 \text{ mag}$

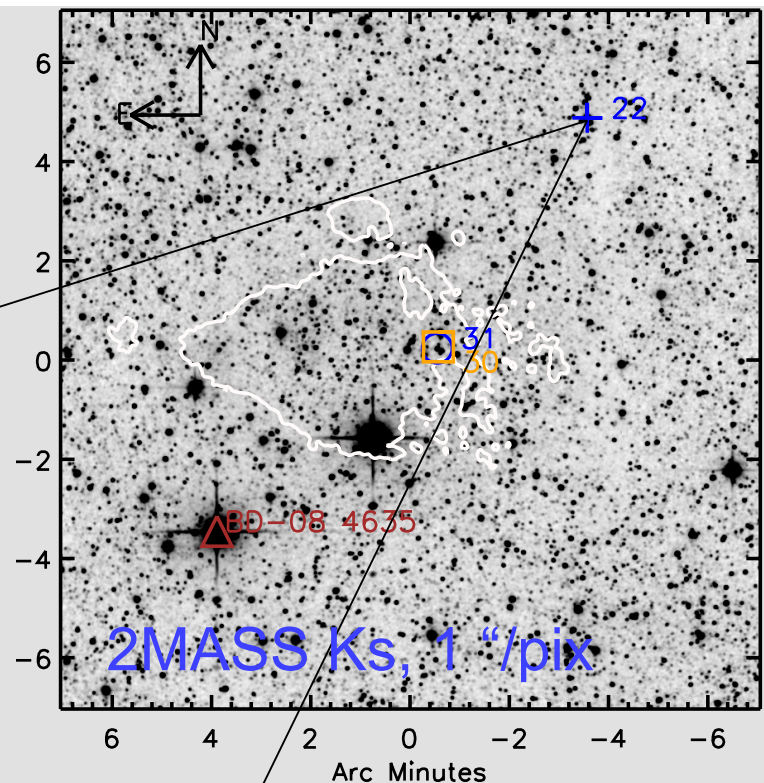
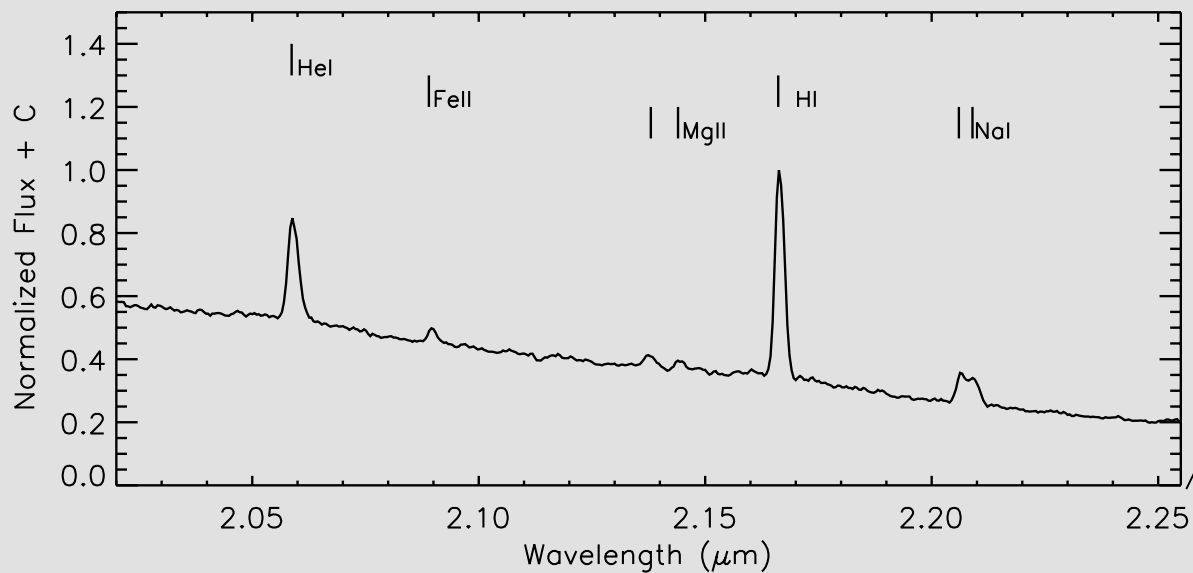
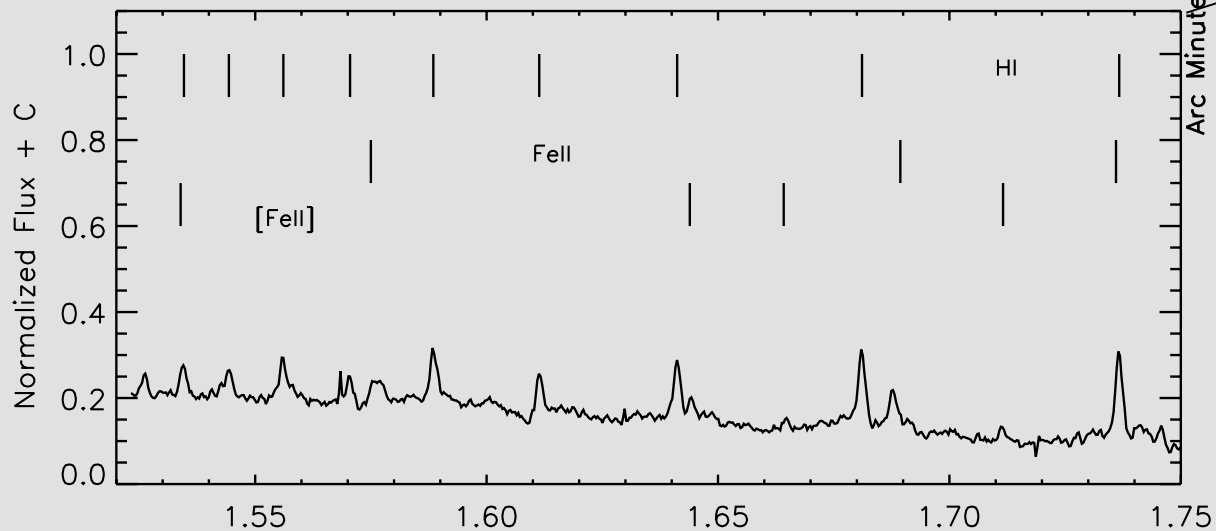
$A_{\text{Ks}} = 1.34 \text{ mag}$

$\text{Dist} = 4.6 \text{ kpc}$

$M_{\text{bol}} = -9.14 \text{ mag}$

$\log L_{\text{um}} = 5.55 \text{ mag} (28\text{-}36 \text{ Msun})$

# B supergiant, cLBV, ESO-SofI



$T_{\text{eff}} = \sim 14000\text{K}$   
 $K_s = 7.63 \text{ mag}$   
 $A_{K_s} = 1.13 \text{ mag}$   
 $\text{Dist} = \sim 4.6 \text{ pm kpc}$   
 $M_{\text{bol}} = -7.90 \text{ mag}$   
 $\text{Log } L/L_{\text{sun}} = 5.06$

## Summary: G23.3-0.3

- The combination of radio and infrared data allow us to detect the parental GMC, which appears rich in HII regions and SNRs.
- Detected eleven massive Of+ stars with ages of 5-8 Myr.  
Detected ten cRSGs with ages of 15-30 Myr.  
G23.3-0.3 has had one of the longest history of star formation known.
- A **cLBV** is found isolated ( we need a **coronagraph for Ks=7 mag**).
- Massive stars are detected in the core of SNR W41, of SNR G22.7-0.2, and of G22.7583-0.4917.
- The **GLIMPSE9** region on the GMC-SNR G22.7-0.2 interaction shows several massive stars (25-50 Msun) sparse on a 8pc wide area (**IMF of a large area. Proper motions**)
- With similar studies of other clusters and giant HII regions we will be able to shed light on the initial masses of the supernova progenitors, and therefore on the fate of massive stars.

# Galactic location of the G23.3-0.3 complex

