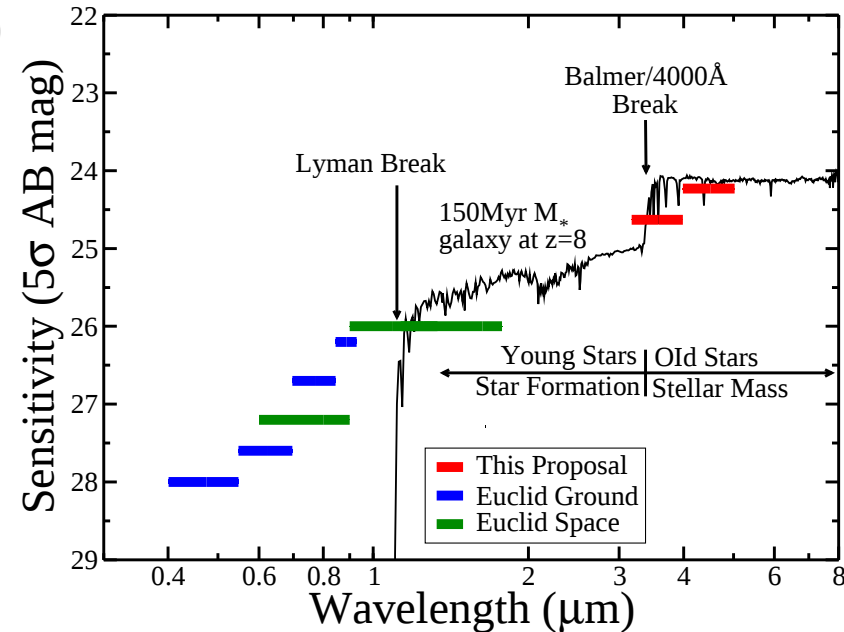


Questions from the Community

- How about a Spitzer 3.6-4.5 μ m survey to determine stellar masses at $z>3$? (few tens of sq. deg.; expensive alternative is JWST) → need wide-area capabilities (radio, FUV, optical)
- Will we have access to the necessary software for grappling with the data?
- What about supporting an open-source-like repository for science (not instrument) software?
- What will be the latency between downlink and distribution? e.g., microlensing and SNe
- Will it be allowed to revisit places already observed in HLS? e.g., transients
- Will the GO 25% be available every year, or is that just an overall mission %? How can we best optimize synergy with other facilities? e.g., JWST would overlap at beginning.



Questions from the Community

- **Combining data with data from other observatories and missions:
Detections assessed at the pixel level, or will we simply be given separate catalogs?
Issues: de-blending, drop-outs, computation ease, assumption of priors**
- **How will archival work be funded? How about an ADAP-like budget purely for HLS data mining? How about making the first archival CFP at “Launch – 2 years”?**
- **Given ~2 intellectual cycles over ~6-year mission, how can we best prepare for Data Day 1? → Provide early on: simulated data, science software, ...**
- **Will there be an effort to simulate what will be seen by the various instruments, in order to optimize instrument development and survey definition?**