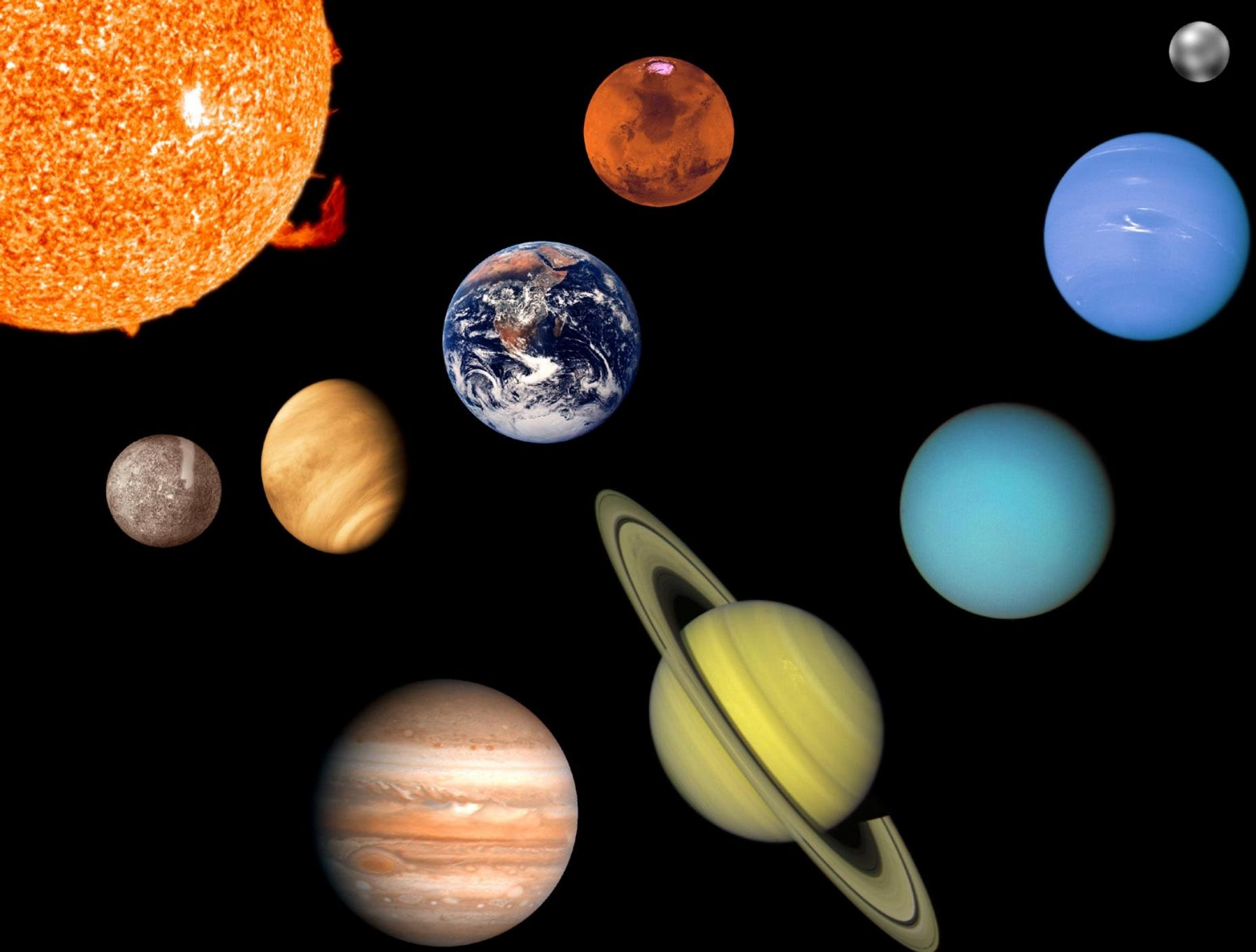




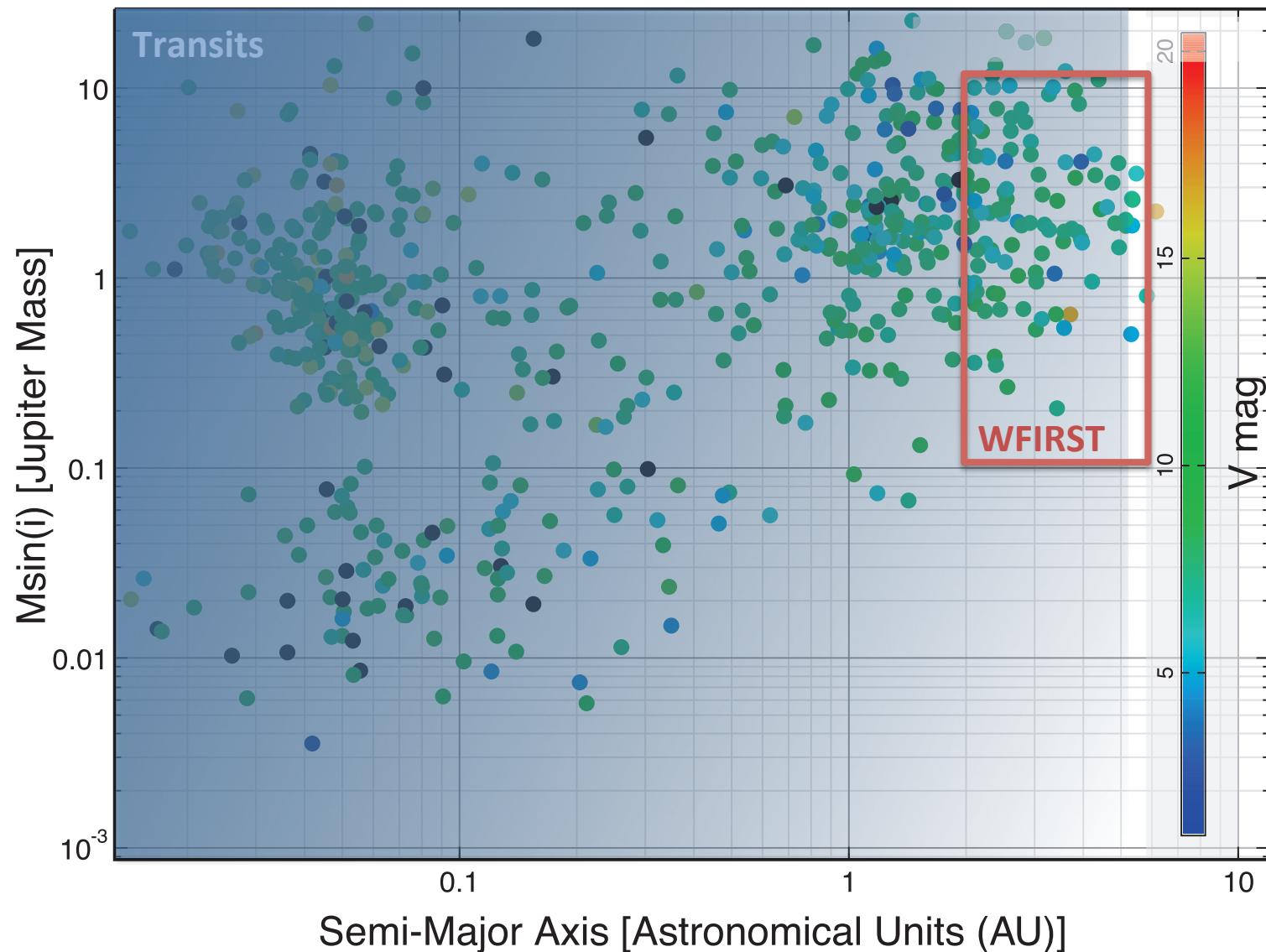
Exoplanet Spectra with WFIRST Cool Planets, Exciting Sciences

Renyu Hu
Jet Propulsion Laboratory

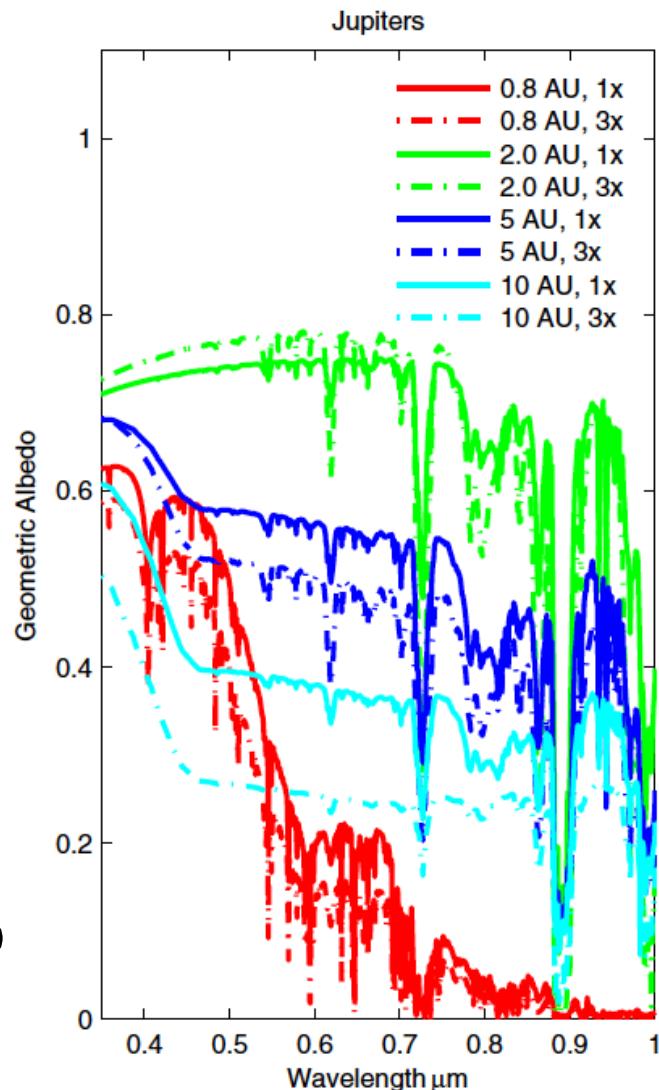
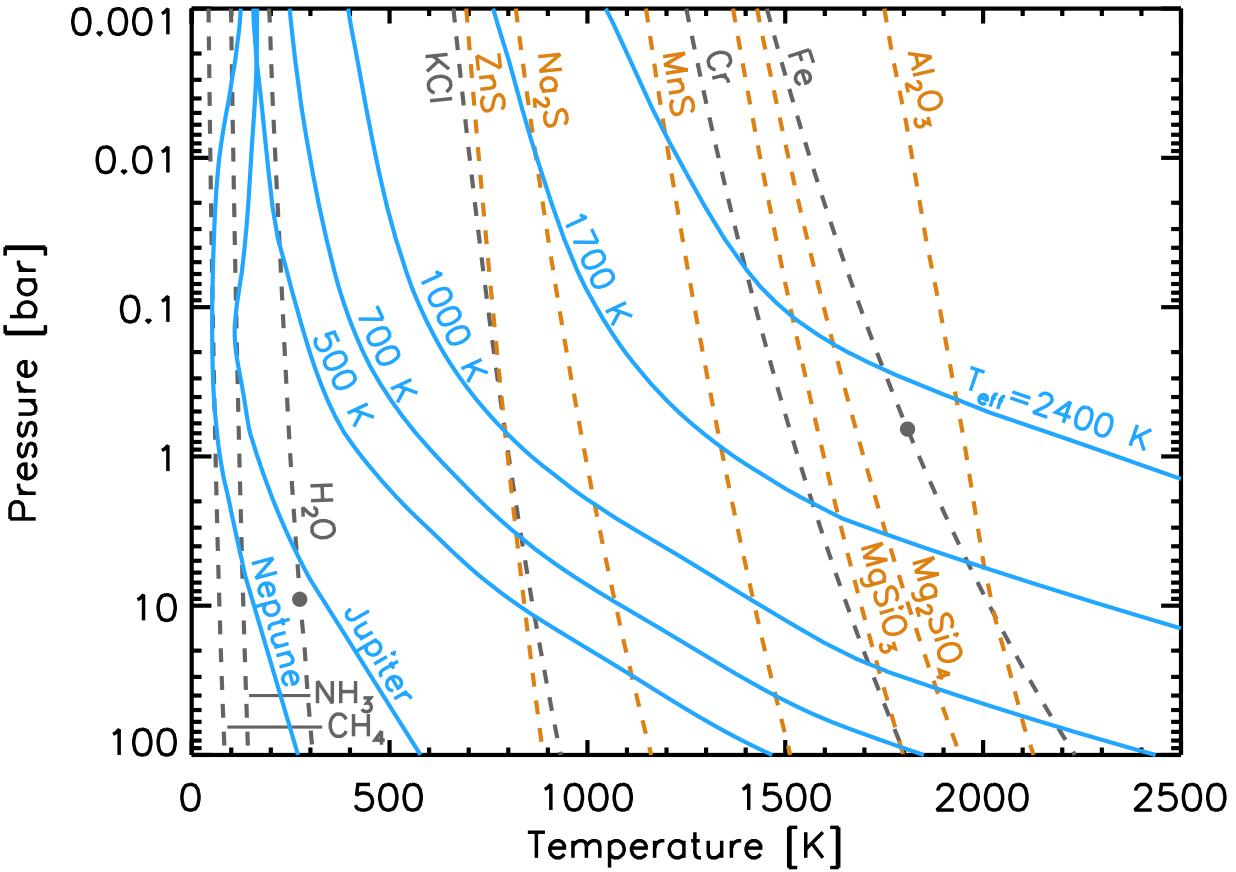
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Government sponsorship acknowledged.



Cool, Gaseous Planets

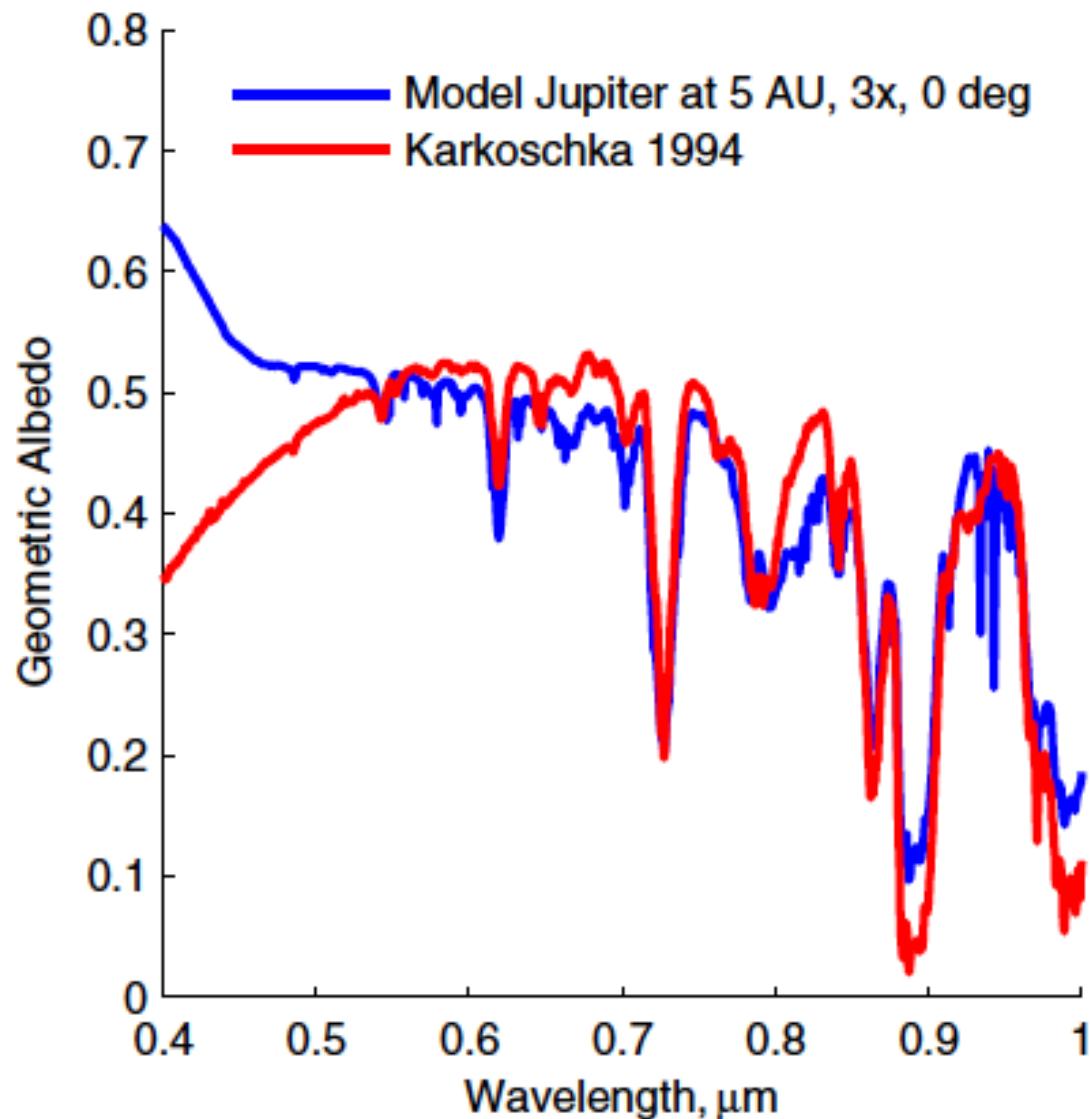


Reflection from Gas and Clouds

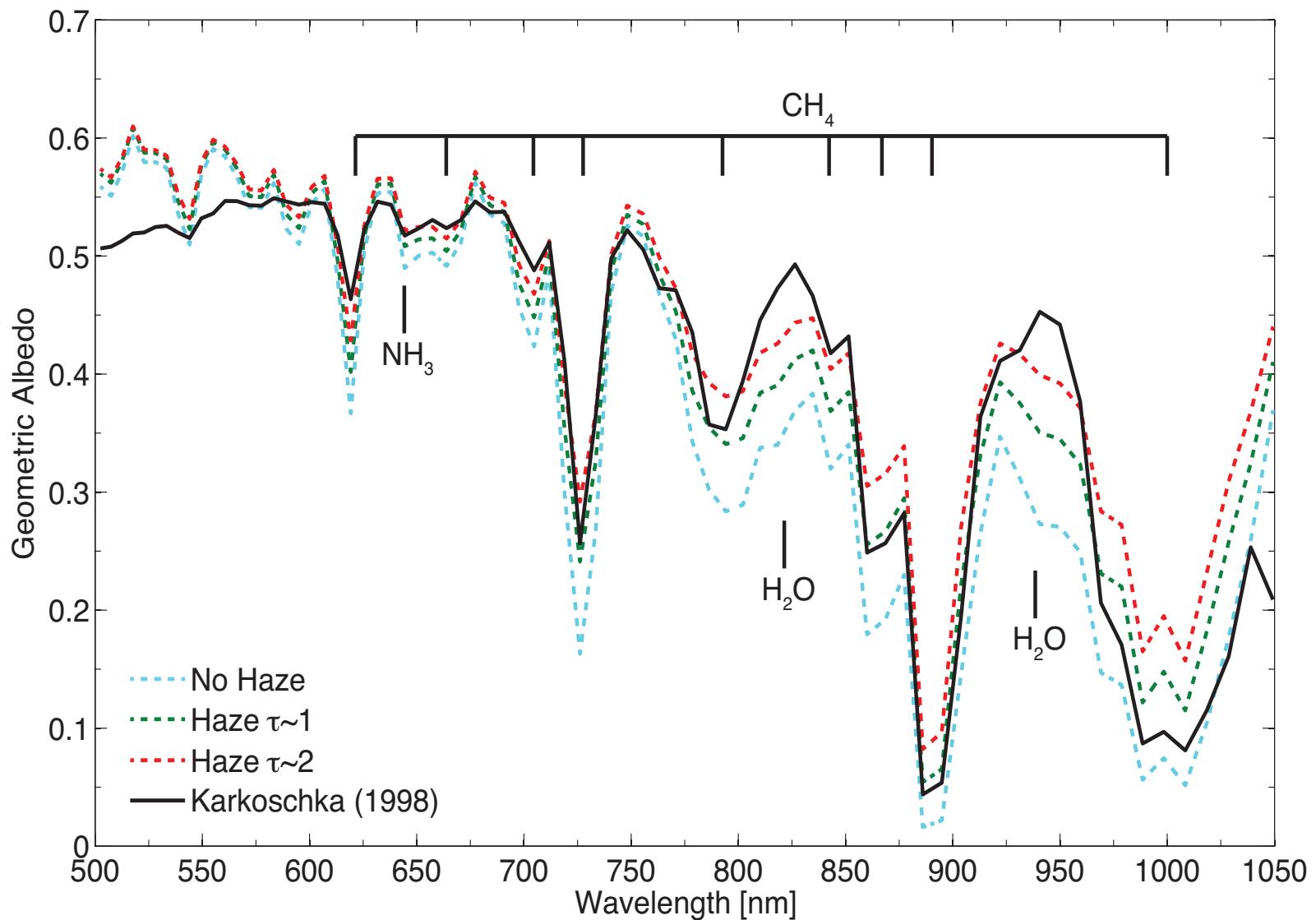


Burrows et al. 1997; Sudarsky et al. 2000;
Cahoy et al. 2010; Marley et al. 2014

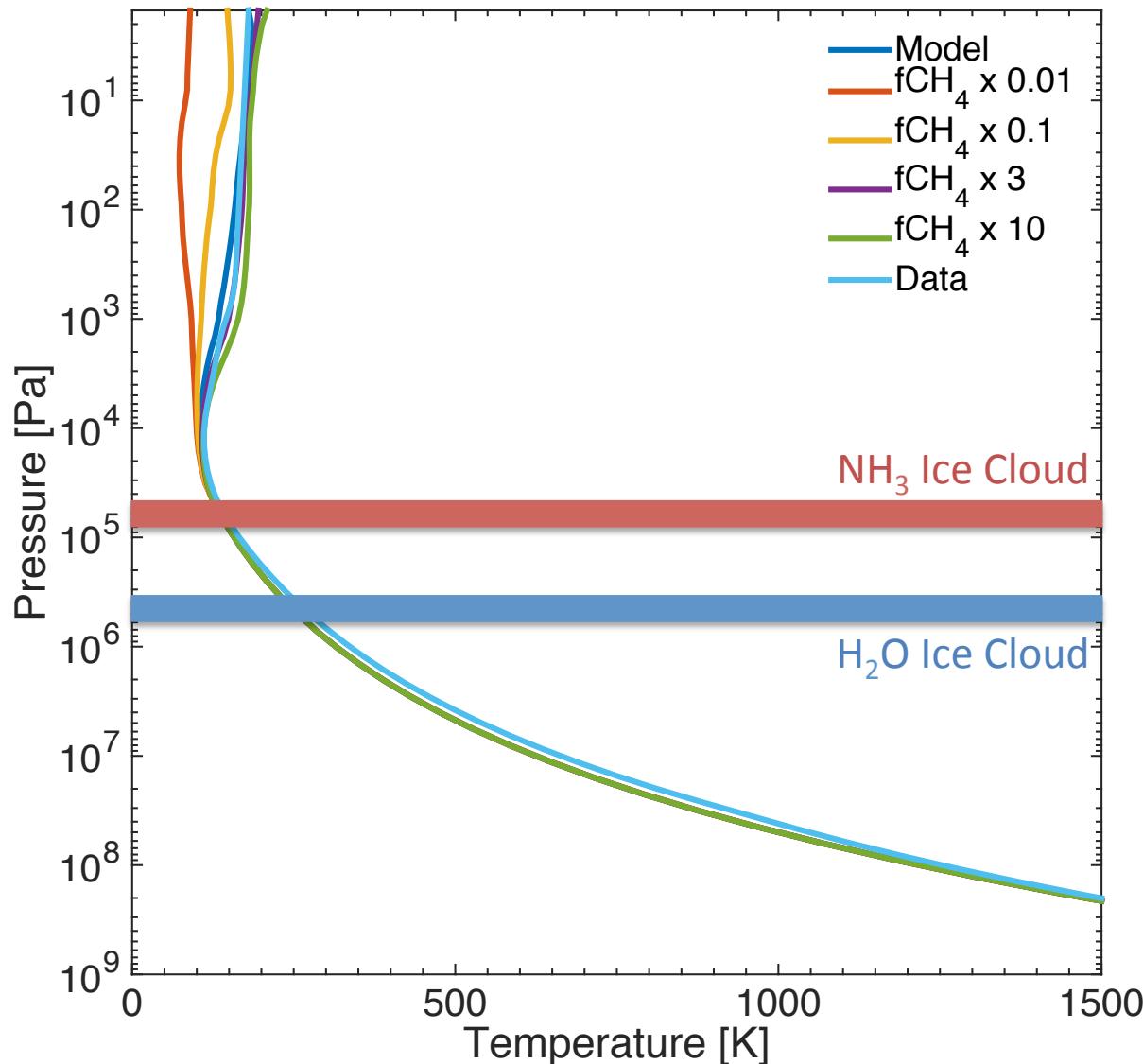
Photochemical Haze



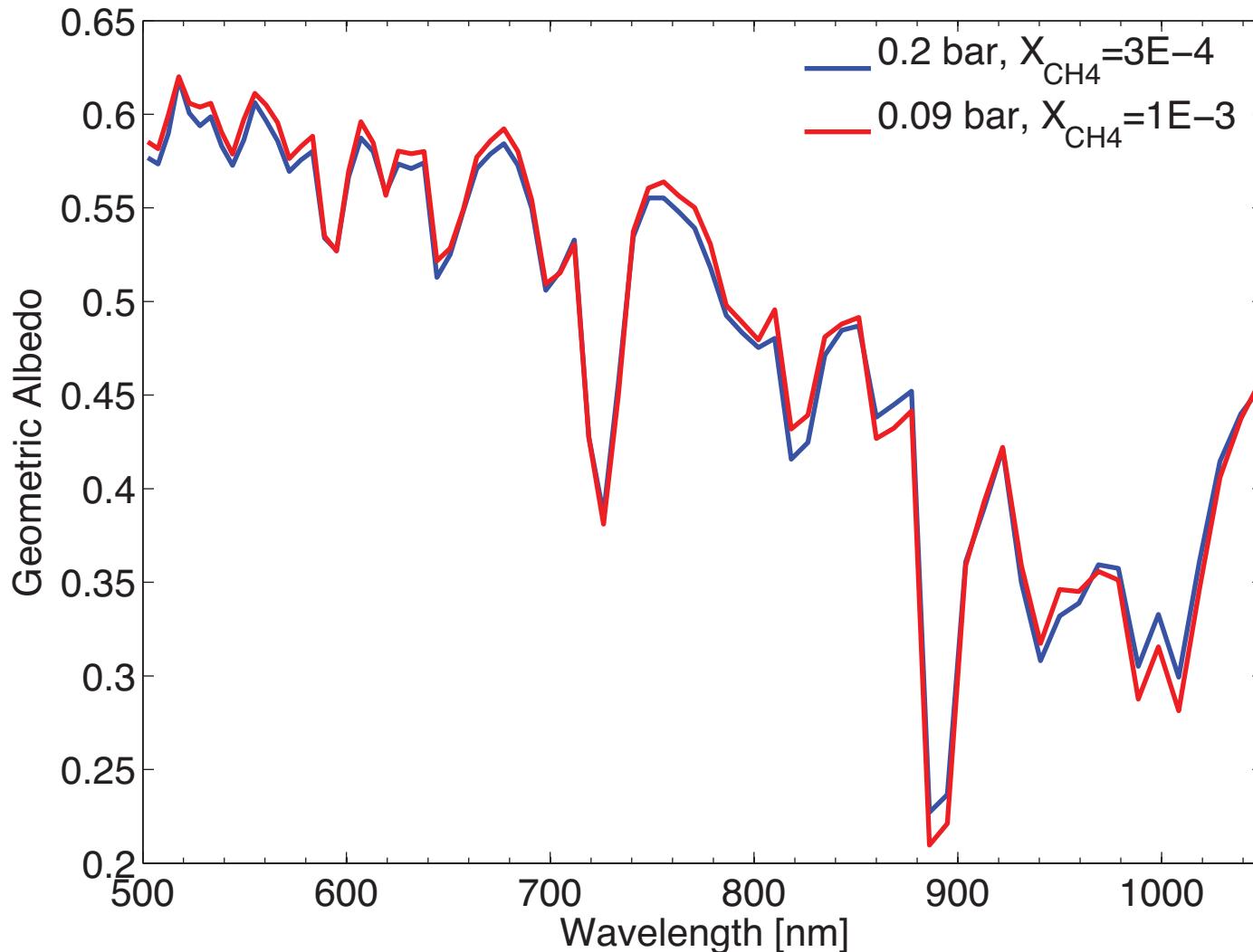
Photochemical Haze



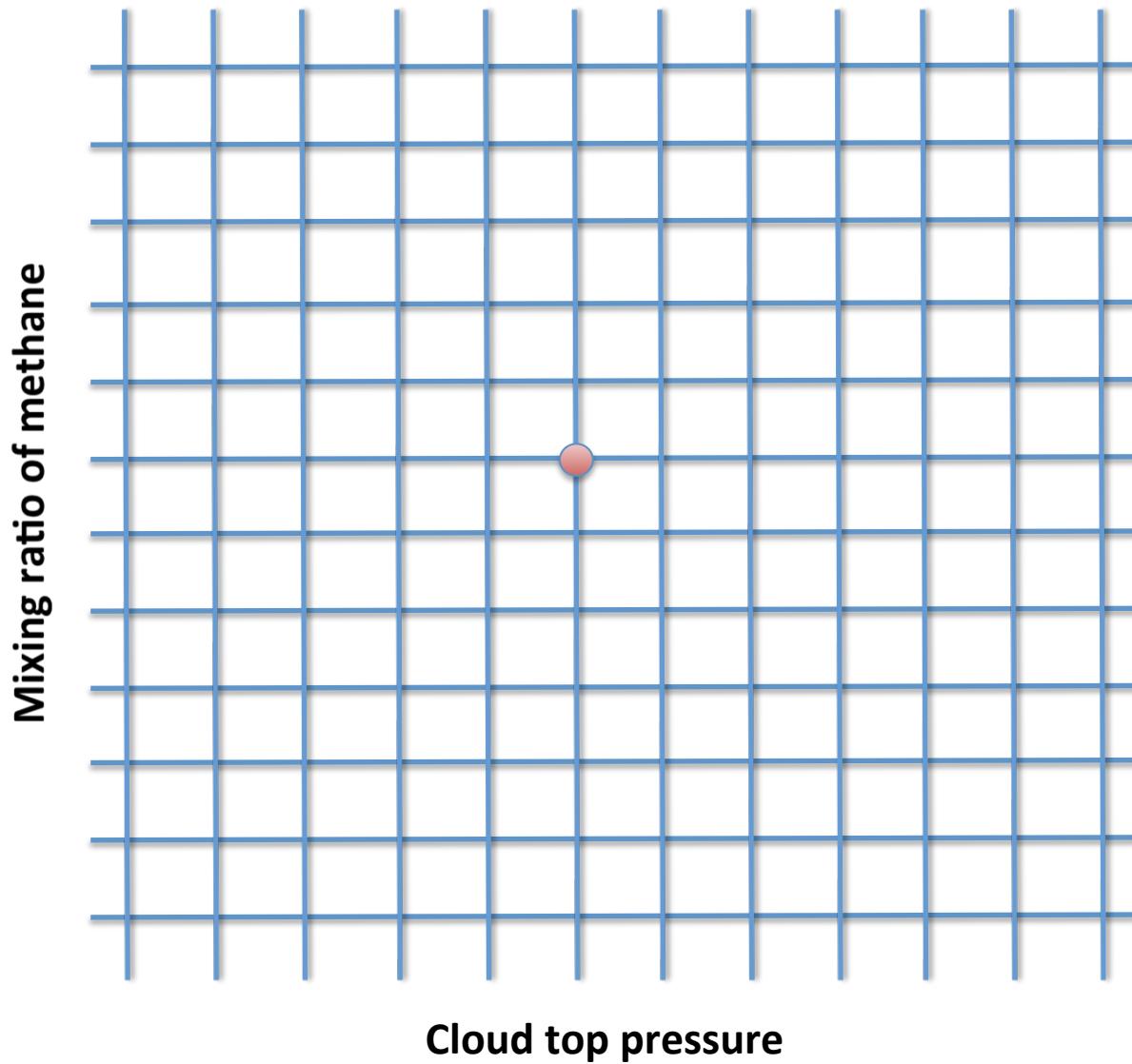
Beyond the Solar Composition



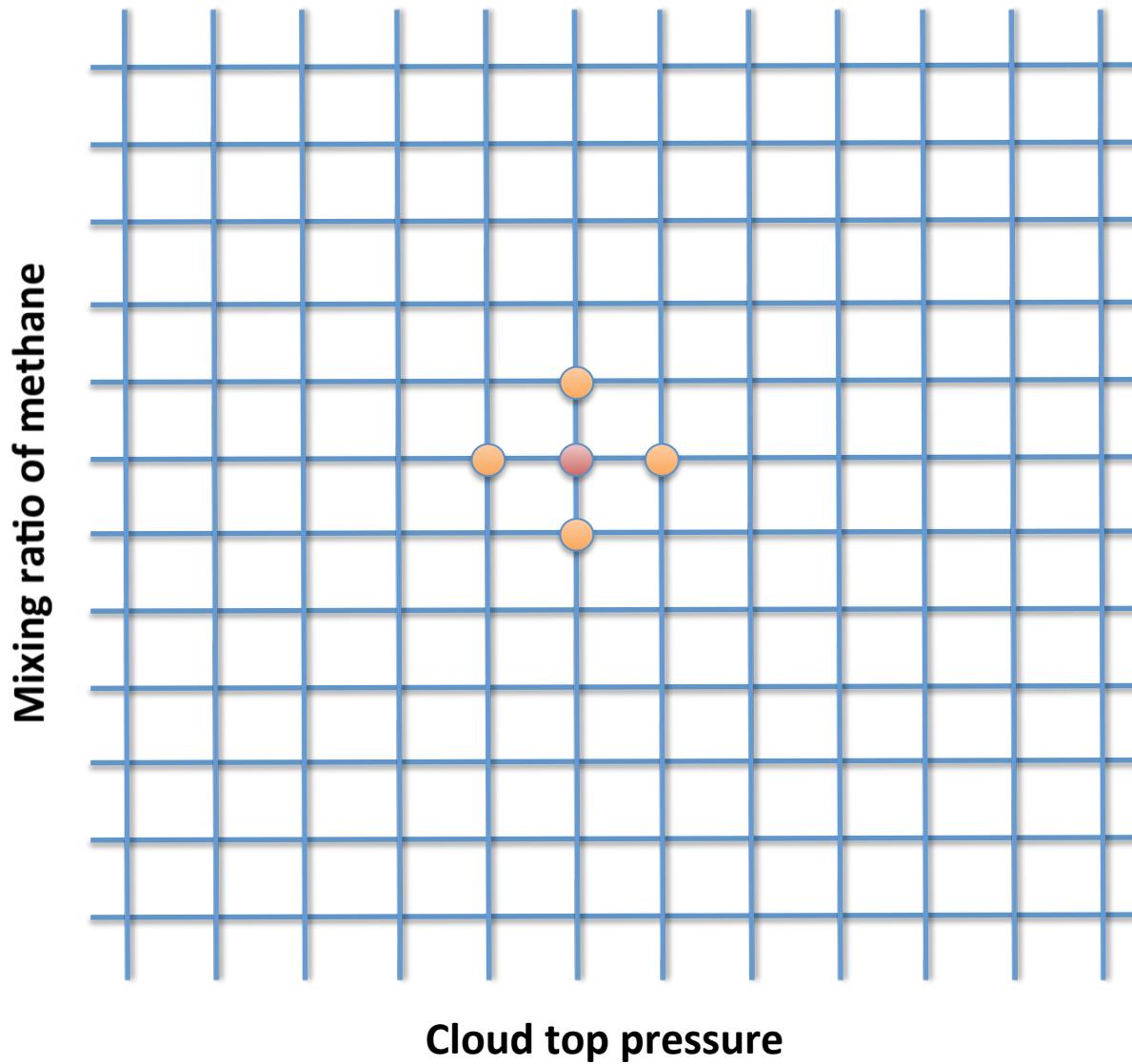
Degeneracy!



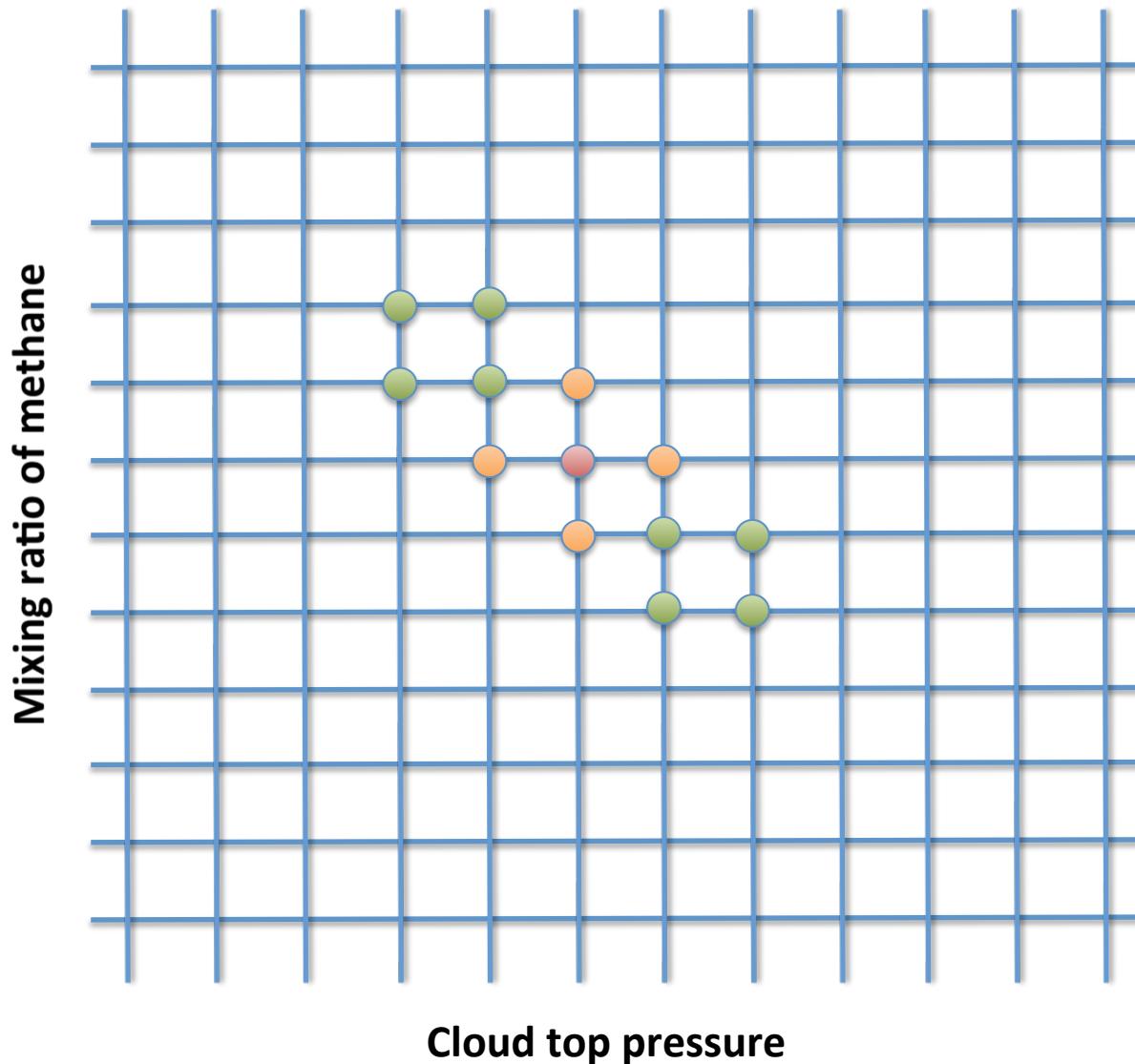
Detectability of Cloud and Gas



Detectability of Cloud and Gas

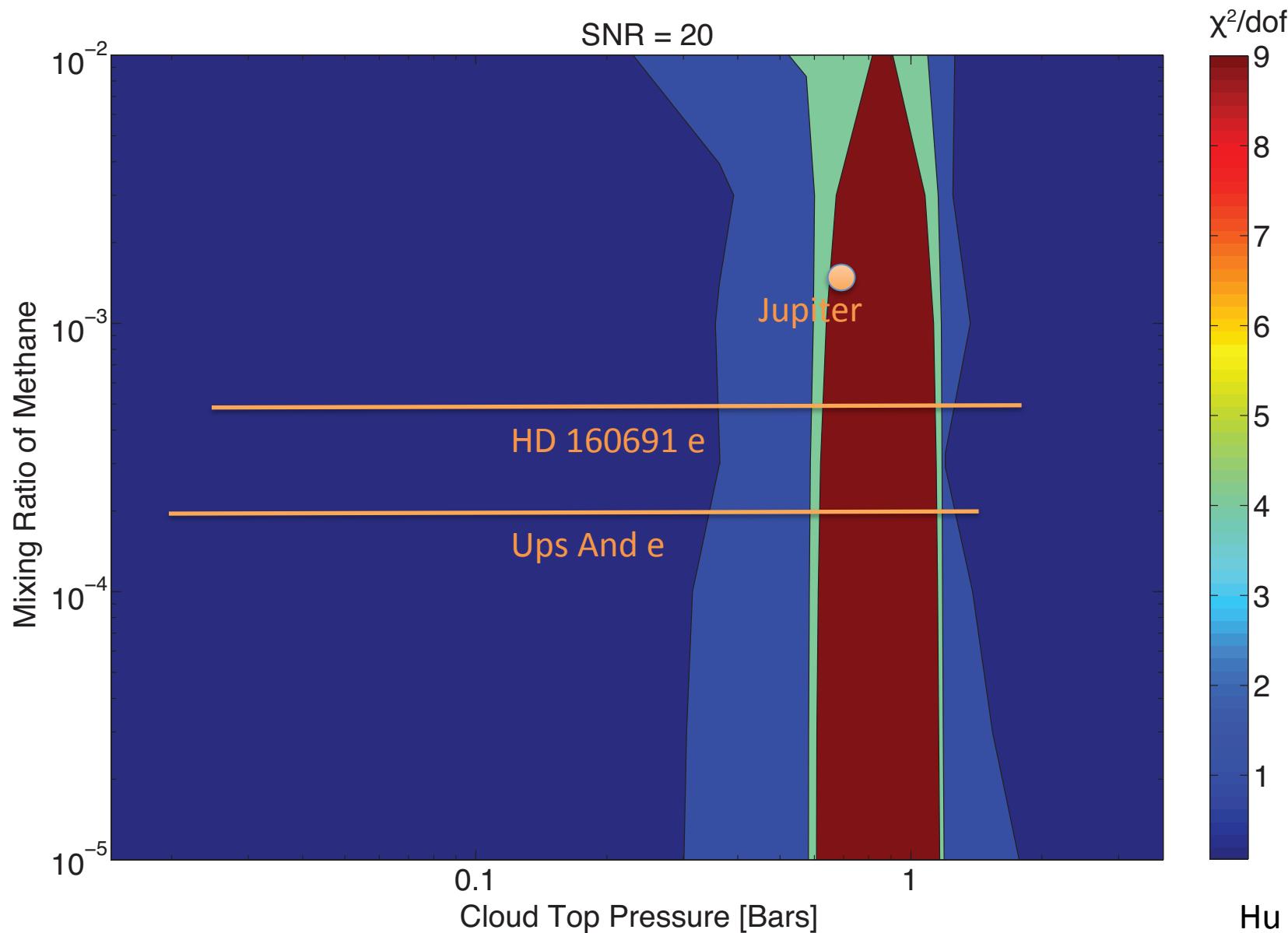


Detectability of Cloud and Gas



χ^2 can be defined as the difference between a model and all adjacent models, for specified spectral resolution and SNR

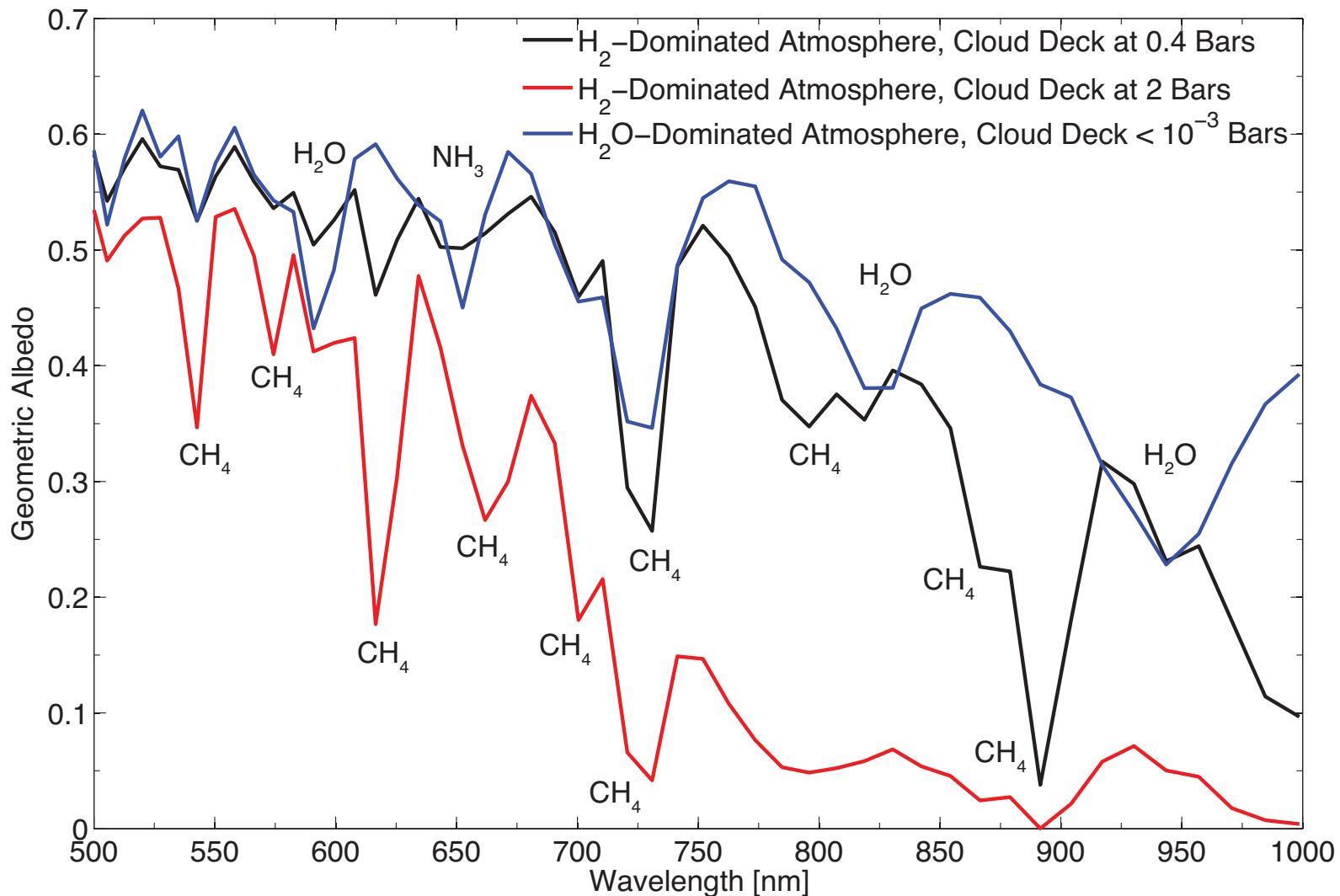
Detectability of Cloud and Gas



Towards Imaging (Super-)Earths



Cool H_2O -rich Super Earths



Conclusion

Exciting scientific discoveries can be anticipated from imaging and characterizing cool, gaseous exoplanets by WFIRST.

Modeling groundwork is ongoing to fully recognize the science potential.

Interplay between clouds and C, N, O abundances

Haze Production in Photochemical Processes

Evolved Neptunes and H₂O-rich Super Earths

Spectral Degeneracies and Information Content

?

Adam Burrows, 2014, **Scientific Return of Coronagraphic Exoplanet Imaging and Spectroscopy Using WFIRST**, arXiv: 1412.6097

Renyu Hu, 2014, **Ammonia, Water Clouds and Methane Abundances of Giant Exoplanets and Opportunities for Super-Earth Exoplanets**, arXiv:1412.7582

Mark Marley et al., 2014, **A Quick Study of the Characterization of Radial Velocity Giant Planets in Reflected Light by Forward and Inverse Modeling**, arXiv: 1412.8440

Thermal Evolution from the Cloud Pressure

