

### WFIRST Microlensing Pipeline

As part of MICROSIT

NASA Goddard

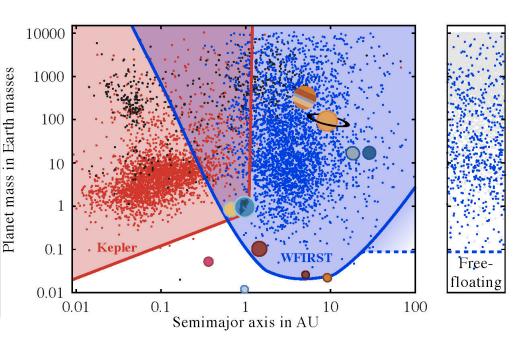
University of Maryland College Park

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#### WFIRST Exoplanet Microlensing Survey



Selected as top priority project in 2010 Decadal Survey Together, Kepler and WFIRST complete the statistical census of planetary systems in the Galaxy.



WFIRST will

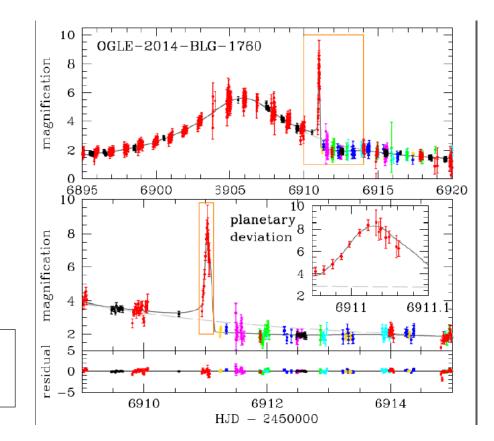
- 1) detect and
- 2) measure mass
- build mass function of planets

2

#### Science Needs

- Light curves to detect the planetary signal
- Mass Measurement using host star detection or parallax measurement
- Detection Efficiency Calculation and yielding Mass Function

Requires: Time series Photometry + Detection and analysis of time varying Modeling, star catalog of proper motions

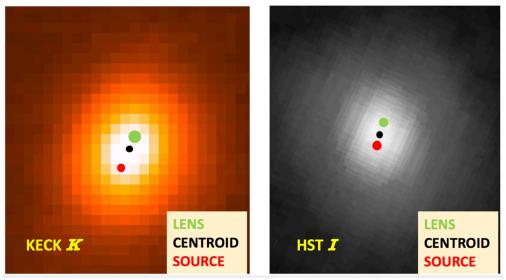


Bhattacharya+ 2016

### Science Needs

- Light curves to detect the planetary signal
- Mass Measurement using host star detection or parallax measurement
- Detection Efficiency Calculation and yielding Mass Function

Requires: High Precision Astrometry specially in crowded fields, Track of neighbors and contaminations from binaries, Uncertainties as a function of positions, brightness and color, aims to measure centroid shift of 3 mas shift at 3 sigma, data handling of >6000 images



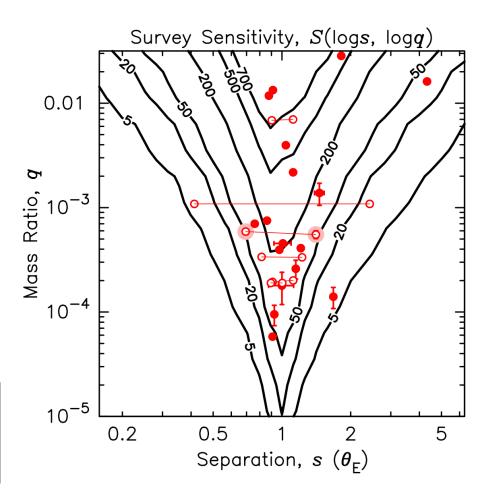
HST -> 7 images, Keck -> 40 images

Bhattacharya+ 2018

#### Science Needs

- Light curves to detect the planetary signal
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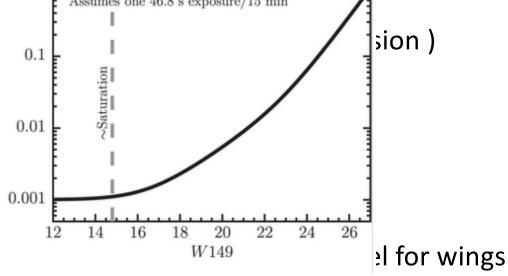
Requires: Injection of planet anomalies in the light curves, Injection of stars with different brightness at image level, reliability of the software



Suzuki+ 2016

# Parts of Microlensing pipeline interesting to other science

- Build Reference image and keen track of star catalog: mag, color, err, positions, proper
- Time series photo
- Difference Image
- Detection of movi
- Crowded field PSF
  Following a hybr
  --Empirical mode



• PSF model varying with space and color

# Parts of Microlensing pipeline interesting to other science

- Build Reference in positions, proper
- Time series photo
- Difference Image
- Detection of movi
- Crowded field PSF
  Following a hybr
  --Empirical mode

g: mag, color, err,

el for wings

Penny+ 2019

## Parts of Microlensing pipeline interesting to other science (continued)

- Microlensing modeling of light curves (including non-planetary)
- Injections of stars of different magnitudes at image level
- Completeness calculation and reliability of software
- Calibration using PRIME telescope IR microlensing survey -science data of 5 yrs before launch
- Photometry and astrometry will be in several batches:
  - Daily photometry
  - End each season photometry
  - End of all season photometry

### Additional Advantage

- ~40000 images over 6 seasons (W149), ~1000 images in other passbands, ~6000 images every season
- Total microlensing field area ~2 deg<sup>2</sup>, 100 millions of stars.
- High precision photometry + astrometry + time series variation
- The pipeline prototype will be built and tested for efficient algorithms and bugs -- Before and during the mission, i.e. till 2031 (season in the beginning and in the end)
- Updated catalog daily, each season and end of season

Huge insight into data – please take advantage of it!! Open to better algorithms or testing algorithm ideas from other SIT Thank you!!